

A_N DY: Measurement of the Analyzing Power for Large Rapidity Drell-Yan Production

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1 Introduction

We propose to measure the transverse single spin asymmetry A_N for Drell-Yan virtual photon production at RHIC. If the current understanding of the space/time structure of color flow in hard scattering processes is correct [1], this asymmetry should have opposite sign to that measured in semi-inclusive deep inelastic scattering (SIDIS) at HERMES [3] and at COMPASS [4]. Color flow in these processes is depicted in Fig. 1. The DY virtual photons will be observed through their decay to e^+e^- pairs which are detected in the apparatus we propose to assemble as shown in Fig. 2. Demonstration that dileptons can be extracted from prolific backgrounds from $p + p$ collisions using event selections described in the original proposal [2], is given by analysis of data from RHIC run 11 in Fig. 3, and is further described below. The total cost for the project is \$1.52M, with \$1.19M in FY12 and \$0.32M in FY13. Details of the theory, simulations, apparatus, and management are presented below.

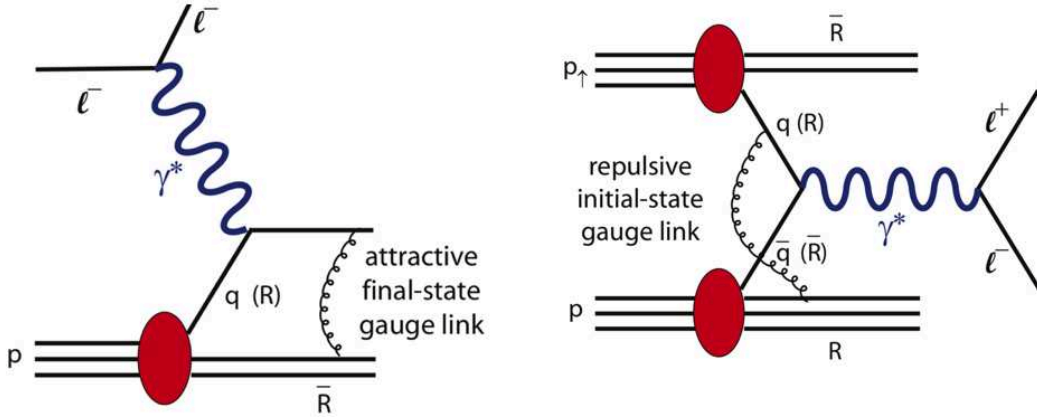


Figure 1: These diagrams show the gauge link in semi-inclusive deep inelastic scattering and Drell-Yan production. Note that color is not annihilated in the SIDIS process, with the interacting quark continuing to feel the attraction of its anti-color partners. In the DY process the forward-going proton sends an anti-color parton to interact with a color-parton from the other proton, with the result that the "spectator" has the same color as the interacting parton from the "target" proton. In both processes, the anisotropy results from orbital motion inside the proton.

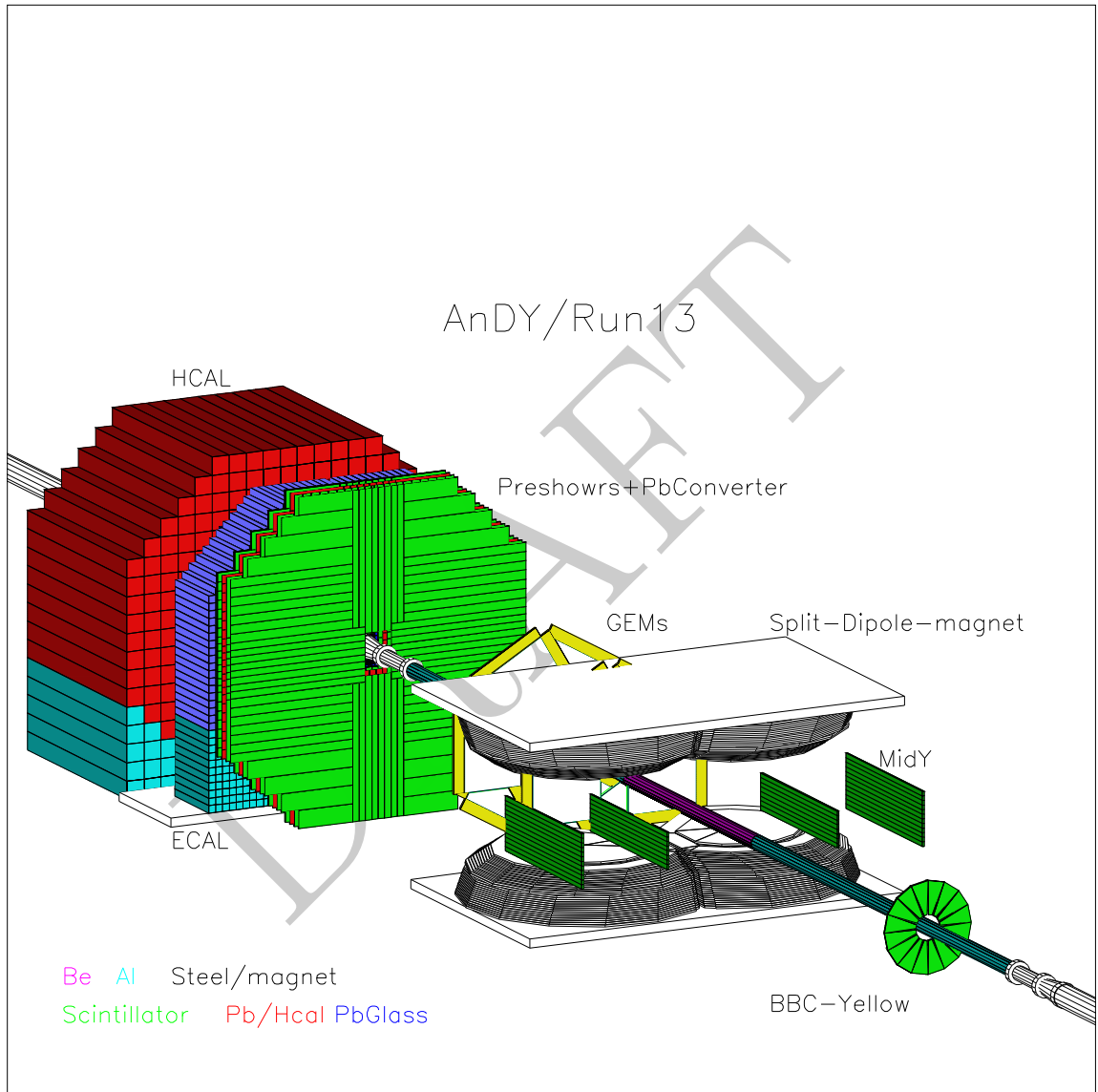


Figure 2: GEANT rendering of the apparatus we propose for observing Drell-Yan production

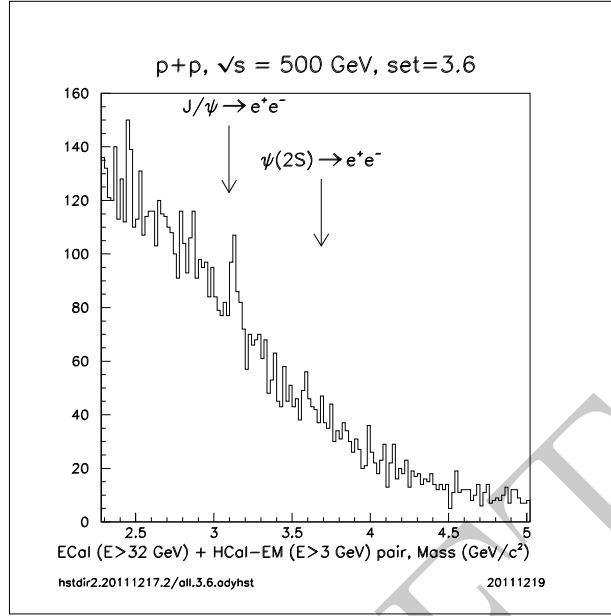


Figure 3: Cluster pair mass distribution including the BBC requirement that clusters have a MIP-like response. Final calibrations are based on individual detector mass distributions. A clear peak from $J/\psi \rightarrow e^+e^-$ is observed, despite the limited statistics provided by the run-11 ECal triggered data set.

¹⁶ A_N DY is an experiment to measure the analyzing power for forward di-
¹⁷ electrons produced by the Drell-Yan (DY) process in polarized proton colli-
¹⁸ sions at $\sqrt{s}=500$ GeV. The experiment is staged at interaction point 2 (IP2)
¹⁹ at the Relativistic Heavy Ion Collider (RHIC) at Brookhaven National Lab-
²⁰ oratory. The physics motivation for A_N DY was described and reviewed by
²¹ the BNL program advisory committee (PAC) in June 2011, and the complete
²² proposal to the PAC is attached as Appendix 1 to this funding proposal. The
²³ committee has endorsed running for A_N DY based on tests we performed in
²⁴ run11 (see ref.html of report), which we summarize here. Measuring DY e^+e^-
²⁵ pairs at RHIC requires multiple simultaneous photon and hadron rejection
²⁶ methods to kill background, the detritus from the spectator, driving the im-
²⁷ plementation presented here. This document proposes the implementation of
²⁸ A_N DY for RHIC runs 13 and run 14, the maximum duration of the project.
²⁹ Transverse single spin asymmetries (SSA) are expected to be vanishingly
³⁰ small in collinear leading-twist perturbative QCD due to the chiral sym-
³¹ metry nature of the theory. However, large spin effects are measured for
³² $p^\uparrow + p \rightarrow \pi^0 + X$ at RHIC energies [5, 6] and at lower \sqrt{s} essentially con-

current with these measurements at RHIC, sizable SSA were also observed for semi-inclusive deep inelastic electron scattering (SIDIS) from transversely polarized proton targets [3]. The large spin effects are not compatible with the collinear leading-twist description of particle production that successfully describes the spin-averaged cross sections and two-particle correlations. Theory can accommodate such large spin effects, however, by going beyond the collinear picture to include transverse momentum dependent (TMD) distribution and fragmentation functions.

Spin-dependent TMD distribution functions describe orbital motion of partons in the proton. A requirement for spin-dependent TMD distribution functions to produce large transverse SSA is that the particle production amplitude has both real and imaginary parts [8]. Without this phase, naive time reversal invariance results in the expectation that spin-dependent TMD distribution functions vanish. The phase arises naturally from gauge invariance and is effectively given by the QCD analog of Coulomb wave functions as the current quark propagates through the color field produced by the spectator partons. For SIDIS, there is an attractive color force between the current quark and the spectators because the virtual photon interacts with an initially color neutral proton. There is a robust theoretical prediction [1] that the attractive color final-state interactions in SIDIS will become a repulsive color initial-state interaction in Drell Yan production, thereby resulting in a change in sign for transverse single spin asymmetries in DY versus those in SIDIS. The goal of A_N DY is to measure the analyzing power for DY production to test this prediction. Achieving this goal would satisfy a DOE performance milestone (HP13).

There are two basic requirements that must be met for A_N DY to succeed: (1) there must be adequate yield for the DY production of virtual photons to provide sufficient statistical precision to test the sign change prediction; and (2) the kinematics for DY production must overlap the kinematics for SIDIS to allow clean comparison. The first requirement is primarily met by luminosity and beam polarization performance for spin-polarized colliding beam operations at RHIC. Detector acceptance is then dictated by the RHIC performance to ensure adequate precision for the measurement. The second requirement sets the range in p_T and x_F where A_N DY must operate. A_N DY will initially concentrate on $M_{\gamma^*} > 4 \text{ GeV}/c^2$, where M_{γ^*} is the DY virtual photon invariant mass, rather than the much lower masses needed to match the Q^2 scale of the SIDIS results, because predicted backgrounds at lower mass appear to preclude measurement of a spin observable. We will,

however, push our analysis below the 4 GeV limit to reconstruct the J/Psi and other charm states, as shown in Fig.3, and to determine the limits of our background rejection.

The importance of the second requirement has been clarified by recent theoretical work [9] that seeks a global analysis of measured analyzing power for pion production at RHIC and transverse SSA for SIDIS. A global analysis of the experimental results is particularly challenging to theory since pion production has no proven factorization theorems that use TMD distribution functions extracted from SIDIS, unlike the case for DY production. Instead, the pion production results are analyzed via a twist-3 formalism for which factorization has been proven [10]. The analysis extracts quark-gluon correlators that can be intercompared to moments of the Sivers function extracted from SIDIS. The RHIC pion production results show large analyzing powers for $x_F > 0.3$ [6]. The SIDIS results span $0.05 < x_B < 0.3$ [3]. Consequently, the data have essentially no kinematic overlap. Accounting for the different gauge links that enter, the conclusion from theory [9] is that there is a sign mismatch between the SIDIS results and the RHIC pion production results. This mismatch can be accommodated if the Sivers function has a node near $x_F \simeq 0.3$ [11, 12]. A_N DY must overlap the kinematics of SIDIS experiments to avoid the complications near this node and to provide a robust test of the theoretical prediction for a sign change. This overlap sets requirements on A_N DY acceptance and on signal:background ratio.

Simultaneous with our measurement of A_N for DY, we propose to measure the neutral pion analyzing power at x_F values that overlap our proposed DY measurements and also overlap our prior measurements at large x_F [6]. This will eliminate any ambiguity in the sign of the spin asymmetries measured.

The primary detector in A_N DY is a high-resolution electromagnetic calorimeter (ECal) built from 1596 lead-glass cells that view particles produced in the forward direction by colliding beams. The discrimination of electrons and positrons from the background of photons and hadrons is accomplished by associating the ECal response with the response of a preshower detector (PS) immediately in front of ECal and with the response of a hadron calorimeter (HCal) immediately behind ECal. A tracking system (TRK) consists of three stations of Gas electron multipliers (GEM) to provide high resolution tracking. These make the association between PS detectors and ECal clusters robust, and help discriminate photons from charged particles. We propose one data set be taken with an existing split-dipole magnet. Tracking through the magnet identifies charge sign and further discriminates showers in ECal

109 produced by electrons and positrons from showers produced by photons pro-
110 duced in the collision.

111 The preshower detector consists of two overlapping planes of 76 scin-
112 tillation counters each followed by a lead converter and a third plane of
113 scintillators. The preshower plays a crucial role in identifying the particle
114 responsible for creating a cluster in ECal by providing a crude measurement
115 of the longitudinal energy deposition profile. Photons will give no response
116 in the first two scintillator planes and a partial shower response in the third
117 plane. Charged hadrons will most probably give the response of a minimum-
118 ionizing particle (MIP) in all scintillator planes. Electrons and positrons
119 will give a MIP-like response in the first two scintillator planes and a partial
120 shower response in the scintillator that follows the converter.

121 The hadronic calorimeter consists of 296 lead-scintillating-fiber cells origi-
122 nally developed for AGS experiment E864. The primary purpose of the
123 HCal is also for longitudinal energy profiling. Hadrons that partially shower
124 in ECal will be rejected by matching the clustered ECal response to the
125 clustered HCal response.

126 The tracking system consists of 24 GEM modules, each 40cm x 50 cm
127 incorporating 2250 strips, with ADC values for each strip allowing interpola-
128 tion to provide $\leq 100\mu$ position resolution. Stations 1 and 2 (TRK1 and
129 TRK2) each consist of 6 modules in overlapping hexagons, providing effi-
130 ciency $\geq 98\%$ for charged particles. Station 3 (TRK3) consists of a single
131 plane of 12 modules. The electrons and hadrons produce single MIP re-
132 sponse, while the e^+e^- pair from a photon conversion in the beam pipe can
133 be distinguished by its 2-MIP signal in each module of TRK1.

134 This document is divided into the following sections. Section 2 describes
135 the run-11 tests for A_NDY with a focus on the primary goals of establishing
136 the impact of collisions at IP2 on STAR and PHENIX and of the operation
137 and calibration of the HCal. Section 3 and section 4 describe the proposed
138 configuration of ECal and HCal. Section 5 describes the proposed configu-
139 ration of the preshower. Section 6 describes a detector to measure vertex
140 location. Section 7 briefly describes the proposed forward pion detector im-
141 plementation that closes as many theoretical loopholes as possible. Section
142 8 describes A_NDY tracking detectors. Section 9 describes the split-dipole
143 magnet proposed. Section 10 describes the triggers we will use and the data
144 acquisition system. Section 11 describes the simulations we use to design
145 and understand our experiment. Section 12 describes our computer needs.
146 Section 13 is a summary of budget and our resource loaded timeline. Section

147 14 is our management plan and Section 15 addresses safety.

148 **2 Run-11 Summary**

149 A preliminary test for A_N DY was implemented at IP2 for RHIC run 11
150 (Fig. 4). The primary goals of RHIC run 11 as described in the 2010 Letter
151 of Intent to the PAC [2] were to

- 152 • establish the impact of initiating collisions at IP2 on operations at IP6
153 and IP8; and
- 154 • to demonstrate that the hadron calorimeter modules (HCal) could be
155 calibrated.

156 This section summarizes the work done prior to RHIC run 11 in which we
157 tested portions of the major detector components we propose to assemble for
158 A_N DY, and summarizes the status to date of understanding collision data
159 from the run-11 A_N DY configuration.

160 **2.1 Preparations for RHIC run 11**

161 The IP2 area is the former location of the magnetic spectrometers built and
162 operated by the BRAHMS collaboration [17]. In August, 2010 the rem-
163 nants of the BRAHMS experiment were still in place at IP2. Led by Charles
164 Folz, different teams from the Collider-Accelerator (C-A) department made
165 a major cleanup of the IP2 area in a 3-month period, effectively removing
166 all remnants of the apparatus built by BRAHMS. In addition, these C-A
167 teams installed AC power distribution and cable tray for the run-11 A_N DY
168 electronics; designed and implemented supports for reconfigured HCal mod-
169 ules formerly at IP10 and used by the PHOBOS collaboration; designed and
170 implemented support stands for small ECal modules populated by 120 lead-
171 glass detectors borrowed from BigCal [14] at Jefferson laboratory; designed
172 and implemented support stands for two beam-beam counter arrays borrowed
173 from the University of Maryland [15]; aided in the construction of preshower
174 detector supports; installed a beryllium beam pipe with specially designed
175 support stands; surveyed the apparatus that got built at IP2 for the run-11
176 A_N DY configuration; and installed fiber-optic communication lines for clock
177 distribution from RHIC-standard V-124 modules to the A_N DY electronics,

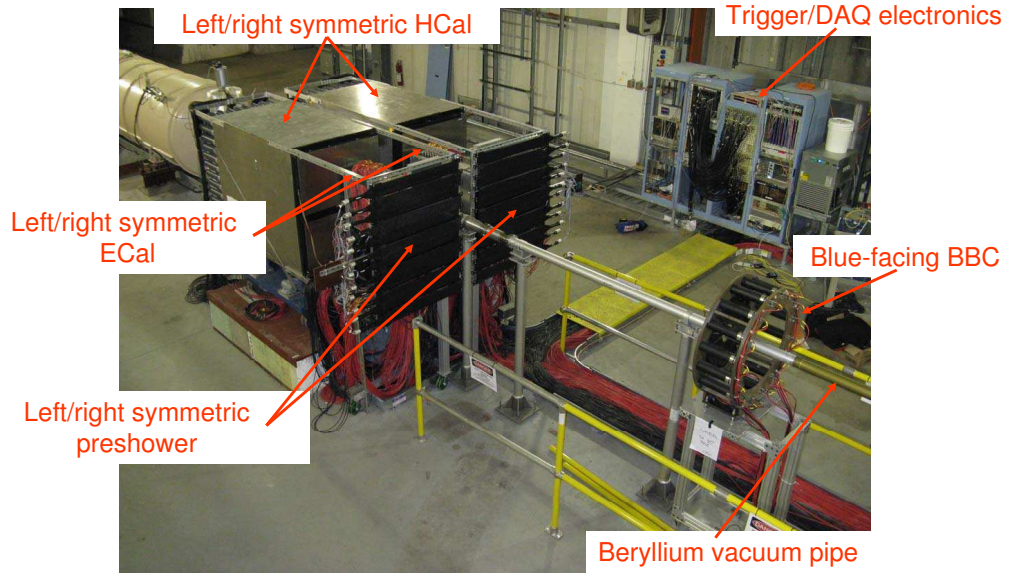


Figure 4: Picture of the A_NDY apparatus for RHIC run 11. Not shown in the picture are the Yellow-beam facing beam-beam counter array and the zero-degree calorimeters. The hadron calorimeter (HCal) modules were built from detectors constructed for AGS-E864 [13]. The electromagnetic calorimeters (ECal) were built from lead glass borrowed from Jefferson Laboratory [14]. The beam-beam counters (BBC) were built for PHOBOS [15]. The preshower scintillators were built for the run-11 A_NDY implementation.

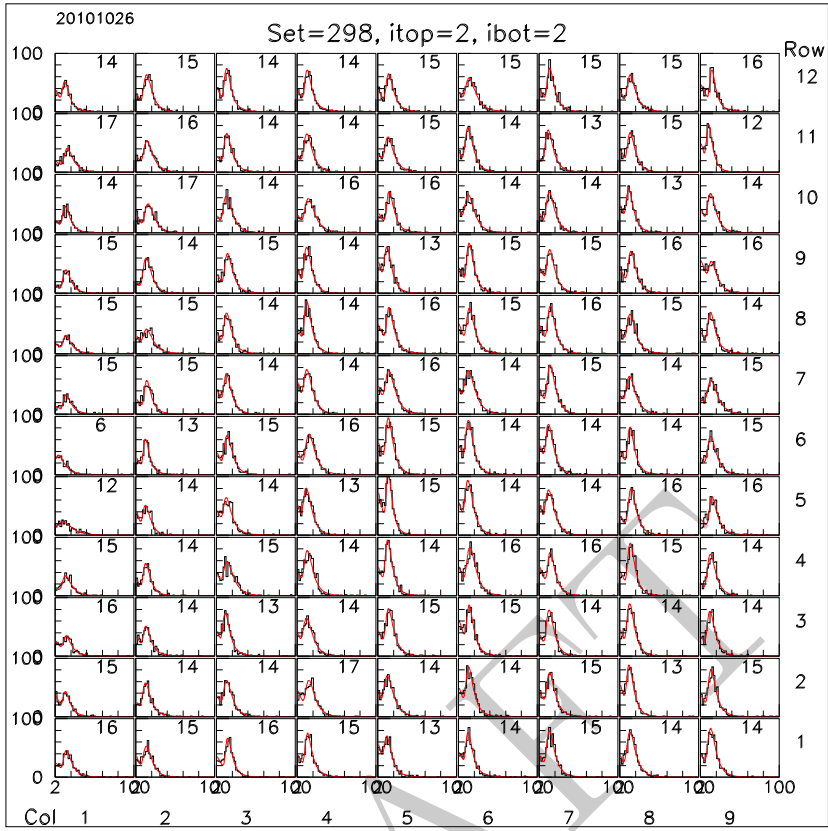


Figure 5: Individual detector ADC spectra from the beam-left HCal module for cosmic-ray triggers taken prior to run 11. The data show clear peaks associated with a minimum-ionizing particle (cosmic-ray muon) passing through the stack. The distributions are fit by an exponential background summed with a Landau distribution. The centroid of the Landau distribution for each detector is shown on the plot.

178 a fiber-optic network used for data acquisition and to provide monitoring
 179 as feedback to the RHIC main control room. The resulting run-11 A_N DY
 180 apparatus is shown pictorially in Fig. 4.

181 Members of A_N DY were busy from August, 2010 through January, 2011
 182 building and testing detector apparatus and the trigger/data acquisition elec-
 183 tronics [18]. The HCal modules that had been used at PHOBOS were moved
 184 to IP2 in August and the cosmic-ray muon calibration of these detectors
 185 began immediately. A key that enabled this effort was that the STAR trig-
 186 ger system [18] had all the ingredients of a configurable trigger and data
 187 acquisition system. That step had been taken in 2009 for cosmic-ray muon
 188 based calibration of a 5-row \times 10-column hadron calorimeter module orig-

inally built for PHOBOS from the AGS-E864 detectors, as reported by a graduate student at a Division of Nuclear Physics meeting [19]. That system was transported to IP2 in August, 2010 and was initially made operational for 64-channel readout and later extended for testing and ultimately for the run-11 A_NDY trigger/DAQ system (see 2.2).

We also had one 32-channel low-voltage distribution board to power the Cockcroft-Walton bases for the first tests of the HCal, of the type fabricated to power all CW bases used for the run-11 A_NDY configuration. The same methods used for relative gain monitoring and testing of lead-glass calorimeters worked very well for HCal (see Section 3.5 for more detail of the fiber-optic distribution of light-emitting diode (LED) flashes). Consequently, a gain curve for each HCal photomultiplier tube was measured. This data proved essential during the RHIC run when first indications of the HCal energy scale for showers from particles produced in collisions became available, and facilitated a gain change of HCal during the run.

The portable trigger/DAQ system was expanded to 128 channels of readout and the rapid fabrication of the low-voltage distribution boards resulted in full 9-column \times 12-row HCal module tests with cosmic-ray muons, with concurrent LED-based gain monitoring, starting in early October. The trigger used for this work ensured complete integration of the cosmic-ray muon signals through the clocked trigger/DAQ system by an effective time-to-amplitude converter on plastic scintillator paddles put above (top) and below (bottom) the HCal stack. Three paddles were put on top of the stack and three below the HCal stack to localize the cosmic-ray muon trajectory along the length of the HCal detector, to confirm that the attenuation length of the scintillating fibers in the detectors was observed. The “beam-left” HCal module calibration was completed in late October with final data as shown in Fig. 5. Similar cosmic-ray calibrations were completed for the “beam-right” HCal module just before the modules were moved onto their support stands for operation in run 11. The portable DAQ system was subsequently dismantled and was reconfigured for the run-11 A_NDY trigger/DAQ system.

A loan agreement between the Yerevan Physics Institute (YerPI) and BNL was signed by BNL, Jefferson Laboratory and YerPI in October, 2010. This agreement enabled the loan of 120 40mm \times 40mm \times 40cm lead-glass bars, FEU-84 photomultiplier tubes and resistive bases from a calorimeter that had been constructed, commissioned and used for experiments at JLab (BigCal [14], further discussed in Section 3. The primary purpose of this loan was to assess whether the lead glass was sufficiently annealed of color-trapping

227 centers after use of BigCal in a high-luminosity experiment at JLab. After
 228 transport to BNL, transparency measurements and visual inspection showed
 229 the glass to be clear, except for ~ 8 bars. These bars were annealed at BNL by
 230 ultraviolet lamps. A secondary application of these borrowed lead-glass de-
 231 tectors was to make a pair of small-scale electromagnetic calorimeter (ECal)
 232 modules by adapting enclosures that were in storage at BNL. The adaptation
 233 required an interface between the SHV feedthroughs on the exterior paneling
 234 of the enclosures and Hypertronix connectors on the resistive bases built by
 235 YerPI and used in BigCal. Optical coupling of FEU-84 to the lead glass was
 236 accomplished by using threaded rods screwed into vanadium blocks glued
 237 to the glass. The FEU-84 was pushed against the lead glass using a thin
 238 pusher plate stamped with the pin pattern of the phototube. Optical grease
 239 between the phototube entrance window and the lead glass was used for the
 240 coupling. This is the same method we propose for these detectors for the
 241 full-scale A_N DY ECal, as described in Section 30. Neutral pion reconstruc-
 242 tion from the run-11 A_N DY ECal modules showed expected mass resolution,
 243 as discussed below.

244 Finally, we built a prototype of the preshower counters envisioned for the
 245 final A_N DY configuration. BC-408 scintillator was ordered from St. Gobain
 246 and delivered by early December, 2010. The scintillator was carefully in-
 247 spected and then wrapped in aluminized mylar reflectors. Forty Hamamatsu
 248 H5010 photomultiplier + resistive base assemblies were borrowed from AGS-
 249 E896 and were used for the 2.5-cm wide \times 90-cm long \times 1.0-cm thick long
 250 scintillator strips. A total of 24 Photonis XP2972 photomultipliers with
 251 Cockcroft-Walton bases were borrowed from AGS-E864 [13] and used for
 252 10.5-cm wide \times 90-cm long \times 1.0-cm thick scintillator strips. The wide
 253 scintillators were individually wrapped in black tedlar to make the detectors
 254 light tight. The construction of the preshower scintillator plane built from
 255 2.5-cm wide strips is described in some detail below, since we propose a sim-
 256 ilar construction method for the A_N DY preshower proposed for RHIC run
 257 13.

258 The construction method for the preshower plane built from 2.5-cm wide
 259 scintillator strips was as follows. Each strip was individually wrapped with
 260 aluminized mylar. After wrapping, the strips were glued to a roughened
 261 aluminized mylar substrate on each side to form a uniform and rigid 50-cm
 262 wide \times 90-cm long plane. This gluing method minimizes cracks between
 263 adjacent scintillator strips. The method had been developed for scintillator-
 264 strip shower maximum detectors used in electromagnetic calorimeters. Fiber-

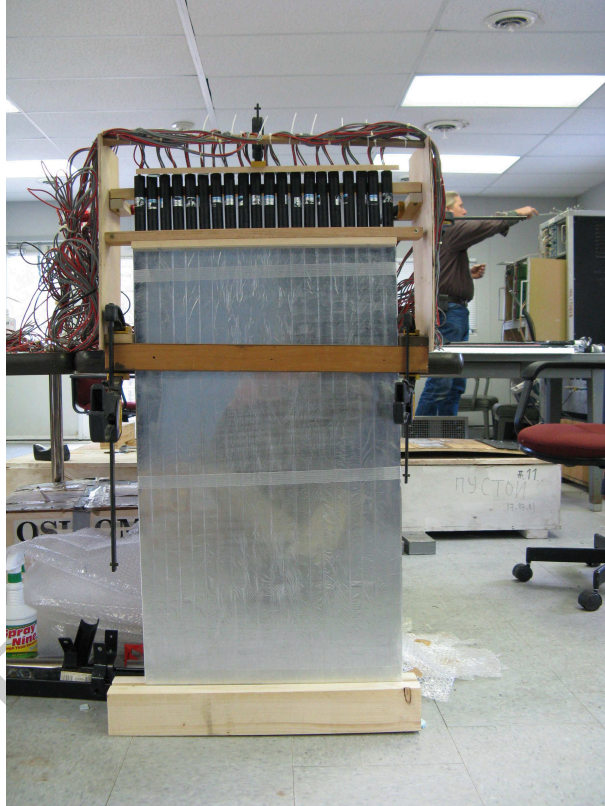


Figure 6: Stand used for gluing Hamamatsu H5010 phototube assemblies to 2.5-cm wide \times 90-cm long \times 1.0-cm thick BC-408 strips. There are 20 such strips glued to make a rigid plane.

glass reinforced tape was wrapped around the glued assembly to reinforce the glue joints. The planned construction for the preshower will eliminate a tape wrap in the center of the resulting plane.

The array of H5010 phototube assemblies were then glued at one end of the scintillator strips. The difficult part of this assembly was to compensate for the torque on the H5010 from the integral cabling for high-voltage to the resistive divider and co-axial signal cables for the anode output. A jig was constructed for the gluing of the phototube assemblies to the strips. The jig enforced the centering of the phototubes on each strip and enabled downward pressure on the assemblies while the glue was setting. A photograph of one plane in its gluing jig is in Fig. 6. After this step, the entire plane was enclosed in black tedlar to make the detector plane light tight. This step will be avoided for the run-13 preshower construction due to the potential for non-uniformities of material at the preshower. Instead, the unwrapped plane will be mounted within the ECal enclosure, as discussed below. The entire process of preparing the preshower scintillators took four weeks, although the actual time involved was mostly associated with preparing and implementing a gluing, followed by ~ 12 hours to allow a complete hardening of the glue.

A final essential component to the run-11 preparations was a GEANT model of the apparatus (Fig. 7). This had initially been prepared to study A_N DY configurations prior to the 2010 letter of intent [2]. That initial model was then extended to include the small ECal modules, the beam-beam counters and the preshower arrays ultimately built and implemented prior to RHIC run 11.

2.2 Description of run-11 trigger/DAQ system

The A_N DY trigger and data acquisition system is based on hardware fabricated for the first generation of the STAR experiment and is described in more detail in Sec.10. A_N DY uses the RHIC standard Zero Degree Calorimeters (ZDC) located at the exit of each final bending magnet (DX magnet) in the beam lines $\sim 8m$ on either side of the IP in setup and triggering. Each ZDC consists of 3 PMTs viewing fibers from a sphagetti calorimeter and a Shower Max Detector (ZDC-SMD) consisting of 8 vertical and 8 horizontal scintillators. The ZDC-SMD is used in local polarimetry, and discriminator outputs from each strip are recorded in a 10 MHz histogramming scaler board for each RHIC crossing. Minimum bias triggers used in run 11 were based on the ZDC and BBC detectors, with high-tower, cluster, and pair triggers

IP2/DY-Run11

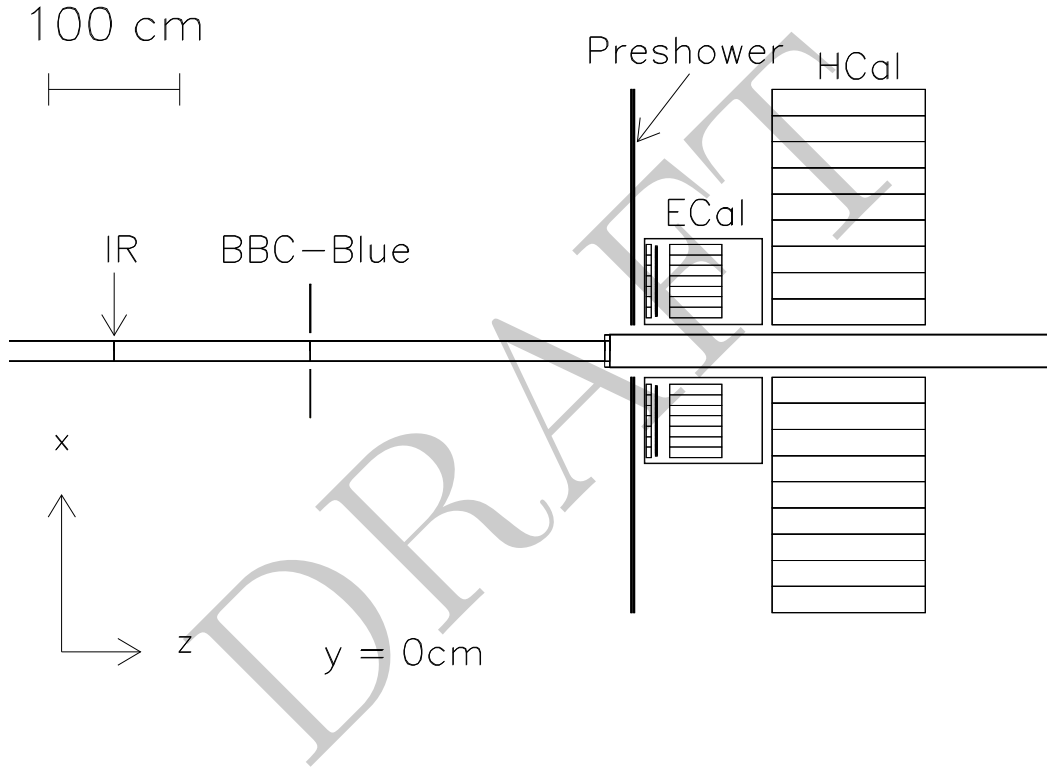


Figure 7: GEANT model of run-11 A_N DY configuration. The Blue beam travels in the positive z direction and the Yellow beam travels in the negative z direction, with the collision diamond centered at $z=0$, as indicated by IR on the drawing. Simulation samples consist of events generated by PYTHIA 6.222 that are then tracked for each particle through the GEANT model.

Trigger Electronics Tree – Run 2011

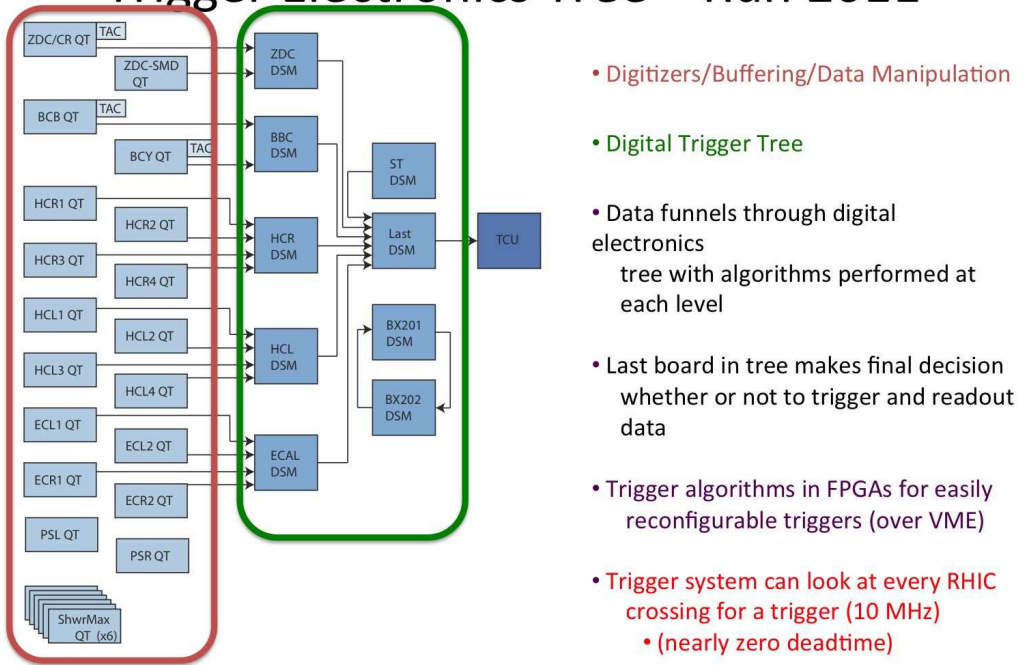


Figure 8: DSM trigger logic tree for run 11.

301 adding information from ECal and HCal as shown in the logic tree of Fig.8.

302 2.3 Impact of IP2 collisions and projections for Lint 303 for runs 12,13

304 Testing the impact of IP2 collisions on operations at IP6 and IP8 was a pri-
305 mary objective of RHIC run 11 [2]. The issue that needed to be addressed
306 is the impact of tune shifts and tune spreads of one beam induced by the
307 other beam (so-called beam-beam tune shifts). In previous years for $\sqrt{s}=200$
308 GeV polarized proton collisions, bunches from the Blue ring that interacted
309 at more points around the ring with bunches from the Yellow ring showed
310 increased rate of ion loss and emittance growth through the store, leading to
311 lower luminosity. RHIC requires bunch intensities approaching 2×10^{11} pro-
312 tons/bunch to achieve the luminosities required for the RHIC spin program.
313 The betatron oscillation tunes are limited by all the orbit instability resonant

314 conditions. Preserving polarization adds different limitations on the tunes.
315 This additional complexity is not present in other colliders making RHIC
316 a unique challenge in the world for accelerator physicists. The effects of
317 beam-beam tune shifts increase linearly with the number of ions per bunch,
318 are inversely proportional to the emittance of the bunch and depend on the
319 number of bunch crossing points in the collider. These tune shift effects,
320 however, do not depend on the ion optics, *i.e.* the β^* used at collision points.

321 The limitations from beam-beam tunes shifts and spreads were fully recog-
322 nized before RHIC run 11. Documentation of the impact of a third collision
323 point was readily available from prior years when beams collided at three
324 points around RHIC. This documentation was a topic of multiple discus-
325 sions between A_NDY proponents and accelerator physicists in the Collider-
326 Accelerator department. These earlier experiences were not sufficient to be
327 certain of the impact of collisions at IP2 on operations at IP6 and IP8 since
328 many different working points in the betatron oscillation tune space have
329 been used to satisfy the additional constraint of preserving polarization while
330 also avoiding orbit instability resonances. The question posed by the 2010
331 program advisory committee was to quantitatively establish the impact of
332 IP2 collisions. The expectation from the prior RHIC runs was that there
333 would be some loss of luminosity at IP6 and IP8. The fundamental ques-
334 tion was the magnitude of this loss, since small losses (*e.g.*, <20%) could
335 accommodate running A_NDY in parallel with the W program at STAR and
336 PHENIX while large losses (*e.g.*, >20%) would be best accomodated by do-
337 ing the experiments serially rather than in parallel. The plan for polarized
338 proton operations at $\sqrt{s} = 500$ GeV in run 11 was to operate RHIC with
339 fractional betatron oscillation tunes near the 2/3 orbit instability resonance
340 so as to improve the polarization of the beams over what had been achieved
341 in RHIC run 9.

342 A further complicating factor for IP2 operation was that the DX magnets
343 at IP2 had in the past exhibited less robust quench training behavior than
344 nominally identical magnets at the five other interaction regions in RHIC.
345 This prompted the plan to start polarized proton operations at $\sqrt{s}=500$
346 GeV in run 11 with a ~ 2 mr bunch crossing angle at IP2. With typical
347 bunch lengths from prior years of RHIC operation, the non-zero crossing
348 angle corresponded to a factor of three less luminosity relative to head-on
349 collisions. Through the run, training quenches of the DX magnets either side
350 of IP2 were done during times when issues precluded RHIC operation. There
351 was a brief period of ~ 1.6 mr bunch-crossing angle operation at IP2 and then

head-on collisions at IP2 after the last training quench on 18 March. The scripted start of IP2 collisions did require changes when the crossing angle changed, since the required currents in corrector magnets to initiate collisions in one step depended on the crossing angle.

RHIC stores have the sequence of bunch injection, acceleration ramp, spin-rotator ramps to precess the spins to make longitudinal polarization for collisions at IP6 and IP8, cogging of the two beams thereby having bunches from each ring cross the interaction regions at the same time and then removal of transverse separations of the two beams at IP6 and IP8 to initiate collisions there. Early in each RHIC store, 10-mm transverse separation bumps of the two beams remained in place at IP2. Late in the store, separation bumps at IP2 were removed to initiate collisions. The beam-beam tune shifts and spreads initiated by the IP2 bump removal would lead to increased ion loss and potentially impact the luminosity at IP6, IP8. The time dependence of the beam intensity, the loss rate of ions in each ring, and luminosities at IP6 and IP8 were available in real time through the run through an interface to RHIC data in the A_N DY counting house, and at STAR and PHENIX. The real time strip charts were recorded for most bump removals in the A_N DY electronic log. An example of an IP2 bump removal record from the A_N DY e-log entry 1845 for RHIC fill 15436 is shown in Fig. 9

Some fills had quantitatively different behaviors than shown in Fig. 9, although the qualitative features of these plots were present in most fills. The qualitatively common features are:

- there was an initial increase in ion loss rate when collisions were initiated at IP2 that subsequently decayed with increasing time.
- The ion loss rate was almost always bigger in the Yellow ring than in the Blue ring, independent of whether there were more ions in Yellow or more ions in Blue at the time IP2 collisions began.
- Some fills had larger peak loss rates than others. This appeared to be associated with the betatron oscillation tunes for the store and whether there were tune adjustments before or after the IP2 bump removal. Typically worse loss rates, or longer times for the loss rate to damp away, were associated with tune adjustments after the bump removal.

The plan to quantitatively assess the impact of IP2 collisions on IP6 and IP8 was to systematically increase the average bunch intensity when

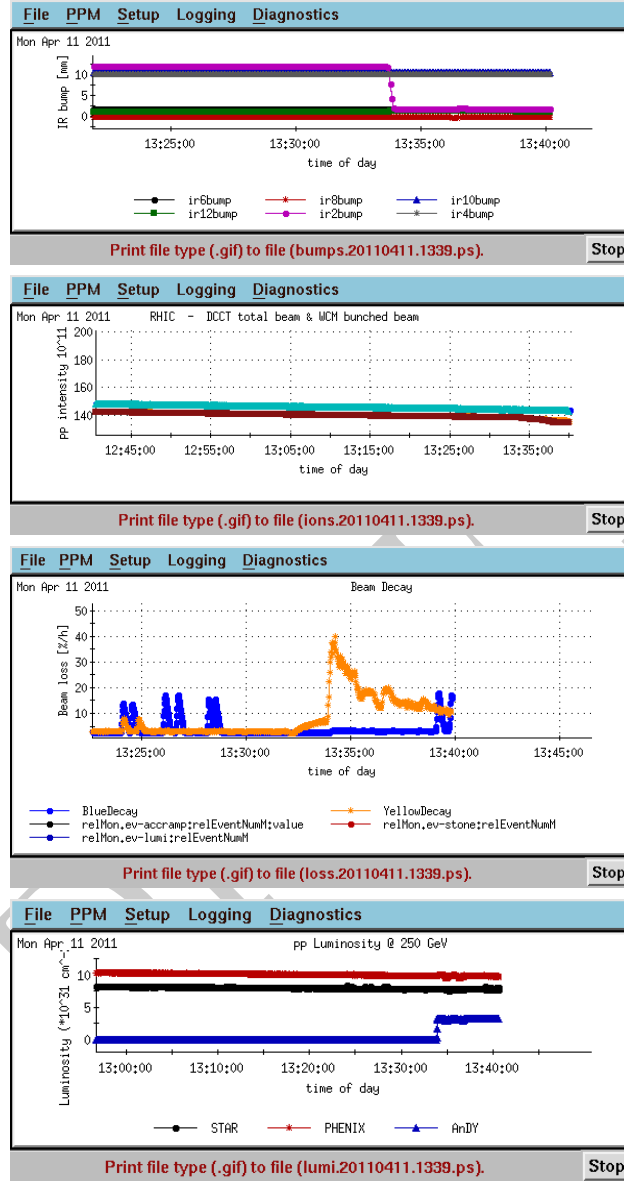


Figure 9: Time dependence strip charts available as real-time feedback that document the start of collisions at IP2 for RHIC store 15436. The four plots from top to bottom are: (1) transverse separation bumps at the six interaction regions around RHIC. At 13:34, corresponding to 197 minutes after the declaration of the start of physics for this store, the separation bump at IP2 is removed; (2) number of ions in the Blue and Yellow ring. A small change in slope is observed for the Yellow ring at 13:34 when IP2 collisions are initiated; (3) the ion loss rate (D) versus time for the Blue and Yellow rings. The peaks are associated with insertions of the carbon ribbon used for Coulomb-nuclear interference (CNI) polarimeters. There is an increase in D in Yellow when collisions are initiated at IP2 that damps away with increasing time; and (4) the luminosity at IP6, IP8 and IP2 as measured by the rate of coincidences observed in zero degree calorimeters common to all interaction regions. The change in instantaneous luminosity at IP6 and IP8 is not evident for this store.

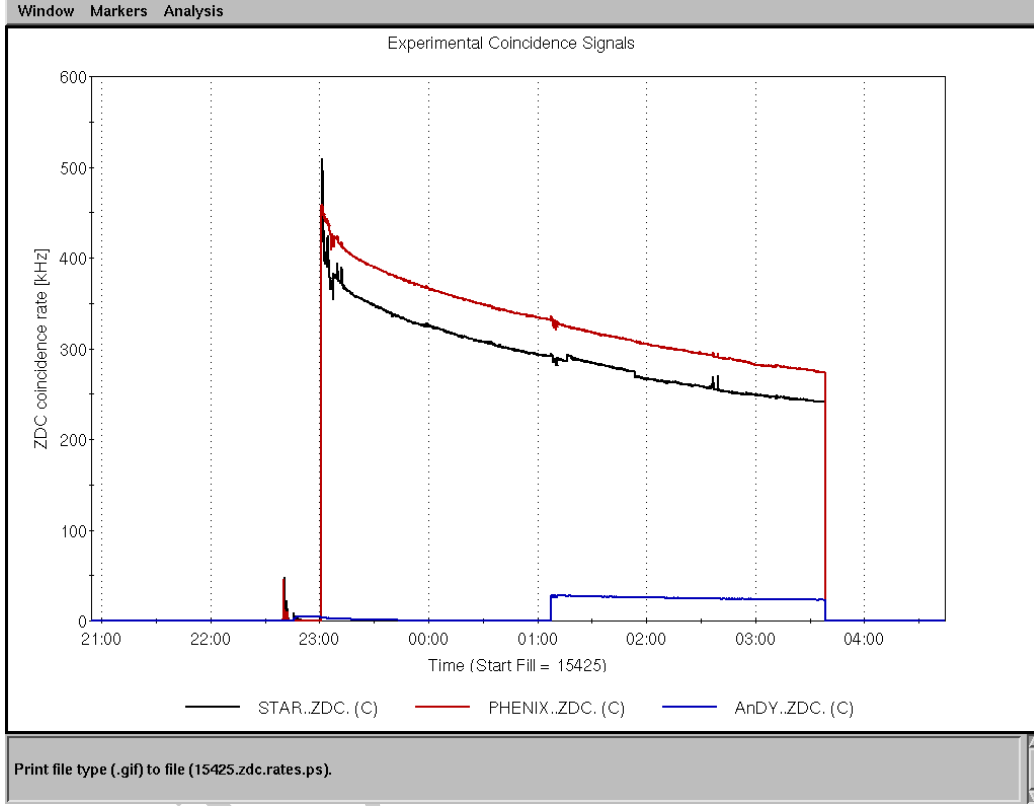


Figure 10: ZDC coincidence rate at STAR, PHENIX and IP2 versus time for RHIC store 15425. IP2 collisions were begun when the bunch intensity had decreased to 1.50×10^{11} ions per bunch. Narrow peaks in the STAR ZDC coincidence rate are associated with backgrounds accompanying CNI polarimeter measurements. The maximum rate of ZDC coincidences at the different interaction points reflects differences in ion optics at the three collision points ($\beta^*=3\text{m}$ at IP2. Smaller values are used for IP6 and IP8, resulting in larger luminosity.) and differences in ZDC operating voltages and thresholds. Both effects are corrected for by vernier scans.

the IP2 transverse separation bumps were removed. The starting point was 0.95×10^{11} ions per bunch. The increment in this threshold was 0.05×10^{11} ions per bunch. This increment sometimes was automatically implemented for successive stores and sometimes was not implemented, depending on the real-time feedback. While this study was underway, there was a separate program to increase the number of ions injected into RHIC to increase the luminosity for $\sqrt{s} = 500$ GeV polarized proton collisions. The program to increase the bunch intensity meant that the time when IP2 collisions began got successively later in store, or meant that the store was dumped before IP2 collisions began. When the luminosity increase program reached its plateau, the time when IP2 collisions began became progressively earlier. The last point studied was the start of IP2 collisions at the average of 1.50×10^{11} ions per bunch. The luminosity at STAR, PHENIX and IP2 for RHIC fill 15425 is shown in Fig. 10. Collisions began at IP2 when the intensity had fallen below the 1.50×10^{11} ions per bunch threshold.

The impact of IP2 collisions on operation at IP6 and IP8 was continually watched and discussed with accelerator physicists and members of STAR and PHENIX during the polarized proton run. Discussions were possible by participation of A_NDY at each daily RHIC operation meeting focused on status and planning; at each weekly Time meeting, used to report A_NDY progress and plans and to hear reports from the accelerator subsystems and the other experiments; at each weekly accelerator/experiment planning meeting; and at special meetings, as needed. Discussions and reports are documented on public web sites. These discussions led to refinements in the general procedure for starting collisions at IP2. For example, one important refinement was to synchronize the transverse separation bump removal at IP2 with scheduled CNI polarimeter measurements, as requested by the PHENIX collaboration since they reduced operating voltages on their muon arm detectors for both activities due to the potential for increased background.

A summary of the integrated luminosity recorded at IP2 is shown in Fig. 11. The vertical axis on this plot is proportional to the number of coincidences from the IP2 beam-beam counters that face the Yellow and Blue beams. Also shown is the integrated luminosity deduced from the number of coincidences from the IP2 ZDC that face the Yellow and Blue beams. The proportionality constant represents an effective cross section. For this plot, the effective cross section for the IP2 BBC is from analysis of a complete PYTHIA/GEANT simulation, discussed in more detail below. The simulation generates non-elastic, non-single diffractive p+p collision events at \sqrt{s}

425 = 500 GeV using PYTHIA 6.222 with default settings. The PYTHIA events
 426 are then run through the IP2 GEANT simulation. Energy deposition in the
 427 BBC detectors modeled in this simulation is then digitized to simulate the
 428 ADC response, accounting for photostatistics and photomultiplier resolution.
 429 An online analysis of the vernier scan at IP2 for RHIC store 15457 resulted in
 430 the effective cross section of 0.98 mb for ZDC coincidences. This is within 7%
 431 of the effective cross section for the ZDC deduced from the simulations. The
 432 systematic error estimates associated with the integrated luminosity mea-
 433 surement still must be completed. As discussed below, it is expected that
 434 the $\sim 6.5 \text{ pb}^{-1}$ of integrated luminosity delivered to IP2 for polarized pro-
 435 ton collisions will result in the first measurement of the analyzing power for
 436 forward jet production. Unbiased jet triggers and reconstruction is possible
 437 with this data obtained with the HCal modules in place at IP2.

438 With a primary motivation of calibrating the IP2 ZDCs using the prolific
 439 100-GeV neutron flux, there were two stores where full-energy Au ions col-
 440 lided at IP2. This is further discussed in section 3.4. Beam-beam tune shifts
 441 and spreads are not a limiting factor for heavy ion operations at RHIC.

442 A conclusion that had been reached and agreed upon by the experiments
 443 and the accelerator physicists was that IP2 collisions could begin at average
 444 bunch intensities corresponding to ~ 3 hours after physics was declared for the
 445 store, with minimal impact on operations at IP6 and IP8. This conclusion
 446 resulted in accelerator physicists estimating that $10 \text{ pb}^{-1}/\text{week}$ integrated
 447 luminosity could be delivered to IP2 during $\sqrt{s} = 500 \text{ GeV}$ polarized proton
 448 operation [20]. To achieve this luminosity, IP2 would need to be operated at
 449 $\beta^* = 1.5 \text{ m}$ in runs 13,14. Even smaller values may be possible.

450 **2.4 Run-11 A_N DY performance**

451 The commissioning of the run-11 A_N DY apparatus involved eliminating single-
 452 channel issues (mostly associated with bad cables, bad connectors at patch
 453 panels, or loose connections), timing adjustments, photomultiplier tube volt-
 454 age adjustments to optimize gain uniformity, trigger development and checks,
 455 and confirmation of expected physics correlations within the data. Some of
 456 these items were addressed before collisions began and some required colliding
 457 beams to finalize. The strategy employed was to first optimize a minimum-
 458 bias trigger, since this provided the most important checks of the perfor-
 459 mance of HCal and ECal. After optimizing the minimum-bias trigger and
 460 recording large data samples, the development of more sophisticated triggers

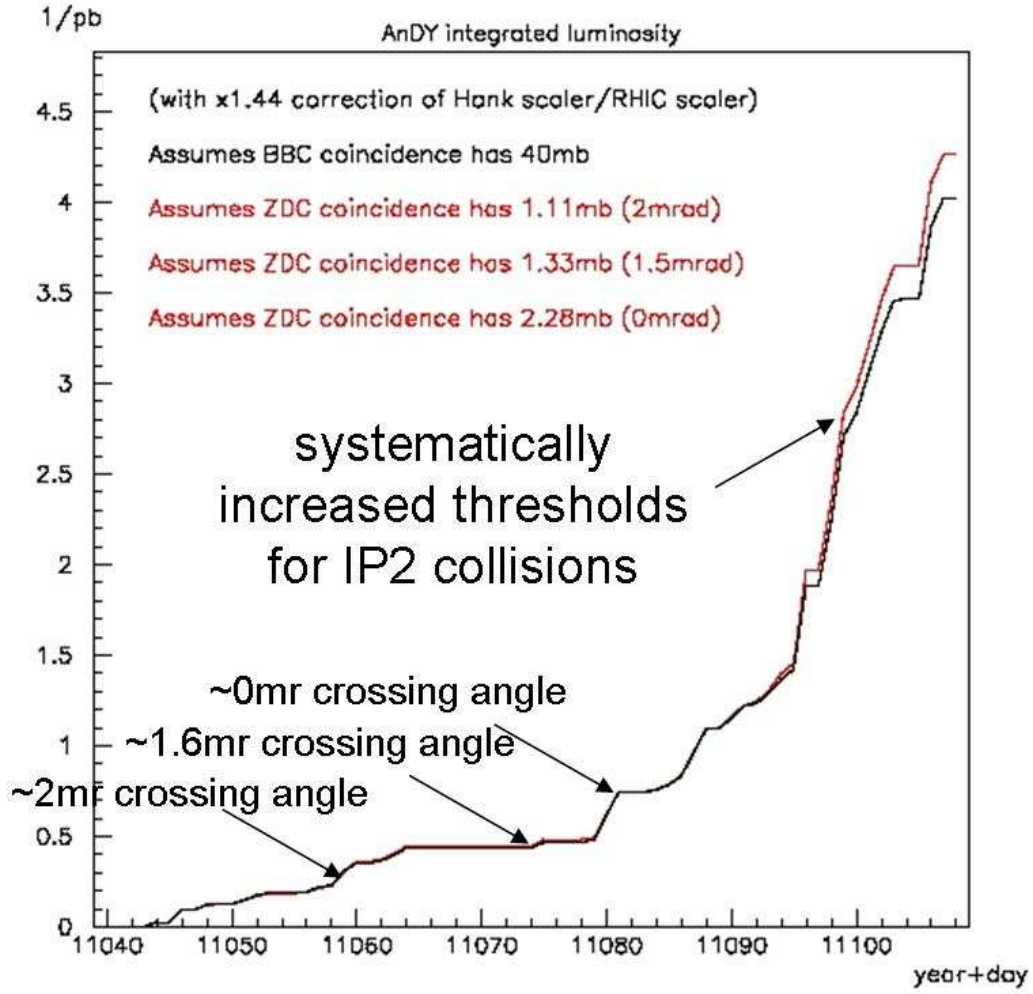


Figure 11: Integrated luminosity at IP2 versus time, as deduced from the number of coincidences from the IP2 beam-beam counters and the zero-degree calorimeters. The effective cross sections used to scale the coincidence count to integrated luminosity are from simulation, as described in the text. The effective cross section from a vernier scan at IP2 during RHIC fill 15457 is 7% smaller than deduced from the simulation, based on online analysis. Systematic errors in the integrated luminosity measurement must still be estimated, but are likely associated with backgrounds in the detectors and RHIC parameters for the store in which the vernier scan was completed.

461 proceeded.

462 The minimum bias trigger optimization involved gain adjustments for the
463 BBC detectors, timing adjustments at the QT boards (see Fig.44), optimiza-
464 tion of the time-to-amplitude signals for the BBC and ZDC detectors from
465 which the collision vertex z position is determined by time differences, and
466 fixing the parameters for the minimum-bias trigger. The primary focus for
467 more sophisticated triggers was on a jet trigger, corresponding to a threshold
468 on a masked sum of HCal detectors from a single module. Its development
469 required optimizing the timing of pulses of the QT integration gate with re-
470 spect to the pulses from particles produced by collisions, setting masks to
471 select jets whose thrust axis was centered within a module, and then setting
472 thresholds. Other triggers used frequently through the run were a module-
473 summed ECal trigger that preferentially selects events that have energetic
474 forward neutral pion production, LED monitoring triggers to monitor the
475 gains of ECal and HCal, a zero-degree calorimeter trigger used for verifying
476 that the A_N DY apparatus could measure a known analyzing power, and a
477 zero-bias trigger used to configure the pedestal corrections made on the QT
478 boards to have robust zero suppression of the data.

479 Fig. 12 shows individual detector ADC distributions for the Blue-facing
480 BBC from minimum-bias triggers. The large peak in each distribution corre-
481 sponds to a single minimum-ionizing particle passing through the detector.
482 The MIP centroid was used to measure gain curves for each phototube and a
483 single iteration in the high-voltage adjustments was made for gain matching
484 the detectors. MIP peak resolutions are comparable for most detectors. The
485 width is primarily associated with photostatistics and phototube resolution,
486 as established by offline comparison to full PYTHIA/GEANT simulations.
487 Evident for most detectors are peaks corresponding to two MIPs passing
488 through one detector in the same event.

489 Fig. 13 shows a vertex- z distribution with and without trigger level cuts.
490 The vertex- z position is proportional to the time difference of the earliest hits
491 on the Blue-facing and Yellow-facing BBC annuli. Hit timing is measured
492 by integrating a current produced by a time-to-amplitude converter (TAC),
493 started when a BBC detector crosses a preset threshold and stopped by the
494 RHIC clock, meaning that large TAC values are early hits and small TAC
495 values are late hits. Time difference is computed by the trigger FPGA coding
496 by first identifying the earliest hit detector on a QT board that has both
497 analog inputs and TAC inputs from each detector in the same 70-ns gate
498 (so-called “max-TAC algorithm”). The vertex z position is proportional to

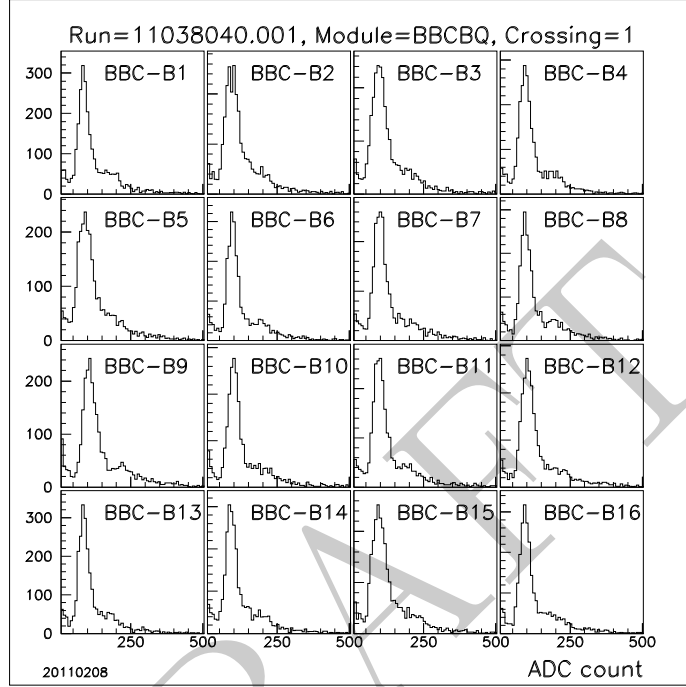


Figure 12: Charge distributions for Blue-beam-facing BBC for minimum-bias events that include a timing cut on the collision z distribution. Detector numbers increase in the clockwise direction around the annulus, when looking towards the interaction point. Gain matching was complete for this run, as were all final timing adjustments. The minimum bias trigger required the summed charge for both the Blue-beam and Yellow-beam facing BBC arrays to exceed 60 ADC counts, and required that the difference between the time-to-amplitude converter channels (see Fig. 13) for the first hits in Blue and Yellow arrays be smaller than ± 400 counts. The large peaks in each detector correspond to a minimum-ionizing particle passing through the detector. Most detectors also show a well defined peak from events where two minimum-ionizing particles pass through that detector.

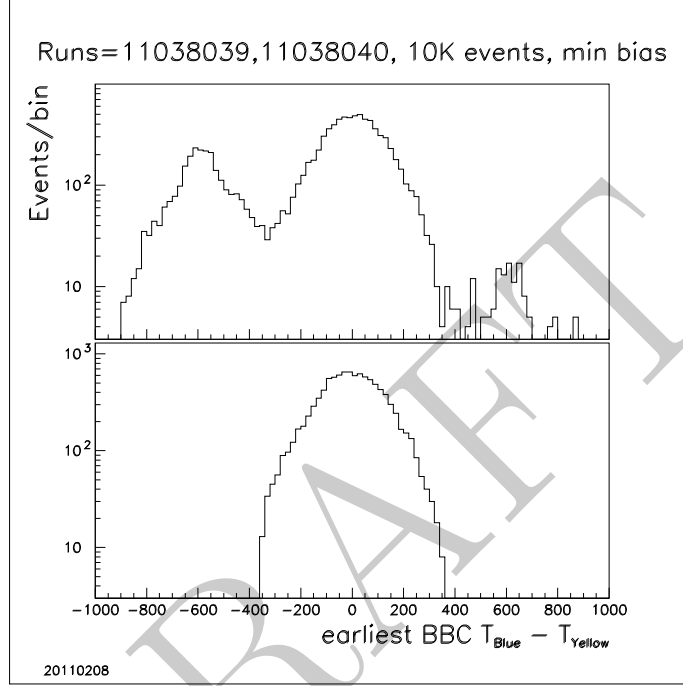


Figure 13: Distributions of the difference of earliest-hit time-to-amplitude (TAC) converter values from the Blue-facing and Yellow-facing BBC annuli. The top distributions do not include a cut on the TAC difference. The central peak is from collisions and is narrow for this run due to the $\sim 2\text{mr}$ bunch-crossing angle at IP2. The peak at negative values is for single-beam backgrounds induced by the Yellow beam. The small peak at large positive values is for single-beam backgrounds induced by the Blue beam. Single-beam background contributions varied from fill-to-fill, as monitored by event data without the TAC difference cut.

the difference between the Blue-facing and Yellow-facing BBC, as confirmed in offline analysis that uses this vertex for neutral pion reconstruction. This TAC difference cut is computed in an FPGA on the vertex DSM after valid data is transferred from the Blue-facing and Yellow-facing QT boards. The TAC difference discriminates collision events from single-beam background events. For collision events, the TAC difference is proportional to the vertex-z position.

During commissioning and to test data integrity, the DAQ was set up to record dedicated minimum-bias collision triggers at a rate of ~ 2.4 kHz with essentially 100% dead time, with the rate limited by recording data from three clock ticks. For each event, data was acquired from the BBC, the ECal, the HCal, the preshower and the ZDC for the RHIC clock tick for which the trigger conditions were satisfied, the RHIC clock tick before the bunch crossing with the trigger (so-called “pre” crossing) and the RHIC clock tick after the bunch crossing with the trigger (so-called “post” crossing). Having the triggered, pre- and post-crossing data for each detector is a powerful means of identifying detector effects (*e.g.*, bad ground connections can produce pulse reflections or oscillations that persist beyond the 70-ns gate) and beam-induced backgrounds.

When the commissioning of A_NDY was complete, post-crossings typically had very small charge coming primarily from afterpulsing in the photomultiplier tube and pre-crossings were generally consistent with pedestal values. This is illustrated in Fig. 14 for the HCal modules. Systematic effects were studied for most fills by collecting data with and without the minimum-bias collision condition imposed in addition to the jet trigger. The resulting patterns of hit cells illustrates that the events are manifestly jet like. **Comic-ray** events (taken when there was no beam by requiring the HCal module sum to satisfy a 30-count threshold) showed clear tracks through the detectors (Fig. 15), confirming that all mappings in HCal were correct and establishing that the “jet-like” patterns for colliding beam events were associated with physics.

The same method of forming module sums and setting a threshold on the resulting sum was used to make a trigger on the ECal modules and a trigger on the ZDC. This is illustrated in a slightly different fashion in Fig. 16. The ECal module-sum trigger that we dubbed as the “ π^0 ” trigger is common to all distributions. Shown in the figure are the BBC sums and the HCal sums, subject to this trigger. The biases are evident: relative to minimum bias events where BBC sums are symmetric, the Blue-facing BBC sees additional

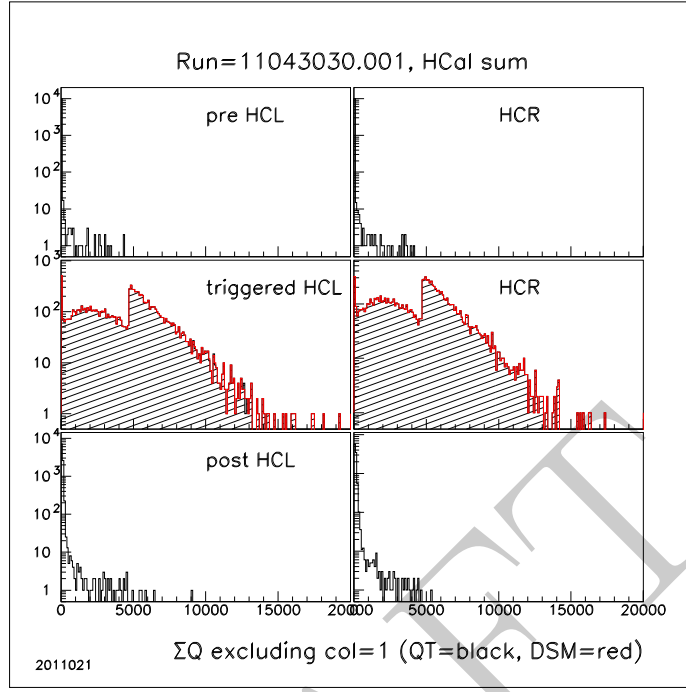


Figure 14: Charge distributions from the HCal jet-trigger. The jet trigger requires a threshold on the masked sum of ADC values from the 108 detectors in each module. For this data, the first column of detectors was masked from the sum, since the most interesting events have the thrust axis of the jet centered in the module. The black-shaded distributions are formed by summing the individual detector ADC values, subject to the same masking as used for the trigger. The red histogram is from the input to a DSM and contains the masked sum computed by the FPGAs. A valid event for the jet trigger has the DSM sum for either the beam-left (HCL) or beam-right (HCR) HCal module crossing a preset threshold. The top row of plots is for bunch crossings that precede the triggered bunch crossing by 107 ns and the bottom row of plots is for bunch crossings that follow the triggered crossing by 107 ns. The “pre” crossing distribution is mostly the 96-channel pedestal sum, with $<1\%$ of the events having charge (likely from room backgrounds). The “post” crossing distributions are again dominated by the pedestal, with a small probability of afterpulsing from the phototubes on HCal. The jet trigger was commissioned on day 46 of the run and was used to record data from essentially all of the luminosity delivered to IP2 (see Fig. 11). When the instantaneous luminosity was increased at IP2, the live time of the jet trigger was kept high by first changing the masking to exclude the outer two perimeters of detectors in a module, and then by raising the threshold.

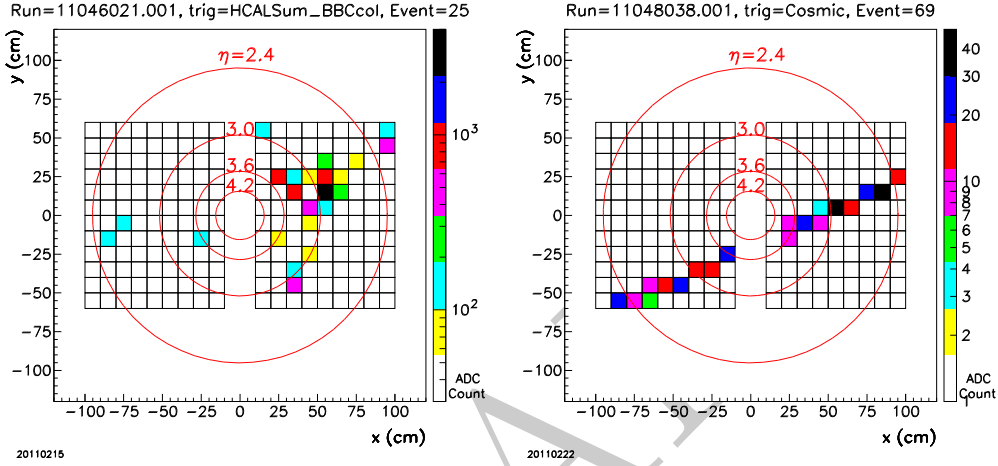


Figure 15: Event displays for the HCal arrays. This display represents the HCal modules as arrays as observed from the interaction point. The red circles show pseudorapidity for events with vertex- $z=0$. Colored boxes have hits in the corresponding detectors, with the color scale set by the ADC value recorded for that detector. (Left) an event from collision data acquired with the jet trigger, showing clear “jetty” structure. (Right) an event acquired two days later without beam with an analogous module-summing algorithm and a low threshold on the sum. This event clearly shows a cosmic-ray muon passing through both modules. An ensemble of cosmic-ray events demonstrates that the mapping of detectors is robustly understood. In total $> 7.5 \times 10^8$ jet events were recorded from the $\sim 6.5 \text{ pb}^{-1}$ of integrated luminosity delivered to IP2. That data will provide first determination of the forward jet production analyzing power.

charge reinforcing the notion that forward neutral pion production at this pseudorapidity are fragments of forward jets. As well, the HCal module sums differ from minimum bias events, having a slower falloff in the event yield as the charge sum increases. This is expected if the neutral pion is accompanied by additional particles associated with a forward jet.

Triggering on the ZDC charge sum is of particular interest, since it is known that far forward neutron production has a non-zero analyzing power [21]. Far-forward neutron production is used for local polarimetry at STAR and PHENIX for $\sqrt{s} = 500$ GeV polarized proton collisions as a means of tuning the spin rotator magnets. Local polarimetry is of interest to many when the beams have vertical polarization. Fig. 17 shows the ZDC charge sum for events where either the Blue-facing or the Yellow-facing ZDC is above a preset threshold. A scintillator-strip shower maximum detector was built for the IP2 ZDCs, from an existing glued plane of scintillator strips recycled from an earlier forward calorimeter project. Scintillation light is collected by a wave-length shifting optical fiber inserted into a centered hole that is made during the extrusion of the triangular cross section strip. The fibers are then mapped onto pixels of a 16 channel Hamamatsu multi-anode photomultiplier tube. The charge from each of the 16 anodes is encoded for event readout. The ZDC shower maximum detector is installed in between the second and third modules of the ZDC. The purpose of the strips is to measure a θ, ϕ for the neutron in the event. The ϕ dependence of the spin asymmetries then provides information. The amplitude of the ϕ dependence is the analyzing power. The phasing of the harmonic dependence of the spin asymmetry on ϕ can provide information about both vertical and radial polarization components.

Fig. 17 shows the module sum distribution for the ZDC and the resulting analysis of spin effects from event data. The event analysis amounts to forming distributions of the number of events in a given ϕ bin for each bunch crossing. The distributions are incremented subject to requirements on the ZDC module sum, the threshold of relatively gain-matched charge in the ZDC SMD strip, and a large enough θ for the event. The polarization pattern is retrieved from a database archived in essentially real time from information broadcast by RHIC. This analysis methodology was then easily extended to bunch-crossing scalars. With suitable bit input to such scalars, the relevant distributions are accumulated at a maximum event rate of 9.38 MHz, with zero deadtime. Offline analysis provides azimuthal dependence of spin asymmetries. The plan is to perform real time analysis of such asym-

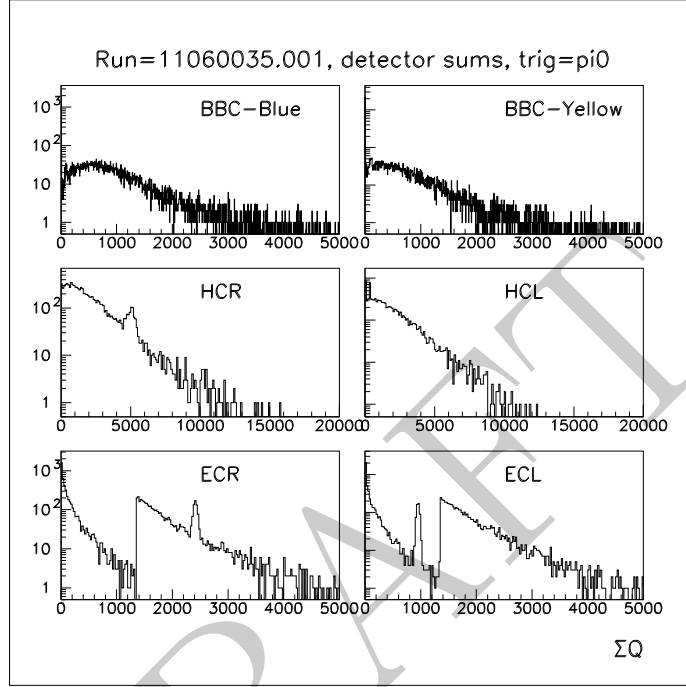


Figure 16: Module sums from the BBC (top), the HCal (middle) and the ECal (bottom) subjected to the neutral pion trigger (threshold on ECal module sum, as evident in the bottom row of plots). No additional conditions are on the data. The narrow peaks in ECL, ECR and HCR are from the light-emitting diode that also satisfies this trigger. The bias from the neutral pion trigger is evident in both the BBC and the HCal. For the BBC, more charge is observed in the Blue-facing BBC than for minimum-bias events, as expected if the neutral pion is accompanied by additional particles. The HCal sum distributions fall off less rapidly with increasing charge sum than for minimum bias, reflecting the additional particles that accompany the neutral pion.

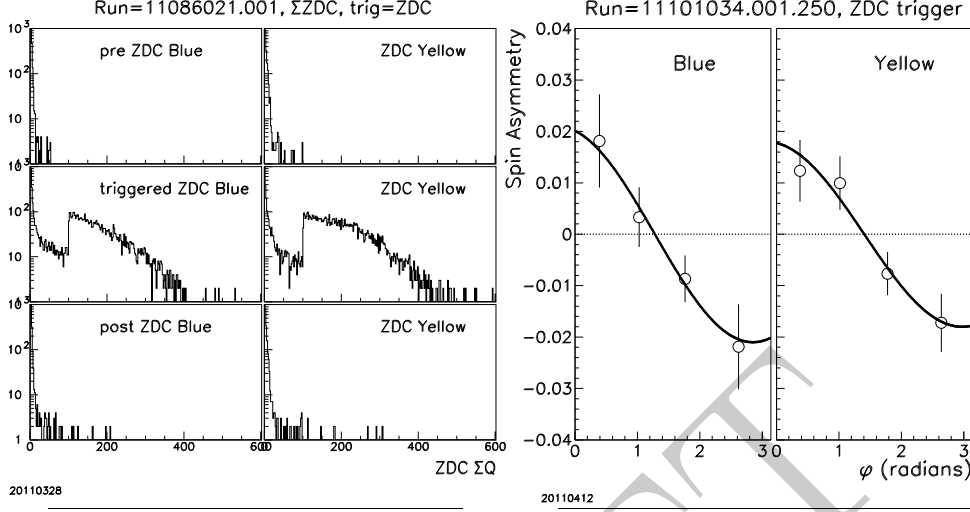


Figure 17: (Left) Module summed distributions from the ZDC subject to a trigger that either the Blue-facing or the Yellow-facing ZDC module sum is above threshold. Calibrations of the ZDC module relative gains were not possible until $\sqrt{s_{NN}}=200$ GeV Au+Au collisions were made at IP2. As for Fig. 14, distributions of the ZDC module sum response for 107-ns before the triggered crossing and 107-ns after the triggered are useful to assess backgrounds. (Right) Spin asymmetry from the Blue- and Yellow-facing ZDC, as analyzed from event data.

metries during RHIC run 13?

The ZDC modules were new to IP2 and were not absolutely calibrated prior to run 11. Consequently, A_{NDY} requested $\sqrt{s_{NN}}=200$ GeV Au+Au collisions so that the ~ 100 GeV neutron flux produced at 0° could be used to calibrate the ZDCs. IP2 was brought into collisions for two stores, after STAR and PHENIX had accomplished their goals. Fig. 18 shows the neutron peaks obtained after high-voltage adjustments were made to the IP2 ZDCs to calibrate the detectors. It will be necessary to look back at the ZDC event data from the polarized proton run to establish the impact of gain mismatching on the spin asymmetries.

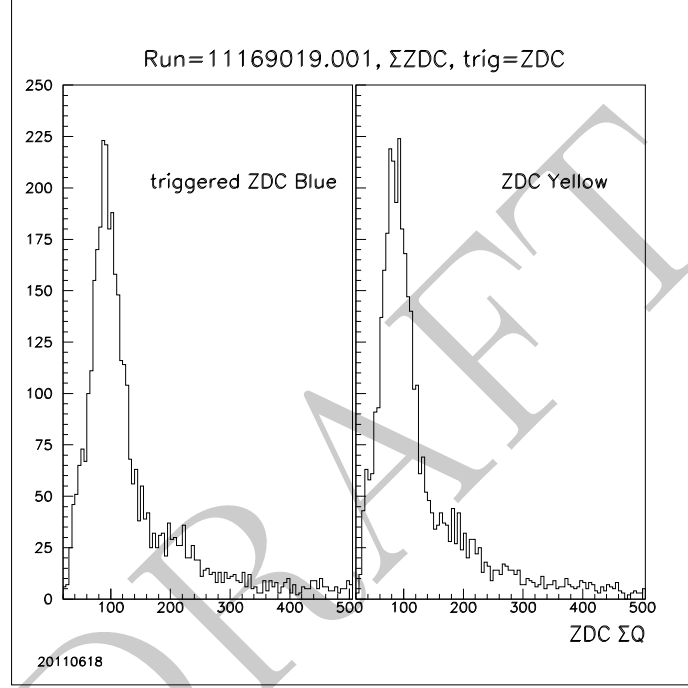


Figure 18: ZDC summed charge distributions obtained during $\sqrt{s_{NN}}=200$ GeV Au+Au collisions at IP2. These collisions were used to calibrate the three separate modules of the Blue-facing and Yellow-facing ZDC at IP2. Such calibrations are difficult in the absence of the peaked response from the ~ 100 GeV neutron beam made at 0° by ultraperipheral Au+Au collisions.

2.5 Offline analysis

Offline analysis of the run-11 data was begun during the polarized proton run and has continued since. The analysis coding is an adaptation of coding developed for other forward calorimetry projects. Changes to this coding specific to A_N DY include an interface to the overall data structure and the specifics of the mapping of the raw data to the detector channels. The latter was greatly simplified by straight mappings of the connections to the readout electronics. At an early stage in developing this analysis coding, mapping of the PYTHIA/GEANT simulation output to become pseudodata was done so that the simulated pseudodata could be analyzed by the same coding as used for the real data. This method greatly facilitates comparisons of real data to simulation. Such comparisons are essential for a calorimeter based experiment to ensure that detector responses are understood and to assure that features in mass distributions are real rather than created from sculpted combinatoric backgrounds.

Association analyses can establish whether features in mass distributions are real. Once the basic data structures and mappings are in place, algorithms that have been well developed over the past ten years [22] can be applied to form energy clusters for the events and to analyze these clusters to attribute an x, y position to the cluster and to reconstruct the relativistic energy of the particle that initiated the calorimeter response from the energy of the cluster. Given that there is no magnet in the run-11 A_N DY configuration, particles produced in the collision follow straight-line trajectories from the event vertex to the x, y impact point at the calorimeter. It is necessary to attribute the z location of where the x, y measurements are made. The z position of the event vertex is reconstructed for each event from the time difference measurement (Fig. 13), and the x, y location of the event vertex is assumed to be zero. With this line and the total relativistic energy, the four momentum is known up to identifying the type of particle that initiates the calorimeter response. The transverse attributes of the cluster are useful for discriminating incident photons, electrons or positrons from incident hadrons. A long-term goal of A_N DY is to robustly discriminate electrons or positrons from hadrons and photons by cross correlating multiple attributes of matched ECal and HCal clusters that are then also matched to the preshower response.

The calibration of the calorimeters is based on reconstruction of particles produced in the collision by identifying peaks in relevant mass distributions.

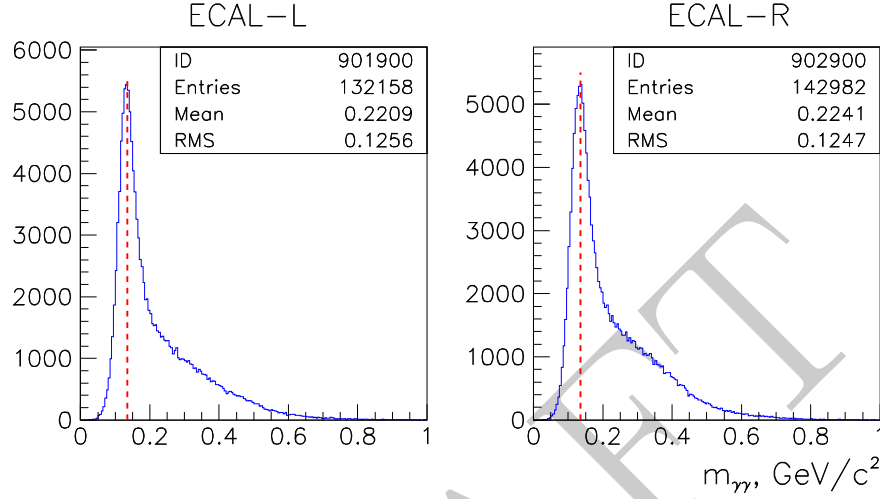


Figure 19: Cluster pair invariant mass from the run-11 A_NDY ECal modules after seven iterations of adjustments to the individual detector gains. A clear peak is observed from $\pi^0 \rightarrow \gamma\gamma$, whose identification is based on prior experience with similar calorimeters where extensive intercomparisons to simulations had been made. Such comparisons for A_NDY have been limited so far to energy spectra from the individual detectors. When comparisons are more complete, it is expected that the tail beyond the neutral pion peak will be identified as the ECal response to incident jet-like multiplicity. Online mass reconstructions allowed adjustments of detector high voltages based on offline gain corrections and measured gain curves for each ECal detector. Two such iterations were completed. More are required for hardware-level gain calibration.

622 This can correspond to cluster pairs, 3-cluster combinations, or more complex
 623 configurations. Association of the reconstructed mass for the event with a
 624 detector identified by the position of the leading cluster enables a means
 625 of establishing both relative and absolute gains for the calorimeter. This
 626 approach does require an iterative solution, especially when the cluster energy
 627 is spread among several detectors. Nonetheless, experience has shown this is a
 628 robust calibration method. For ECal, one step is taken after the clustering.
 629 This step involves fitting the parameters of shower shape functions whose
 630 basic form is determined from test beam data [23]. The fit results are then
 631 used in place of moments of the clusters. The result for the ECal stacks in
 632 place for the run-11 A_N DY configuration is shown in Fig. 19.

633 A primary goal of run-11 as noted in the introduction was to establish a
 634 robust calibration procedure for HCal, for which A_N DY proponents had no
 635 prior experience. A basic clustering algorithm (identifying groups of HCal
 636 towers with energy deposition above a threshold, bounded by towers at or
 637 below this threshold) was applied to the data (Fig. 20), albeit with many
 638 studies that varied the threshold of the cluster boundary, since the HCal
 639 matrix is formed from transversely large cells ($10\text{cm} \times 10\text{cm}$) thereby result-
 640 ing in greater chance for cluster merging. Furthermore, the transverse size
 641 of hadronic clusters is intrinsically larger than those made by incident pho-
 642 tons, electrons or positrons. A complete association analysis of pseudodata
 643 from PYTHIA/GEANT simulations was an essential tool in the studies that
 644 culminated in determination of the HCal energy scale.

645 The association analysis amounts to using the primary particle four mo-
 646 mentum from the PYTHIA event generation and the simulated vertex po-
 647 sition for the event to reconstruct masses. Straight-line propagation of all
 648 particles from the vertex is done to determine their possible impact point
 649 at the HCal. This impact point is then compared to impact points deduced
 650 from the clustering analysis of the full pseudodata. If the impact points
 651 are within proximity, a match is declared and the identity of the cluster is
 652 then known. Fig. 21 shows the results from the proximity-based association
 653 analysis for special conditions on the clusters. Namely, the clusters required
 654 in this analysis had only a single tower, bounded by other towers that were
 655 below an energy deposition threshold of 0.08 GeV , and the towers had to be
 656 beyond the shadow cast by ECal. This event selection emphasizes incident
 657 photons. The cluster position requirement is imposed since ECal is built from
 658 16 radiation length lead glass bars, meaning that most of the incident photon
 659 energy is converted into electromagnetic showers in ECal. The association

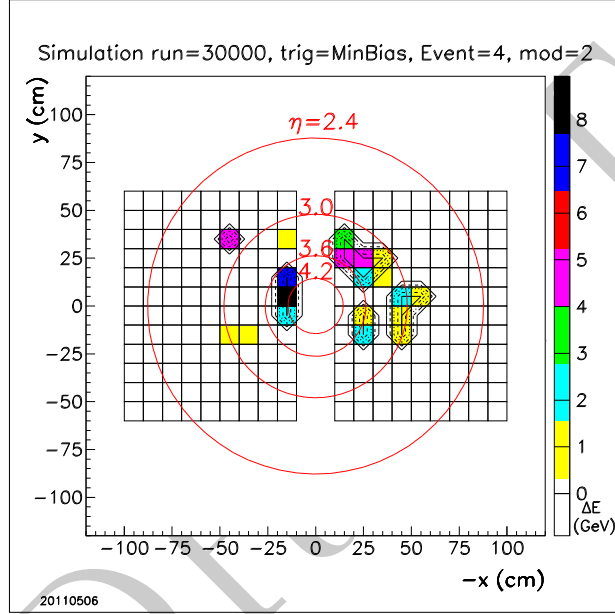


Figure 20: Single event display of pseudodata for HCal from PYTHIA/GEANT simulations and the contours representing the threshold-bounded clusters identified in the event. There are many examples where clusters are merged due to the coarse cell sizes that require careful tuning of the threshold for the cluster boundary

660 results for the position have the expected form. The matching of the cluster
 661 centroid (taken to be the center of the detector cell, since these clusters have
 662 no energetic neighbors to enable a position interpolation) to the impact point
 663 from the PYTHIA primary is essentially a truncated uniform distribution.
 664 The truncation is at ± 5 cm as set by the Moliere radius of electromagnetic
 665 showers in HCal detectors and the requirement for single-tower clusters. Also
 666 shown in Fig. 21 is the pair mass distribution from PYTHIA/GEANT sim-
 667 ulations, where the association analysis allows identification of the incident
 668 particle that deposited the energy. It is clear from this analysis that neutral
 669 pions are reconstructed in the cluster pair mass distribution. This is not
 670 so surprising, since neutral pions are by far the dominant source of photons
 671 produced in p+p collisions at $\sqrt{s} = 500$ GeV. What may be somewhat sur-
 672 prising is that the photon showers are well measured in both position and
 673 energy. Most hadron calorimeters are shadowed from collision vertices by
 674 electromagnetic calorimeters, so cannot see photons produced by collisions.
 675 The fiber density in the HCal detectors is large (47×47 matrix of 1-mm
 676 diameter scintillating fiber [13]), therefore allowing a good measurement of
 677 electromagnetic showers produced by incident photons.

678 The same analysis method as used for the simulation is then applied
 679 to 20 million minimum-bias triggered events, that include a vertex cut, ac-
 680 quired from three different RHIC fills. The resulting mass distributions are
 681 then overlayed with analogous distributions reconstructed from 20M non-
 682 elastic, non-singly diffractive events generated by PYTHIA 6.222 that are
 683 run through GEANT. No scaling factors are applied to any of the distribu-
 684 tions for the overlay shown in Fig. 22. There is excellent agreement between
 685 data and simulation. The mass peaks in data are then quantitatively ana-
 686 lyzed to establish that the absolute energy scale of HCal is determined, at
 687 present, to $\sim 5\%$.

688 A summary of the run-11 effort is that the primary goals set in the 2010
 689 letter of intent [2] have been accomplished.

- 690 • Collisions of the intense polarized proton beams can occur at IP2 with-
 691 out significant impact on operations at IP6, IP8. The integrated lumi-
 692 nosity projected for runs 13, 14 at IP2 while IP6 and IP8 are acquiring
 693 events for their measurements of parity violation for W production is
 694 sufficient for the first transverse spin Drell-Yan production experiment,
 695 using the apparatus we propose to build as described later in this doc-
 696 ument.

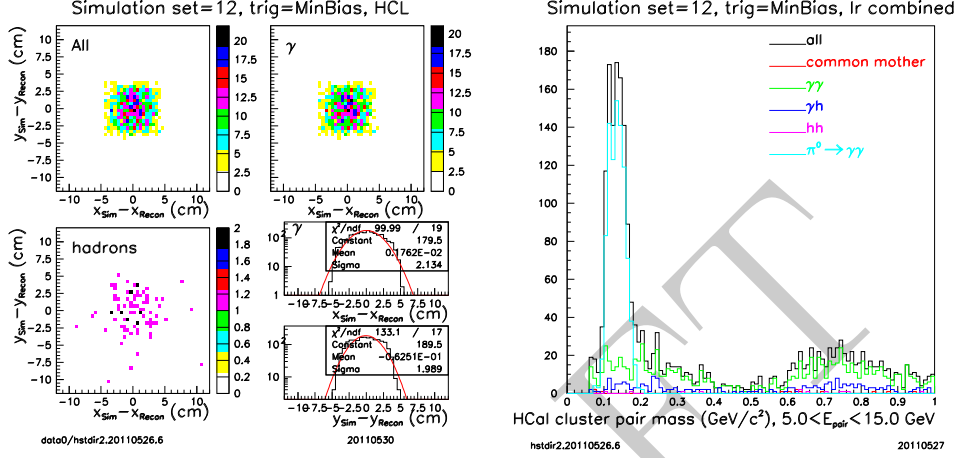


Figure 21: (left) Results from the proximity-based association analysis of full PYTHIA/GEANT simulations showing the position differences between projections of particles generated by PYTHIA to reconstructions of PYTHIA/GEANT simulations. This step is preceded by an intercomparison of data and simulation that establishes that the PYTHIA/GEANT model gives a good description of the data, as expected from such comparisons made with other forward calorimeter projects done at RHIC. The reconstructed clusters are subjected to event selections that significantly suppress the hadronic response: single-tower clusters bounded by towers with $\Delta E < 0.06 \text{ GeV}$ with an addition requirement that the x, y location of the cluster is outside of the shadow cast by ECal. (right) The selected clusters are then converted into four momenta of incident particles, assuming they are produced at the event vertex and that they are photons, using a scaling factor to convert cluster energy into incident total relativistic energy. Invariant mass is then computed from inclusively pairing all such clusters for the event. The association tag then allows a decomposition of the pair mass spectrum to show that the peak at $0.135 \text{ GeV}/c^2$ is due to $\pi^0 \rightarrow \gamma\gamma$ decay.

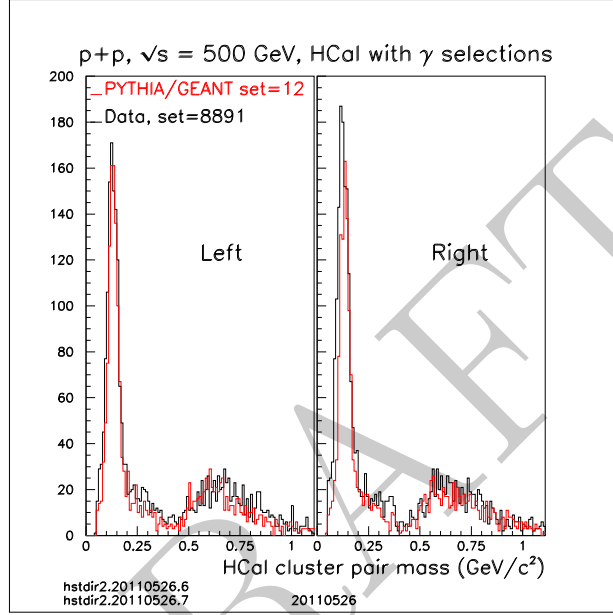


Figure 22: Comparison of pair-mass distributions reconstructed from data with those reconstructed from simulation for clusters subjected to photon-like requirements described in the text and in the Fig. 21 caption. The peak near $0.135 \text{ GeV}/c^2$ is from $\pi^0 \rightarrow \gamma\gamma$ events and establishes that the energy scale of HCal is known to $\sim 5\%$, as confirmed by an association analysis of the simulations (Fig. 21). The near quantitative agreement between data and simulation is a strong indication that the PYTHIA/GEANT model gives a good description of the data. The peak near $0.6 \text{ GeV}/c^2$ is from sculpted combinatoric backgrounds as shown by the association analysis!

- The HCal modules configured for run-11 have been calibrated to $\sim 5\%$. Further work on calibration is required to establish that the hadronic showering model in GEANT properly accounts for the deviations from a perfectly compensated calorimeter.

When that work is complete, jet clustering algorithms can be applied to the data and a jet-energy scale will be established. The first measurement of the analyzing power for jet production will then be possible.

The first comparisons of PYTHIA 6.222/GEANT simulations to the data show good agreement. This supports that the QCD background estimates made for the A_N DY proposal (Appendix 1) are valid. Further work is required here to benchmark the simulations. The run-11 data set can provide some indication of the spin-dependence of the QCD background to the Drell-Yan production signal that later runs will identify. The matching of clusters between ECal and HCal for the run-11 data is an important step for the benchmarking.

Any concerns about significant differences in single-beam backgrounds being worse at IP2 than at IP6 have been laid to rest by the run-11 data. The background environment in this forward region is well controlled at RHIC by the cryostats containing the ring magnets and by the collimation system. This provides RHIC a unique opportunity to make the first transverse spin DY production measurement. Although backgrounds were not found to be large, attention will be required in future runs to optimize single-beam backgrounds at IP2. The focus on run-11 was simply to establish collisions and precluded significant attention to single-beam background optimization.

The run-11 A_N DY experience has also established that the analysis tools developed over the past ten years are flexible and can be readily morphed into different calorimetric configurations. We propose to construct such a different configuration in run-13. Care and effort will be required to assure that the calorimeter detectors are well constructed and properly mapped to the electronics. The experienced group that has been formed for A_N DY is fully cognizant of the quality requirements and is capable of achieving them. Reconstruction coding will be an essential component of commissioning, so the flexibility demonstrated from run 11 bodes well for the future. With all of this said, the A_N DY goals for run-13 of clear observation of the $J/\psi \rightarrow e^+e^-$ and $\Upsilon \rightarrow e^+e^-$ decays and the intervening di-electron mass distribution from a newly constructed experiment are indeed ambitious. The A_N DY proponents believe the implementation described below for run-13 is robust

734 and can be completed, although we are fully cognizant of the significant effort
 735 that will be required. The project risks are understood by the proponents
 736 through the best possible design tool: namely model building. The run-11
 737 implementation can be viewed as the model that demonstrates the viability
 738 of the proposed configuration for run-13.

739 2.6 Dilepton Reconstruction from Run-11 Data

740 The robust reconstruction of $\pi^0 \rightarrow \gamma\gamma$ in both the ECal and HCal invites the
 741 question of whether A_N DY methods can be applied to run-11 data to extract
 742 a di-lepton signal; *i.e.*, $J/\psi \rightarrow e^+e^-$. Such a signal can provide confidence
 743 that the proposed A_N DY apparatus will be capable of measuring Drell-Yan
 744 production.

745 It must be kept in mind that the run-11 apparatus has limitations compared to what we propose here:
 746

- 747 • The preshower that we propose has substantially more granularity than
 748 what was in place for run 11. Consequently, background suppressions in
 749 run11 will not match expectations from GEANT studies in the A_N DY
 750 proposal (see Section 5). Indeed, the Blue-facing beam-beam counter
 751 turns out to provide the most critical background suppression, as de-
 752 scribed below. The proposed preshower for A_N DY is ten times more
 753 granular.
- 754 • The acceptance for dileptons in the run-11 apparatus is predominantly
 755 from the HCal. Electromagnetic showers are preferentially selected
 756 from HCal by considering single-tower clusters, bounded by towers with
 757 small energy deposition. The position resolution for HCal clusters is
 758 at least $10\times$ worse than for ECal, due to the coarse granularity. This
 759 poor position resolution impacts the ability to identify which preshower
 760 detector to query for the photon/electron discrimination.

761 Even with these caveats, it is instructive to proceed, given the prospects
 762 of understanding better how the proposed apparatus will perform. As de-
 763 scribed below, we observe $\sim 120 J/\psi \rightarrow e^+e^-$ events (Fig. 3) from a data
 764 sample consisting of 3.5×10^7 triggers. The factor of 3×10^5 event suppres-
 765 sion represents a major part of what is required to access DY events.

766 Data obtained with the ECal trigger (Fig. 16) are the natural source
 767 for attempts to extract a di-lepton signal from the run-11 data. Only a

limited data sample was acquired with this trigger, since the focus for run-11 had been on the jet trigger (Fig. 14). To accompany this data, the run-11 GEANT model was used to simulate detector responses from 0.5 pb^{-1} of integrated luminosity from non-elastic, non-singly-diffractive PYTHIA 6.222 events. These events were first filtered by requiring that the summed total relativistic energy of particles projected into the ECal acceptance exceed 10 GeV. Prior experience has shown that such filters are crucial to generating simulation samples that can be compared to data. Computer resources (CPU and disk space) are minimized by filtering events before they are run through GEANT.

The first step towards this goal is to compare data and simulation for the HCal response. Such comparisons were made for the jet-triggered events [24]. In general, simulation is found to explain data distributions. The events are dominated by QCD jets, as expected. The data/simulation comparison shows that the neutral pion based calibration of the HCal provides a good starting point, although $\sim 20\%$ corrections for hadron initiated showers are still required.

The next step is to extend the association analysis discussed in Sec. 2.5 to the ECal+HCal pair mass distribution, so as to understand the composition of the background. The clusters considered from the HCal for this pair mass distribution are described above. HCal clusters are excluded from consideration if they can be paired with a second HCal cluster to yield an invariant mass $M < 0.2 \text{ GeV}/c^2$. The simulation is used to determine the primary particles that give rise to the large-mass cluster-pair background (Fig. 23), using proximity association described above. It is found that the background is dominantly from photons. This is not so surprising given that imposition of a minimum x_F in this rapidity interval kinematically suppresses jet pairs. The two-photon decay of the neutral pion is a way to generate pairs at large mass through the opening-angle distribution for this decay.

Given the photon dominance of the background, the beam-beam counter provides the primary means of distinguishing electrons and positrons from photons, since this array is the first active element particles see upon exiting the beam pipe. Photon conversions and electron/positron **brehmsstrahlung** in the BBC means that the preshower detectors see partial showers, so are unable to make the necessary discrimination. The requisite discrimination is established by requiring the BBC element matched to the projected position of the cluster from either ECal or HCal to have a charge corresponding to a minimum-ionizing particle (*i.e.* the large peaks in Fig. 12). The

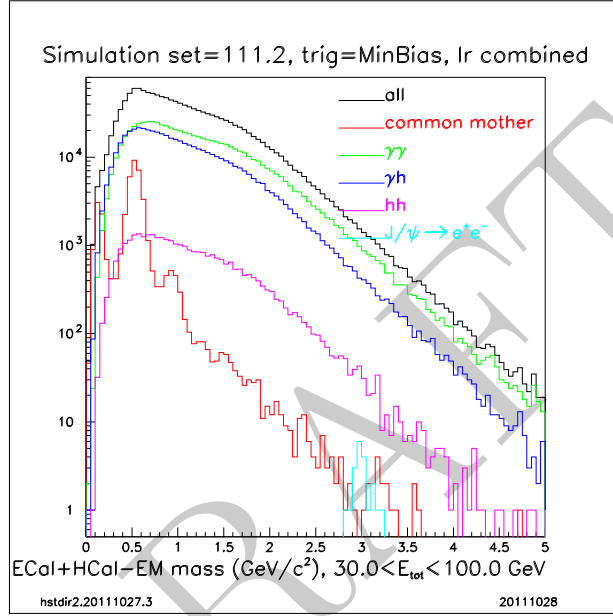


Figure 23: Association analysis of the mass distribution when an ECal cluster is paired with an HCal cluster. Filtered PYTHIA 6.222 simulations require summed total relativistic energy > 15 GeV for primaries projected from a vertex distribution into the acceptance of ECal. GEANT simulations provide the detector response for PYTHIA events meeting the filtering criteria. Reconstruction proceeds on these simulated events, followed by a proximity association of reconstructed clusters with projected primaries. The large-mass background is primarily from photons.

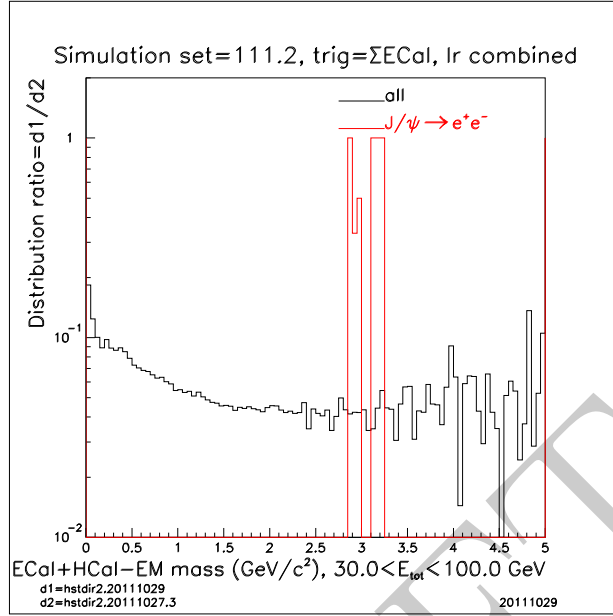


Figure 24: Ratio of the pair-mass distribution including the BBC requirement that clusters have a MIP-like response to the distribution without this requirement. The BBC requirement suppresses the background by a factor of 25, and mostly preserves the di-lepton signal from *Jpsi* decays.


806 PYTHIA/GEANT simulations show (Fig. 24) that the background is sup-
807 pressed by a factor of 25 after imposing the BBC requirement. It is also
808 evident in this figure that dileptons from J/ψ decay are mostly unaffected
809 by this requirement, although PYTHIA 6.222 grossly underpredicts the J/ψ
810 yield in this rapidity range, so the simulations have limited statistics for J/ψ
811 production relative to the data. The NRQCD model in PYTHIA 6.425 in-
812 cludes both color-singlet and color-octet matrix elements for onia production,
813 and has been found by the ATLAS collaboration to give a good description
814 of onia yields at the LHC [25]. Naively, we expect $100\times$ more photon sup-
815 pression than what is achieved in Fig. 24, based on single-particle studies
816 presented in the A_N DY proposal. The position resolution from the HCal
817 and the occupancy in the BBC both serve to reduce the effectiveness of the
818 discrimination by multi-particle effects: *e.g.*, a single BBC tile can have both
819 a charged hadron and a photon pass through it, thereby mimicking the re-
820 sponse expected for an electron or positron. The proposed A_N DY apparatus
821 remedies both position resolution and granularity issues.

822 The final step is to refine the calibrations, including the critical recon-

823 struction of the z -component of the vertex and the relationship between
824 deposited and incident particle energy for the HCal. For the z_{vertex} recon-
825 struction, single-beam background particles provide one calibration of the
826 TAC difference between Blue-facing and Yellow-facing BBC, assuming these
827 particles have $\beta \rightarrow 1$ and using the measured positions of the arrays. The
828 single-beam background measurement provides a $\sim 15\%$ larger calibration
829 constant than provided for by optimizing neutral pion masses, for cluster
830 pairs in both ECal and HCal. The source of this discrepancy is still unre-
831 solved. The z_{vertex} reconstruction plays a much more important role for J/ψ
832 reconstruction, since $\delta M/m \propto \delta z/z$. The HCal calibration was refined by
833 association of the mass distribution for each event with the leading cluster
834 of the pair. In that manner, peaks from $\pi^0 \rightarrow \gamma\gamma$ are observed for individual
835 HCal detector, except for those shadowed by the ECal module. Iteratively
836 adjusting calibration factors to assure that the peaks are at the physical
837 neutral pion mass was used for fine tuning the HCal calibration.

838 After these steps were taken, the cluster pair mass distribution observed
839 in the run-11 data obtained with the ECal trigger are shown in Fig. 3. A
840 clear peak from $J/\psi \rightarrow e^+e^-$ is observed in the data. A strong indication for
841 $\psi(2S) \rightarrow e^+e^-$ is also observed, as expected from earlier di-lepton studies.

842 The reconstructed $J/\psi \rightarrow e^+e^-$ yield provides ~ 120 events in the peak.
843 Accounting only for geometric acceptance of ECal and HCal and the energy
844 thresholds, 420 events are expected from the NRQCD model in PYTHIA
845 6.425 in a 0.5 pb^{-1} data sample. Although a full evaluation of the reconstruc-
846 tion efficiency is not available at this time, inefficiencies from the geometrical
847 acceptance of the BBC array and the clustering parameters used for HCal
848 are crudely estimated to provide 25% efficiency. The NRQCD model does
849 not include possible contributions from intrinsic charm components of the
850 proton wave function. The yield observed for $J/\psi \rightarrow e^+e^-$ is reasonable.

851 The NRQCD implementation in PYTHIA 6.425 can then be used to ad-
852 dress the ratio of $J/\psi \rightarrow e^+e^-$ production events to Drell-Yan production
853 events, subjected to the geometric and energy conditions used in Fig. 3. With
854 the further requirement that the virtual photon mass is larger than $4 \text{ GeV}/c^2$,
855 PYTHIA 6.425 predicts this ratio is 170. Consequently, further background
856 reductions of at least this magnitude are required from the proposed A_NDY
857 apparatus. There are two primary sources that can provide this requisite
858 reduction 

- 859 • The granularity of the proposed preshower detector and the elimination

of the position uncertainties in the association of calorimeter clusters with specific preshower detectors will enable us to approach the single-particle limit of 98% rejection of photons, from the current level of $\sim 80\%$ rejection. This is a further suppression of a factor of ~ 10 per cluster, or a factor of ~ 100 suppression of the pair mass background.

- When photons are sufficiently suppressed, hadron backgrounds will be the dominant background source. Estimates from the original A_N DY proposal are that $\geq 97\%$ of hadrons can be rejected from analysis of the longitudinal shower profile available from the HCal and preshower detectors. Given that the lepton daughters will provide independent rejection capability from this shower analysis, another factor of >1000 suppression of the pair mass background is possible.

The ratio of $J/\psi \rightarrow e^+e^-$ to Drell-Yan events for the ECal+HCal electromagnetic cluster pairs used in Fig. 3 also enables an assessment of how many DY pairs will reconstruct in the full A_N DY apparatus. The estimate is that the dileptons in Fig. 3 represent $\sim 10\%$ of what will be achieved with the full A_N DY apparatus. The acceptance limitations of ECal are the primary factor that will be addressed in going from the run-11 apparatus to the proposed A_N DY configuration. Another key is to lower the energy thresholds for the electron/positron observation. Extracting signal above background at lower energies is a key factor, beyond just the acceptance, to realize the 9400 DY events in a 100 pb^{-1} data sample.

Demonstration that $J/\psi \rightarrow e^+e^-$ can be extracted from QCD backgrounds in the run-11 data provides confidence that Drell-Yan can be extracted with the full A_N DY apparatus.

3 Electromagnetic Calorimeter (ECal)

The heart of the A_N DY apparatus is the BigCal electromagnetic calorimeter we have borrowed from JLab. Groups from Jefferson Laboratory (JLab) who had previously measured the electric form factor of the proton to high Q^2 using BigCal have joined the A_N DY project (see Fig. 25). The terms of the loan agreement are to return the calorimeter to JLab by 31 July 2014 for its planned use in experiments with the 12-GeV electron beam. This return date is non-negotiable and sets the timescale for A_N DY.

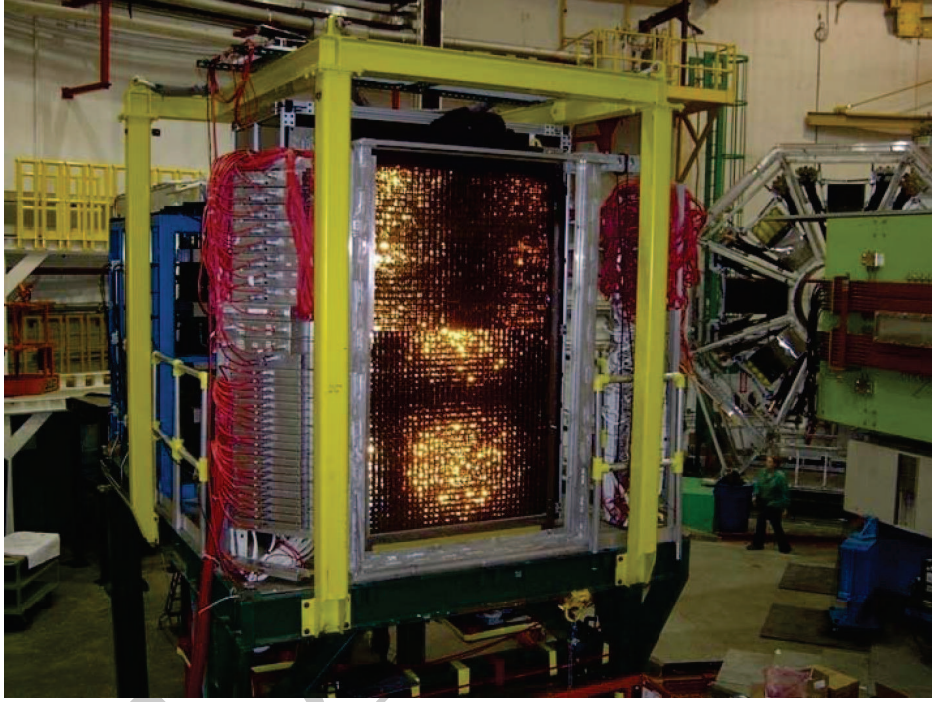


Figure 25: Photograph of BigCal at Jefferson Laboratory. The lead-glass calorimeter was constructed for measuring the electric form factor of the proton. It consists of a $24\text{-row} \times 30\text{-column}$ submatrix of $(40\text{ mm})^2 \times 40\text{ cm}$ lead-glass detectors atop a $32\text{-row} \times 32\text{-column}$ submatrix of $(38\text{ mm})^2 \times 45\text{ cm}$ lead-glass detectors. Each detector is viewed by an FEU-84 photomultiplier tube. In total, there are 1744 lead-glass detectors in BigCal. BigCal is on loan for the $A_N\text{DY}$ project until 31 July 2014.

893 The primary conclusion from our run 11 tests is that beams can collide
894 at IP2 with minimal impact on operations at STAR or PHENIX. The run-
895 11 experience has also allowed the Collider-Accelerator physicists to project
896 that $10 \text{ pb}^{-1} / \text{week}$ of luminosity can be delivered to IP2 for $A_N\text{DY}$ during
897 $\sqrt{s}=500 \text{ GeV}$ polarized proton operation. For a 10-week run, this is a smaller
898 data sample than initially assumed in the Letter of Intent [2]. Imposing a
899 requirement $x_F < 0.3$, as required to test the predicted sign change, also
900 results in fewer DY events. Consequently, an optimization of the $A_N\text{DY}$
901 acceptance has been done.

902 Multiple geometries for ECal have been considered in the optimization of
903 the $A_N\text{DY}$ acceptance. The best choice proves to be an azimuthally complete
904 ECal implemented by two stacks of lead-glass detectors, each on a base plate
905 supported by bearings sliding on Thomson rails. The motion of each stack
906 is transverse to the beams. The azimuthally complete ECal is realized when
907 the two stacks are moved together. Each stack has an embedded insert that
908 forms a hole for the vacuum pipe for the beams when the ECal is in its closed
909 position. Motion of the stacks away from the beam is required to allow a
910 close packing of ECal and HCal and the servicing of ECal. The stacks are
911 also moved away from the beam to limit the formation of radiation-induced
912 color-trapping centers in the glass during heavy ion running.

913 The inner and outer calorimeters have cells with a size mismatch. The
914 central $30\text{-row} \times 30\text{-column}$ hole of the outer calorimeter is too small using
915 nominal transverse cell sizes for the $32\text{row} \times 32\text{-column}$ matrix of $(38 \text{ mm})^2 \times$
916 45 cm lead-glass bars. The size mismatch can be accommodated by strategic
917 placement of 0.79 mm thick FR4 shim pieces that are readily available from
918 commercial suppliers. We have used this method in earlier calorimeters.
919 The shim can be included in the GEANT model of the calorimeter. Exact
920 accounting of shim pieces awaits transverse cell size measurements for the $(40$
921 $\text{mm})^2 \times 40 \text{ cm}$ lead-glass bars. Nominal sizes are used for Fig. 26, and 0.79
922 mm thick shims are assumed to be placed between every row. The mismatch
923 in the columns is accounted for by a gap at $x = 0$. Such a gap is present in
924 any case to allow for the enclosure.

925 The modular design of $A_N\text{DY}$ (model=2 in the LOI) was our initial plan
926 in the Letter of Intent [2]. There was a 3-year plan for $A_N\text{DY}$ specified in
927 that document. The first year was the run-11 configuration, with its mod-
928 ular left/right symmetric HCal. The second year, covered by this proposal,
929 was to have left/right symmetric ECal constructed in front of the existing
930 HCal. We were forced to abandon this option to ensure sufficient statis-

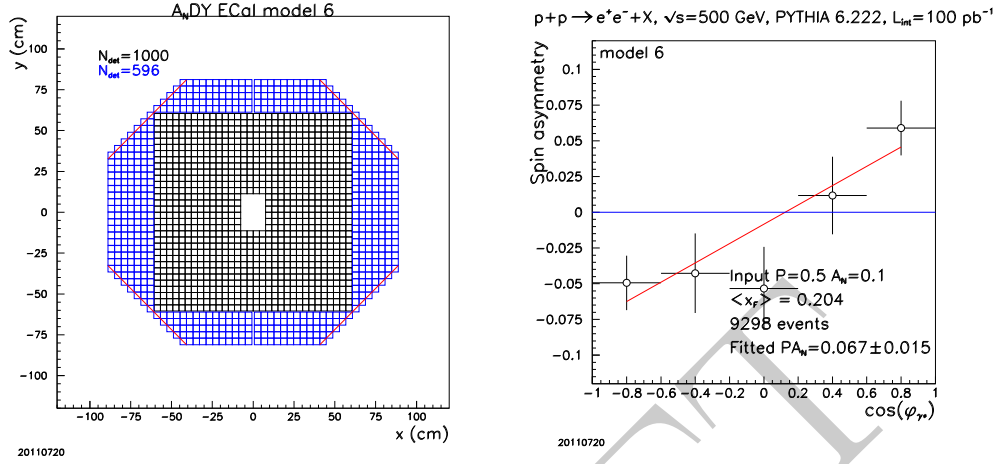


Figure 26: (left) Proposed configuration of the A_N DY ECal as seen from IP2. The inner calorimeter is a 32-row \times 32-column matrix of $(38 \text{ mm})^2 \times 45 \text{ cm}$ lead-glass detectors with a central 6-row \times 4-column hole for the vacuum pipe for the beams. The calorimeter is surrounded by $(40 \text{ mm})^2 \times 40 \text{ cm}$ lead-glass detectors from BigCal, with corners clipped to match the channel count. The full symmetry is broken by the $|y|$ extent of the outer calorimeter being smaller than the $|x|$ extent, as set by mechanical constraints. This outer calorimeter has a central 30-row \times 30-column hole into which the inner calorimeter fits. The row/column matching of the two detector types is identical to that used in BigCal. This view shows the calorimeter in its closed position. The calorimeter separates into two halves that can be translated to larger $|x|$. (right) Simulated sensitivity to the spin asymmetry for Drell-Yan production assuming the acceptance of ECal as given by the ECal configuration and 100 pb^{-1} of polarized proton collisions at $\sqrt{s}=500 \text{ GeV}$. This simulation assumes 100 pb^{-1} of integrated luminosity for polarized proton collisions at $\sqrt{s} = 500 \text{ GeV}$ with 50% beam polarization. It further assumes an $A_N=0.1$ for DY production. The product $\epsilon = P_{beam} A_N$ is what determines the slope of the linear dependence on $\cos\phi_{\gamma^*}$. Kinematic restrictions on the virtual photon are $M > 4 \text{ GeV}/c^2$, $p_T < 2 \text{ GeV}/c$ and $x_F < 0.3$ as required to test the theoretical prediction of a sign change for the analyzing power for DY production relative to that measured for semi-inclusive deep inelastic scattering.

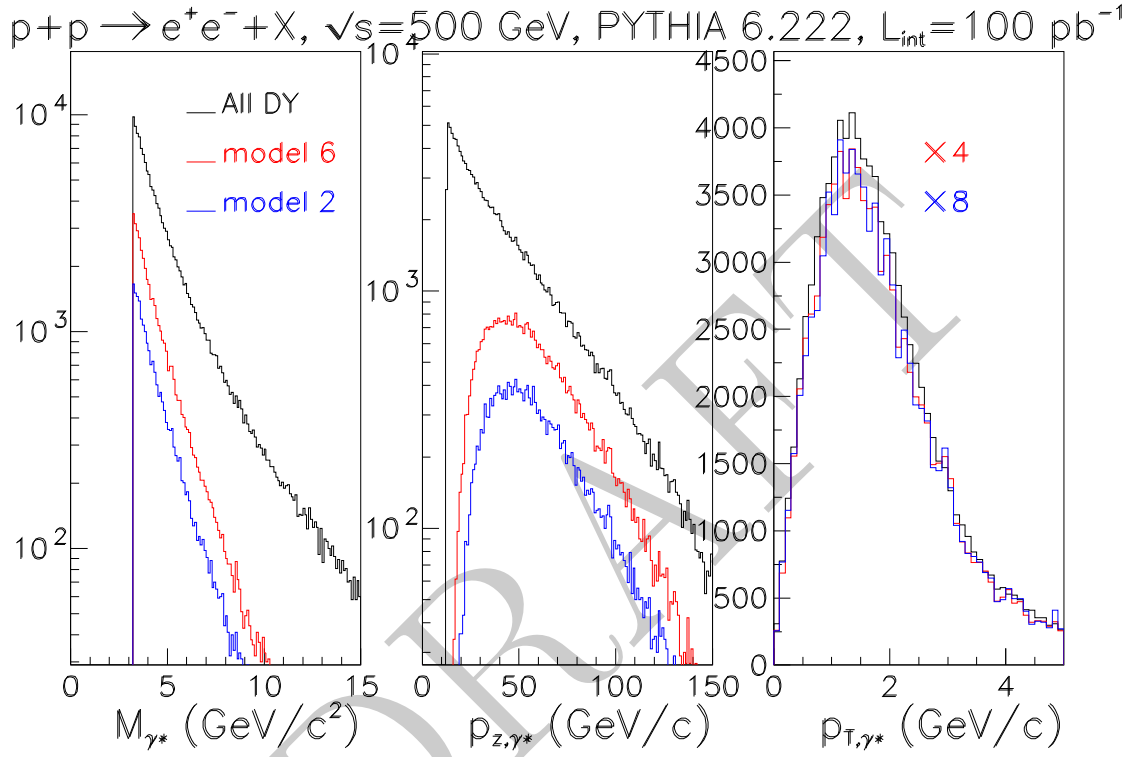


Figure 27: Distributions of virtual photon kinematic variables subject to the acceptance of different models of the A_N DY apparatus. Model=6 includes the A_N DY acceptance described in this document. Model=2 is the original modular design. The increased acceptance of model=6 offsets the difference between initial expectations for integrated luminosity and projections based on the run-11 experience.

tical precision for the sign change measurement. The basic requirements of ECal for A_N DY consist of overall acceptance, hermiticity, resolution, gain adjustability, gain stability, and serviceability. Acceptance is discussed below. Hermiticity ensures minimal transverse shower leakage. This requirement is met by close-packing the glass in the calorimeter. The inner calorimeter will not require any shim pieces. Prior experience indicates that the transverse dimensions of the Protvino glass have a narrow distribution. Close-packing can be facilitated by suitable selection of the detectors in stacking ECal. The resolution requirement can be intrinsically met by lead glass (typically providing 1000 photoelectrons in response to 1 GeV electron-energy equivalent of incident energy), so long as individual detectors share a common ground only at the HV power supply and at the readout electronics, and the noise levels of the readout electronics are small. Gain adjustability and stability sets requirements on the HV system, discussed in section 3.4. The gain stability will be measured by a monitoring system described in section 3.5.

The acceptance requirement is set by estimates of DY event totals assuming delivered luminosity of $10 \text{ pb}^{-1}/\text{week}$ and 10 week runs. The objective is to measure the analyzing power for DY production to sufficient precision to test the theoretical prediction of a sign change, relative to SIDIS. This test requires restricting $x_{F,\gamma^*} < 0.3$ to match the kinematics of SIDIS. Acceptance simulations demonstrate that the acceptance is adequate to test the sign change with the optimized design. Comparisons are made between the optimal acceptance and the original modular concept (model=2)

3.1 ECal major mechanical design

Each ECal stack will be supported on a strong base plate, attached to bearings that are mounted on Thomson rails to provide motion of each half perpendicular to the beams. ECal will be elevated to beam height using vertical lift tables obtained from AGS-E949. Finite element analysis (FEA) calculations have been completed to establish that deflections of the vertical lift tables and the base plates are minimal. A technique that has worked in the past to make an azimuthally complete calorimetric coverage is to construct an insert that gets included with the lead-glass detector stack. This insert is designed to handle the load of the glass stacked on top of it. FEA calculations have established that sufficient strength for the insert, including a safety margin

A light-tight enclosure will be built around each calorimeter half on a

$p+p \rightarrow e^+e^-+X$, $\sqrt{s}=500$ GeV, PYTHIA 6.222, $L_{int}=100$ pb $^{-1}$

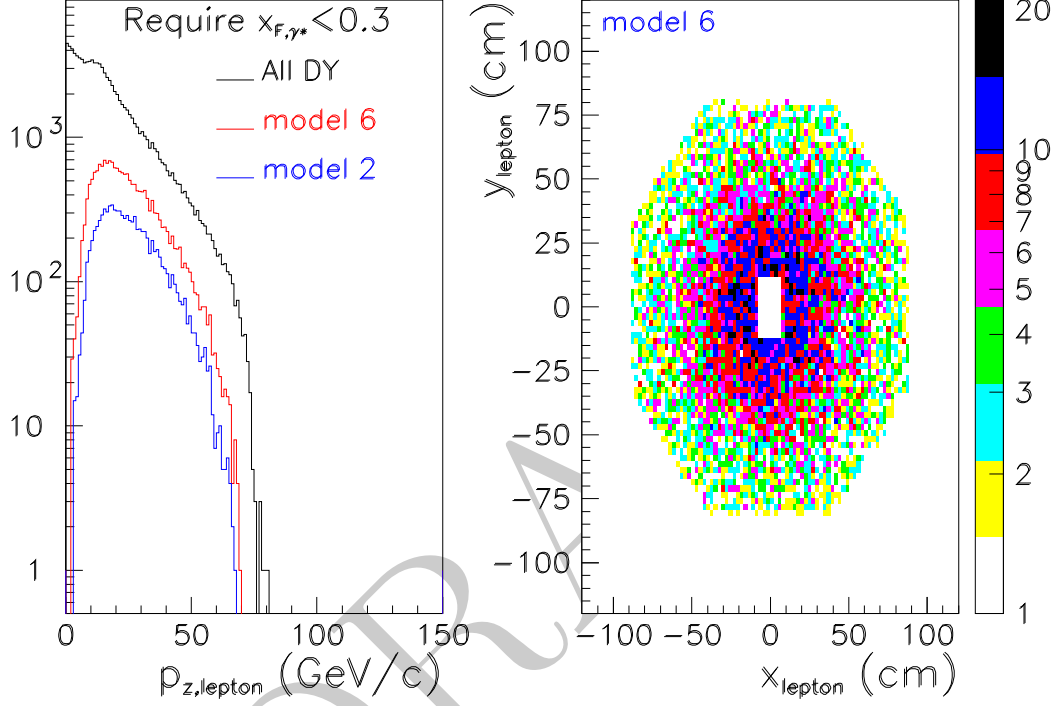


Figure 28: Electron and positron energy distribution (left) and impact position correlation (right), subjected to requirements on virtual photon $M > 4$ GeV/ c^2 , $p_T < 2$ GeV/c and $x_F < 0.3$. The p_T cut on the virtual photon assures sensitivity to spin- and transverse-momentum dependent distribution functions. Model=6 includes the A_N DY acceptance described in this document. Model=2 is the original modular design. The increased acceptance of model=6 offsets the difference between initial expectations for integrated luminosity and projections based on the run-11 experience.

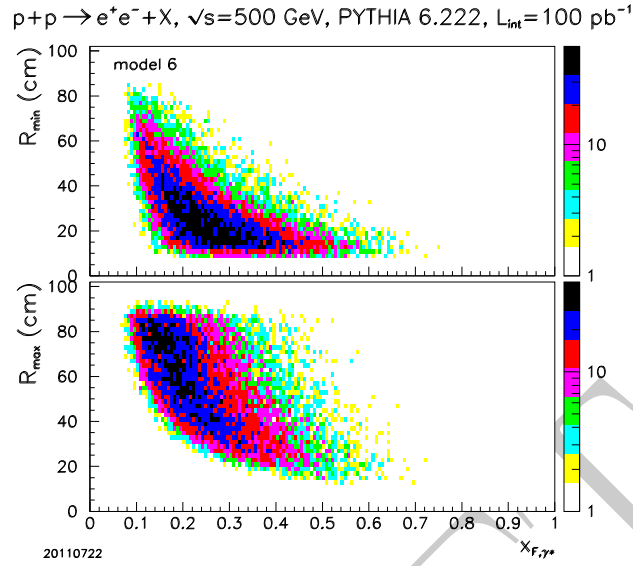


Figure 29: The distance R of the lepton from the beam when projected to ECal is correlated with the Feynman x . The challenge for the acceptance is to ensure large enough areal coverage and still provide acceptance close to the beams.

framework attached to the base plate. The enclosure will include patch panels to feed through the current pulses from each photomultiplier tube and control signals for the proposed Cockcroft-Walton bases. Feedthroughs will be on the large $|x|$ side of the enclosures so that the separation between ECal and HCal can be minimized, as required to ensure robust matching of ECal and HCal clusters. The mean transverse momentum of particles produced in hadronic showers will negatively impact the cluster matching if the z separation is too large. As discussed in Section 3.5, a rack of ECal readout electronics will be in proximity to each ECal half to minimize the cost for cabling.

3.2 ECal construction plan

Experience with the small-scale ECal implemented for RHIC run 11 indicates that the annealing done after the high-luminosity experiment at JLab was quite effective. As per agreement with our JLab colleagues, a fixed light source will be used to characterize the transparency of each lead-glass bar. In these measurements, the basic operation of FEU-84 coupled to that detector

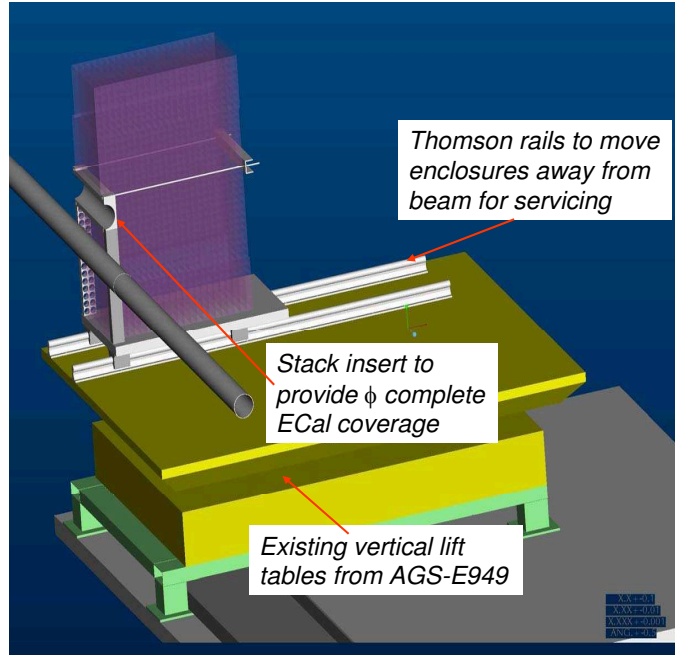


Figure 30: Engineering design of one A_N DY ECal half showing the platform to elevate ECal to beam height, the stack insert to enable the two halves to close around the beam pipe and the Thomson rails used for translating the calorimeter halves perpendicular to the beam pipe.

983 will get verified and the effective gain of the FEU-84+lead-glass bar will get
984 measured. In addition, six measurements of transverse glass sizes will get
985 made so as to optimize the hermiticity of the glass stack.

986 The FEU-84 will get optically coupled to the lead glass in one of two
987 methods. The $(40\text{ mm})^2$ have substantial vanadium blocks glued to the
988 glass to serve as phototube holders. Threaded rods will be screwed into
989 pre-existing tapped holes (insert picture). A phototube holder prestamped
990 with the pin pattern of the FEU-84 is then captured by two threaded rods
991 and pushes the phototube against the glass. A thin film of optical grease
992 between the FEU-84 entrance window and the lead glass will then complete
993 the optical coupling. The $(38\text{ mm})^2 \times 45\text{ cm}$ lead-glass bars will employ a
994 phototube holder matrix constructed for BigCal at JLab. This matrix is a
995 substantial aluminum plate with holes precisely drilled for a matrix of FEU-
996 84. Each hole is accompanied by threaded holes for threaded rods and result
997 in a comparable method of achieving the optical coupling as for the other
998 detectors. The fixed matrix of holes for the FEU-84 will require a matching
999 of the physical dimensions of the glass in the stack, so that accumulated
1000 deviations of the actual glass dimensions does not lead to a mismatch between
1001 the precision hole spacing of the phototube holder over the 32 columns that
1002 the phototube holder covers. One difference of the A_N DY configuration over
1003 BigCal is that the central rows of the detector have an insert. We plan
1004 on building special phototube holders for these central rows that double as
1005 structural support for the insert. Finally, the bottom rows of $(40\text{ mm})^2 \times 40$
1006 cm lead-glass bars will have analogous phototube holders fabricated, so as to
1007 support the existing phototube holders for the upper rows.

1008 After shipment to BNL, the glass was staged in the climate controlled
1009 1002B building. The plan is to do the transparency measurements, photo-
1010 tube coupling, effective gain measurements and lead-glass size measurements
1011 after moving portions of BigCal into 1002D. From there, the plan is to im-
1012 mediately begin stacking ECal. It is estimated that the individual detector
1013 measurements and stacking can be completed in ~ 50 days, by sustaining a
1014 throughput of 32 cells per day.

1015 **3.3 ECal high voltage (HV) system**

1016 The JLab BigCal detector provides 660 cells (Yerevan) whose bases use stan-
1017 dard LeCroy High Voltage (HV) distributed on SHV cables for each FEU-84
1018 PMT channel and 1100 cells (Protvino) that use LeCroy HV augmented

by Hewlett-Packard (HP) low-voltage (LV) booster supplies for the last 4 dynodes. The high voltage connections to the existing Yerevan bases use hypertronix connectors. This system requires a total of 7 LeCroy 1440 systems and 4 HP high-current LV supplies. While we can borrow the LeCroy and HP units (we believe, although we have not yet identified sources), the cost for cables and patch panels to implement this at RHIC will be nearly the same as the cost for replacing all the bases with Cockcroft-Walton (CW) bases with computer-control. If we have to purchase LeCroy mainframes the cost will exceed the CW costs. The CW solution would be used at JLab when we return the detector, representing an overall savings for DOE.

These considerations have resulted in our plan to construct Cockcroft-Walton bases for the FEU-84.

The requirements on the CW system include:

- provide 1300-1800 V for 1700 FEU-84 PMTs having 12 dynodes each
- remote setting of HV values with 2 V sensitivity to match PMT gains in situ
- remote readout of set voltage value
- provide HV with better than 2
- less than 100 mW power consumption per base
- provide gain stability over operating range of 1 to 10^5 Hz or up to $100\mu\text{A}$
- switching noise of < 0.25 pC in 100 ns gate
- fit into the ECal enclosure with room to allow cable access for HV and control
- total length for each base of < 15 cm
- provide 4' pigtailed from each base for signal and control lines
- allow signal connection to BNC patch panel

Many versions of CW bases have been designed for small photomultiplier tubes like the FEU84, including solutions at JLab [26] and at BNL. We have

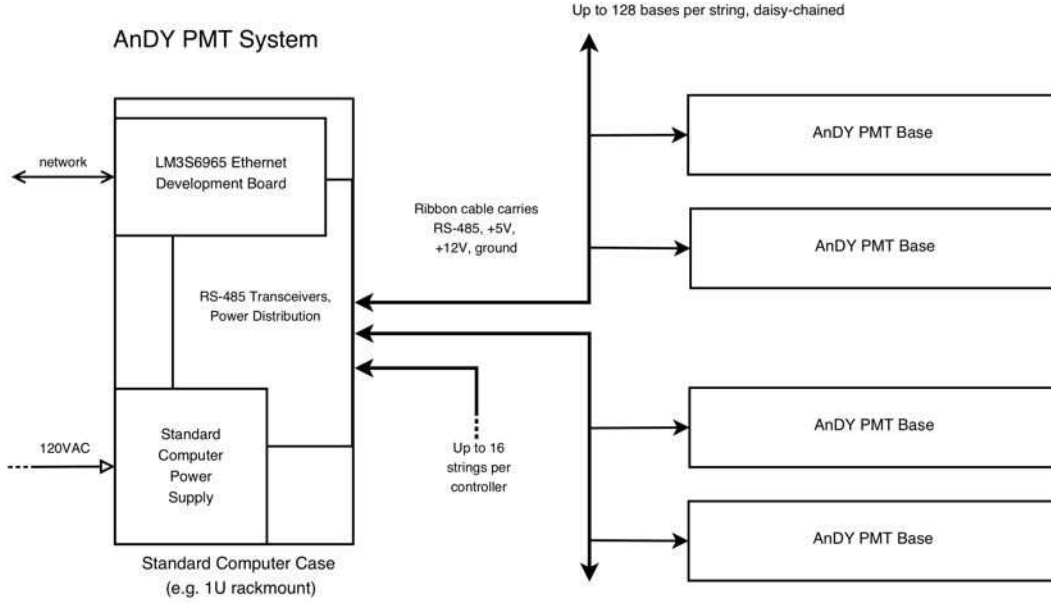


Figure 31: ECal PMT system diagram

borrowed heavily from these but intend to implement our solution with currently available much simplified computer control of the driving voltages. We do not need any elaborate electronics or control to check variation in PMT gain because our detectors all have individual LED fibers to provide continuous monitoring of PMT gain. We have used this LED-flasher system designed and fabricated by UCB/SSI for many years and have already produced the basic elements of the system.

Input power will come from a single standard computer power supply, 5V for electronics and 12 V to drive the diode pump. The ECal PMT HV system is shown schematically in Fig.31.

For ECal, a typical cell produces ~ 500 kHz of 4 pC signals or 2A in minimum bias operation. A single pe at gain of 2×10^5 typical for an FEU84 operated at 1500 V produces 0.03 pC. Our calorimeter cells produce typically 1000 photoelectrons for 1 GeV of electron-equivalent energy loss, or 30pC/GeV. Operating at 1500V then our typical base will draw 3mW average power. A single bin can thus provide power for all of our PMTs.

Each base will accept input power and diode pump voltage from a 32 channel control board. The pump voltage will be controlled for each channel on the control board using a simple DAC. Communication with this board

will be over RS-485 from a dedicated computer. We expect to produce 80 of these boards, using 54 for the ECal, 11 for PS, and 10 for HCal, although only the ECal will use our CW bases. The PS and HCal use XP2972 PMTs with Nunnemaker bases and will use our control board only for controlling the driving voltage.

The CW base has been designed by Mike Ng at UCB/SSL. He will be responsible for both the CW base and the control system for it. We will produce 1800 of the bases, including sockets, 12-dynode HV supply, and 4' signal and control pigtails using commercial fabrication houses once we have shown that our prototype is acceptable.

3.4 ECal readout and triggering

The forward calorimeter projects we have been involved in over the past ten years at RHIC have employed a variety of readout electronics. First prototypes have used commercial gated analog-to-digital converters (ADC) and CAMAC readout (LeCroy 4300B) that remain readily available. These ADC cannot be gated at the RHIC clock frequency (9.38 MHz) due to the slow speed of the ADC. They require a fast trigger to open the ADC gates and delay lines for the individual phototube current pulses. This is not a good solution for A_N DY. The first flash ADC solution used for forward calorimeter projects was a 16-channel 8-bit ADC 9U VME board developed at Lawrence Berkeley Laboratory for fast triggering detectors at STAR. This is also not a good solution for A_N DY, since the 8-bit granularity is too coarse and this flash ADC board was notorious for high-frequency electronic noise. The issues with this flash ADC board were addressed in the QT design created at Berkeley/Space Sciences Laboratory. The QT board (see Fig.44) has proven to be a nearly ideal solution for calorimetry. It is a 32-channel 70 MHz flash ADC with 0.25 pC/count granularity housed on a 9U VME board. The noise levels are < 0.25 pC. Each 9U board consists of a mother board and four 8-channel daughter board. Signals are input to the device through 8-channel Positronix cable assemblies. Each daughter board has a field-programmable gate array (FPGA) to provide the first level of a digital trigger. The mother board also has an on-board FPGA to drive 32-bits of data on the backplane of the VME crate.

The proposed layout for ECal will require 50 total QT boards housed in four 9U VME crates (see Fig.48). Each left/right half of ECal will be readout by a rack containing two 9U VME crates that hold in total 25 QT boards

within the two crates of each rack. A total of 25 twisted-pair ribbon cables will carry data from each rack from the on-board FPGA to the trigger crate for trigger decisions by the DSM. Separately, cat-5 cabling and optical fiber will be required for communication with the MVME 2306 processor in each crate. A twisted-pair clock line is required for each crate. Clock signals are distributed to the individual QT boards on the backplane of the 9U VME crate. 208 VAC lines will provide the estimated 20 kW of power to each rack. A chiller/radiator system is needed for each rack to provide cooling to the electronics. Care must be taken with this system, since the IP2 area is not climate controlled. Elevated dewpoints when the area heaters are off can cause condensation with potentially disastrous impact on the electronics.

A pair trigger on ECal is critical so that the 4kHz bandwidth of the DAQ system is not saturated by the 500 kHz interaction rate from colliding beams. The factor >100 reduction from interaction rate to DAQ rate can be easily accomplished with a pair trigger. The essential hardware for the trigger is a tree structure of field-programmable gate arrays, see Fig.47, although the full flexibility is limited by the availability of hardware. Each DSM board (Fig.45) has 128 bit input and either 16 or 32 bit output. A total of four DSM boards will allow 8 bits from each QT board to make a pair trigger. A natural mapping of the four DSM boards is then quadrants of ECal. Dilepton pairs of interest will have an average separation at ECal of 100 cm with 99% of the events where the electron and positron are in different quadrants. It will be necessary to check that the distance from the quadrant boundary is sufficiently large to ensure that the pair trigger is not flooded by lower mass events.

3.5 ECal monitoring

Online monitoring of all detector elements is done during all runs. The individual detector spectra can identify issues and monitoring software allows for rapid response to such issues. A precise determination of each detectors relative gain is provided by a light-emitting diode distribution system. A method that has worked very well for relative monitoring in the past is to employ a fast current pulse generator that drives a blue Nichia B500S light emitting diode (LED). Each LED is coupled to a light mixer and then to a bundle of 30 1-mm optical fibers. The optical fibers are then glued into a plexiglass panel that has a matrix of holes to match the matrix of the glass. Since this monitoring system is intended for relative gains, requirements on

light uniformity and alignment of the fiber panel relative to the glass stack are not severe. Instead, the primary requirement is mechanical stability, so that once in place the light injection into the glass changes minimally with time

We propose a similar monitoring system for the A_N DY ECal. This will require 50 current pulser boards, LED and light mixers. Each fiber mounting panel will inject light into a quadrant of ECal through 399 optical fibers bundled at one end and glued into the panel at the other end. Electronics to drive the TTL trigger for the current pulser boards is also required, as is a power distribution for the current pulse. The trigger to the current pulser boards needs to be synchronized with the clock to the QT boards, to assure that the current pulses from each PMT of ECal are phase locked to the QT gate. An earlier design for this trigger board exists and would need to be duplicated for the A_N DY ECal monitoring system. The trigger board is built around a commercially available FPGA board, packaged with standard TCP/IP communications. The FPGA distributes a prescaled clock pulse through buffer chips to the individual current pulser boards.

4 Hadron calorimeter (HCal)

A primary means of discriminating electrons from hadrons in A_N DY is the matching of clustered energy deposition in ECal with clustered energy deposition in HCal, positioned immediately behind ECal. There is a $\sim 50\%$ probability that a hadron incident on ECal will begin its shower, as described in detail in Appendix 1. To facilitate cluster matching, it is a requirement that the transverse cluster size in HCal is kept as small as possible. This requirement is met by minimizing the longitudinal separation of ECal and HCal. The optimization of ECal to provide adequate acceptance for electron+positron pairs from Drell-Yan production necessitates an optimization of HCal, so that background rejection is preserved.

As described in Section 2, two 9-column \times 12-row matrices of HCal detectors, originally built for AGS-E864, were staged at IP2. The optimization of the run-13 HCal requires modifications to HCal already staged at IP2. The basic scale of the optimized HCal is determined by the acceptance of ECal, as shown in Fig. 32.

A resulting $x - y$ view of the proposed HCal is shown in Fig. 33.

The integration of the 80 additional HCal detectors with the existing

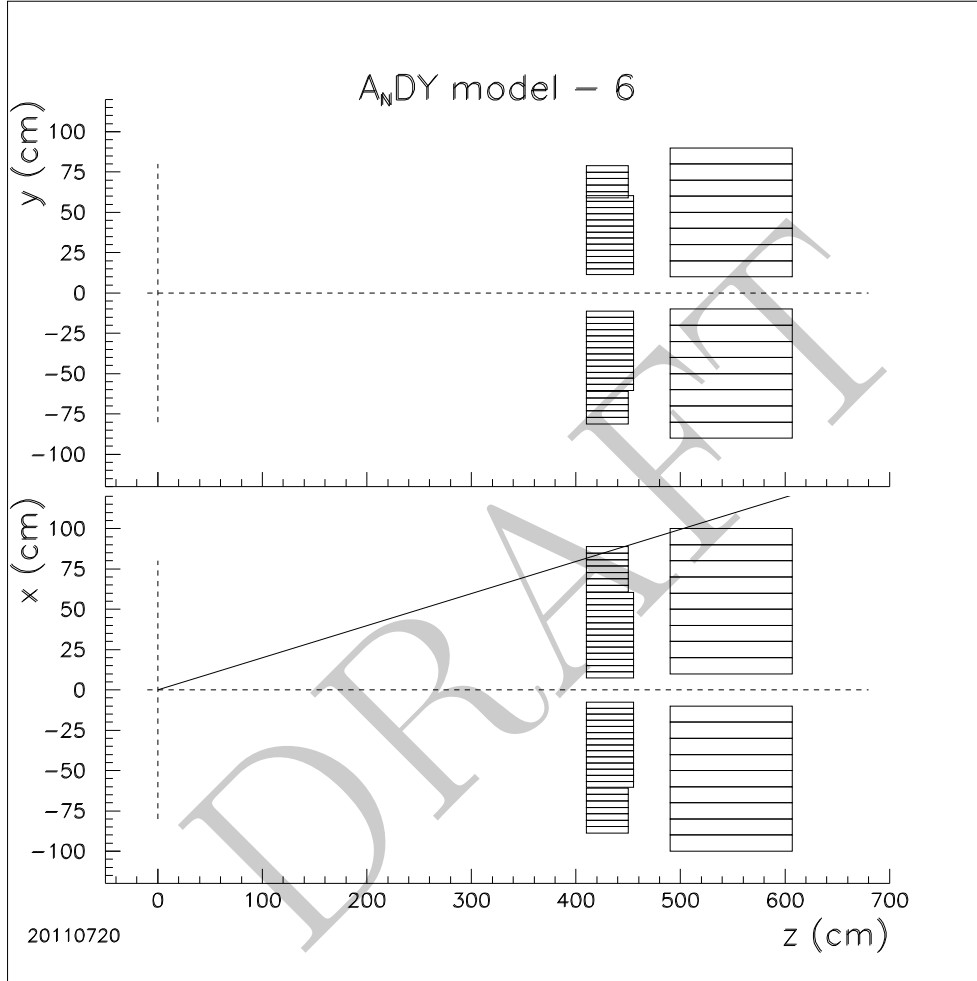


Figure 32: (Top) $y-z$ view at $x = 0$ showing the relationship of the model=6 ECal and HCal, reconfigured from its run-11 implementation. (Bottom) $x-z$ view at $y = 0$ showing the relationship of ECal and Hcal. These schematic views show only the lead-glass of ECal and the active portion of HCal.

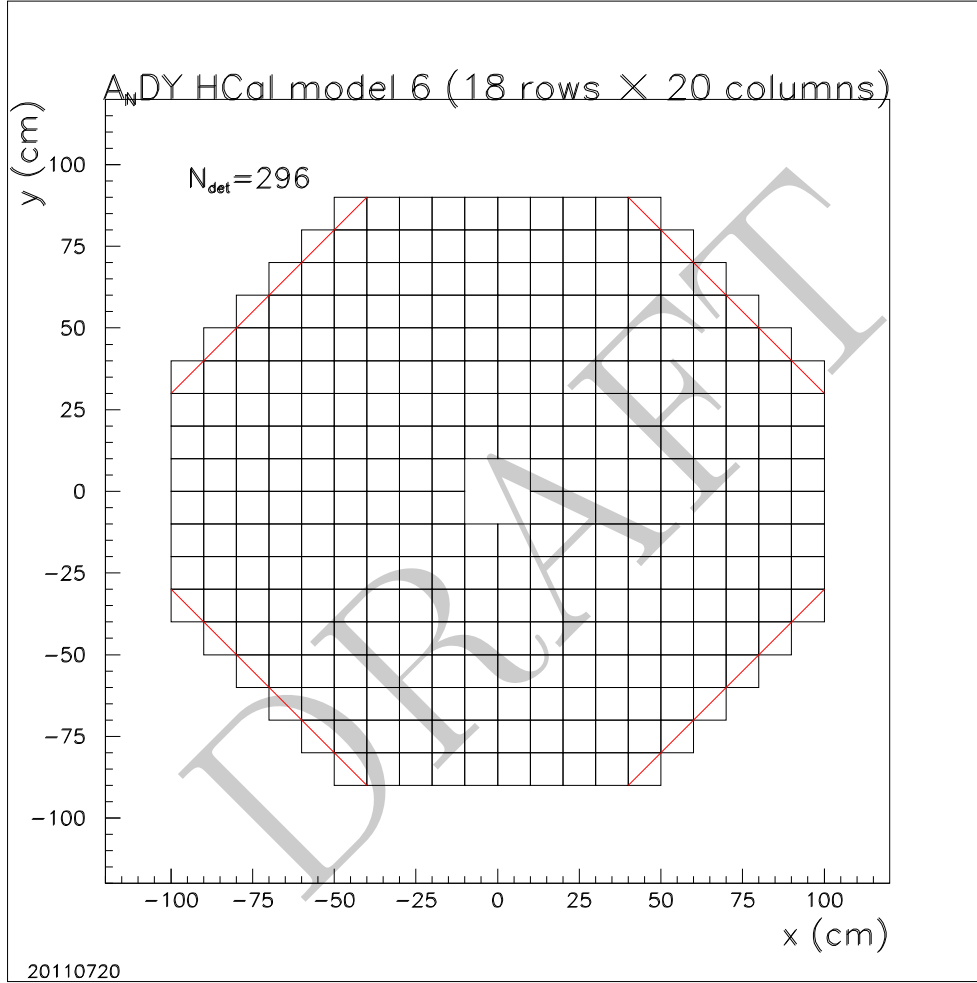


Figure 33: Schematic view of the proposed HCal as seen from the IP.

1174 HCal modules can be done by:

- 1175 • A rigging operation in which each HCal module is lifted off its sup-
1176 port blocks and the support blocks are removed. The blocks are to
1177 be replaced by tracks used for the original design of the HCal mod-
1178 ules. Replacement of the blocks by tracks will lower the existing HCal
1179 modules by 30 cm, corresponding to three rows of detectors.
- 1180 • Six additional rows of detectors will get mounted onto the existing base
1181 plate, resulting in the 18-row HCal schematically shown in Fig. 33.
- 1182 • A 2-column \times 8-row submatrix of HCal detectors will be built between
1183 the two modules.
- 1184 • An aluminum U channel will be supported from the bottom 2-column
1185 \times 8-row submatrix. A 2-column \times 8-row submatrix will be built atop
1186 this U channel support. FEA calculations show that the U channel
1187 provides sufficient strength to support this submatrix above the beam
1188 pipe, including a safety margin.
- 1189 • A 20-column \times 6-row backplate addition will get attached to the pair of
1190 existing 9-column \times 12-row backplates. These backplates are required
1191 for mounting the photomultiplier tubes to the HCal detectors, and also
1192 form the light-tight enclosure for the HCal.

1193 5 Preshower detector (PS)

1194 The essential function of the A_N DY detector systems is to discriminate di-
1195 electron pairs from background. Particle identification is accomplished by
1196 characterizing the longitudinal and transverse shower profiles in each element
1197 of the apparatus. The first step in this characterization is to match clustered
1198 energy in HCal and ECal, and to match these clustered responses to the
1199 preshower response. The preshower refers to a system comprising two arrays
1200 of scintillator detectors (hereafter referred to as $PS_{1,i}$ and $PS_{2,i}$ where i refers
1201 to the segment number of the plane) that view the interactions with minimal
1202 intervening material, a planar converter of known thickness and uniformity
1203 and a final array of scintillator detectors ($PS_{3,i}$) after the converter. The
1204 two planes before the converter satisfy a hermiticity requirement, discussed

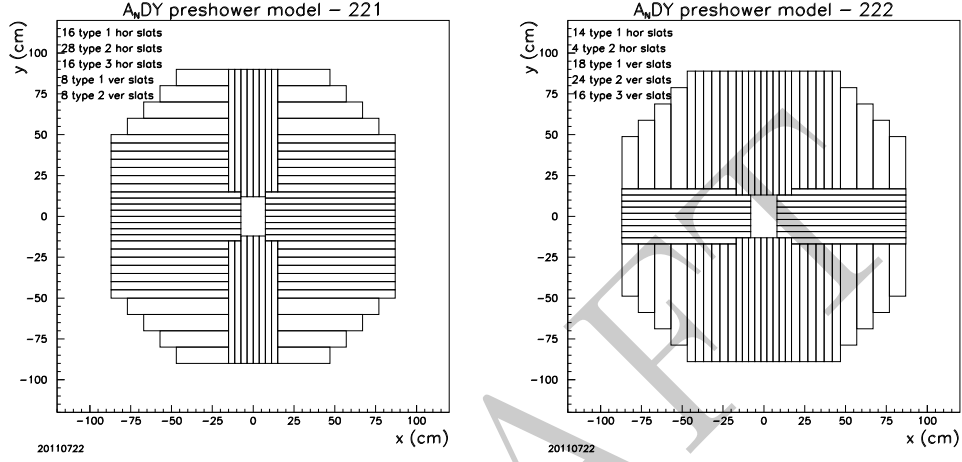


Figure 34: Proposed arrangement of preshower planes 1 (PS_1 , left) and 2 (PS_2 , right). Each plane consists of 76 total detectors each having thickness 0.5 cm of BC-408 (primary component anthracene). There are three different slat widths used to tile the full acceptance of ECal: 3.75 cm; 5.0 cm and 10.0 cm. The PS_1 and PS_2 planes are located at smaller z values than a lead converter plate. Two planes are required for hermiticity. Parallel strips in each plane are staggered by half a strip width to ensure that cracks between adjacent strips are not holes in the combined action of the two planes. Quadrant sections at larger distance from the beam overlap to form pixels between the two planes (see Fig. 35). A third plane of the same layout as PS_1 is located at larger z than a 0.5-cm converter plate. The hermiticity requirement is not as severe for the PS_3 plane since the converter most probably has initiated showers for electrons and positrons.

below. The design assumes that a valid hit in either or both of the first two planes of the preshower will suffice for particle identification.

A critical step in the particle identification is the matching between matched ECal and HCal clusters with the response of the preshower detectors that lie along the line between the clusters and the interaction point. An identified electron or positron will most likely have a single minimum ionizing particle (MIP) in $PS_{1,i}$ ($PS_{2,i}$) and a $PS_{3,i}$ response as given by multiple MIPs, an energetic cluster in ECal and little to no response in HCal. An identified photon will most likely have little to no $PS_{1,i}$ ($PS_{2,i}$) response and $PS_{3,i}$ response as given by multiple MIPs, an energetic cluster in ECal and little to no response in HCal. An identified charged hadron will most likely have a $PS_{1,i}$ ($PS_{2,i}$) and $PS_{3,i}$ response as given by a MIP, an energetic cluster in ECal and an energetic cluster in HCal. An identified neutral hadron will most likely have little to no $PS_{1,i}$ ($PS_{2,i}$) and $PS_{3,i}$ response, sometimes no response in ECal and sometimes an energetic cluster in ECal and an energetic cluster in HCal.

We have considered multiple geometries for the preshower. A robust construction method is a requirement. This sets constraints on geometries that can be realized. It is also a requirement that photostatistics do not limit the resolution of the detector. We look for > 50 photoelectrons for a minimum-ionizing particle (MIP) passing through the detector. Consequently, wave-length shifting fiber readout of pixilated preshower is not an option. The preshower is also required to be as uniform as possible so that robust modeling of it can be completed in simulation codes. The uniformity requirement rules out configurations where photomultiplier tubes are used within the preshower acceptance to, for example, collect light from segments at large pseudorapidity. The preshower, especially the plane(s) before the converter, must be highly efficient to ensure robust γ /electron/hadron discrimination. Inefficiencies are mostly related to gaps between adjacent detectors. Our design first minimizes gaps by the construction methodology and secondly involves two scintillator planes before the converter to allow staggering of gaps. A valid hit in a detector from either plane will work for identifying electrons and positrons from DY production. We do not plan to include the preshower in the trigger, so requirements on gain uniformity are not critical. Hardware level gain uniformity to within a factor of two will suffice. Gain spread larger than that would then face limitations from the 12-bit granularity of the analog-to-digital converters. Software corrections can be done offline to achieve the gain uniformity needed in the analysis, using well

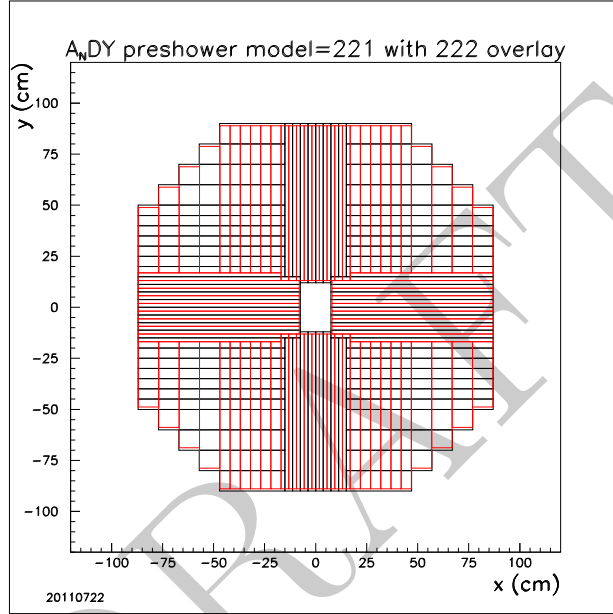


Figure 35: Overlay of the PS_1 and PS_2 scintillator preshower planes. Horizontal (vertical) strips near $y = 0$ ($x = 0$) are offset by one half strip width between the two planes. At larger $|x|$ and $|y|$ values the PS_1 and PS_2 strips are orthogonal to each other, effectively making pixels. It is assumed that good di-electron events will have hits from a single particle in either (or both) the PS_1 and PS_2 planes. This good hit requirement sets the requirement on detector granularity.

defined MIP peaks for calibration. In general, the preshower requirements are best met by a strip solution.

The matching of preshower, ECal and HCal response is complicated by the multiplicity of produced particles. When two particles interact with the same detector element the response of that element is confused. For example, there are two straight lines from the primary vertex to two ECal clusters with measured values (x_i, y_i, E_i) . These two lines may intercept the same preshower detector $PS_{1,j}$. This preshower detector will provide the summed energy deposition of the two particles that produce the ECal clusters. A method of dealing with this complexity is to reject events when critical detector elements have multiple hits. The question then is impact of this rejection method on observing DY di-electrons.

Simulations have addressed the impact of multiple occupancy in the preshower strips on DY detection efficiency. These multiple occupancy estimates eliminate contributions from inner bremsstrahlung by merging closely spaced clusters in ECal, and account for the photon hits from inner bremsstrahlung that are nearby to the electron or positron as a single particle. Approximately 8% of DY events are lost when one or both lepton daughters intercept preshower detectors in both PS_1 and PS_2 that already have energy deposition from other particles. Approximately 89% of DY events will have singly-occupied PS_1 and PS_2 detectors for both the electron and positron thereby adding information to the particle identification.

The geometry proposed for the PS_1 and PS_2 planes of the preshower is shown in Fig. 34. These planes consist of vertical and horizontal slats. The scintillator thickness is assumed to be 0.5 cm of BC-408. A total of 152 strips from the two planes will be read out by 5 32-channel QT boards, leaving 8 spare channels. The detectors are readout by XP-2972 photomultiplier tubes at their ends distant from the beam. Each XP-2972 is powered by a Cockcroft-Walton base. The XP-2972 tubes and CW bases were purchased for AGS/E-864 and 400 sets have been loaned to A_NDY. Tests of these phototubes and bases are underway.

The PS_1 and PS_2 planes of the preshower will be followed by a sealed lead converter plate of thickness 0.5 cm (approximately 1 radiation length). Sealing of the lead will be accomplished by encasing it in adhesive-backed carbon fiber sheets. This encasing will then provide a robust means of mounting the converter in the preshower enclosure. We will work with a vendor to ensure their delivered product matches our flatness criteria and that the material composition will be quantified. Given the softness of lead, even when

1281 hardened by making an alloy with antimony, it is possible to flatten it upon
1282 receipt from a vendor. One radiation length is a tradeoff between initiating
1283 showers for incident electrons, positrons and photons versus not initiating
1284 hadronic showers for incident hadrons.

1285 The sealed lead converter plates will be followed by a third plane of scin-
1286 tillator. The object of this plane is to detect particles from showers initi-
1287 ated in the lead. Finite shower sizes make this plane less prone to cracks
1288 compared to scintillator before the lead. Consequently, we will rely on con-
1289 struction methodology to minimize cracks between PS_3 detector elements.
1290 A glued construction will minimize cracks in any case. The geometry of PS_1
1291 will be chosen for PS_3 (see Fig. 34). In total, the preshower will have 228
1292 strip detectors and require 8 32-channel QT boards for readout. Connec-
1293 tions between the preshower detectors and the QT boards will require 32
1294 Positronix cable assemblies. The CW base has a K-lock connector for the
1295 co-axial signal. Short lengths of RG108 cabling with K-lock connectors at
1296 each end will connect the preshower detectors to a patch panel, built from K-
1297 lock/BNC bulkhead feedthroughs. K-lock/BNC bulkhead feedthroughs are
1298 also required for the H5010 photomultiplier tube assemblies. This will allow
1299 most of the cabling to use the much less expensive BNC connectors. A total
1300 of 8 low-voltage distribution boards will be required to power the 228 CW
1301 bases. Two low-voltage boards from the run-11 implementation will also be
1302 used.

1303 Each distribution board serves 32 PMTs and has 2 - 16 channel control
1304 boards. The design of these boards is very similar to the design of the ECal
1305 CW base controllers described above and shown in Fig.31.

1306 We propose the construction of the preshower planes will be done in
1307 quadrants, since this corresponds to a manageable size. The construction
1308 method will involve wrapping individual strips by aluminized mylar reflectors
1309 and then gluing the wrapped strips to a mylar substrate. Two quadrants
1310 will be mounted inside a preshower enclosure boxes compactly positioned
1311 at smaller z than the ECal enclosure. An enclosure for the preshower will
1312 eliminate the need for wrapping the glued assembly with black tedlar that
1313 quite likely will make the materials in front of the scintillators non-uniform,
1314 thereby violating a requirement for the preshower. The preshower enclosure is
1315 separate from the ECal enclosure to ease the mechanical design requirements
1316 and to allow ECal to be moved away from the beam, thereby enabling HCal
1317 to view collisions through the preshower for at least calibration via $\pi^0 \rightarrow$
1318 $\gamma\gamma$, as per the method used in run 11 (Fig. 22). Mounting the preshower

inside an enclosure will limit, but not preclude, access. The substrate will be painted black to limit the chance that photons from room lights would inject noise into the readout. We do require that noise levels are ~ 0.25 pC (RMS) so that robust modeling of the detector can be done. The plan is to glue photomultiplier tubes to the scintillator strips for strips < 10 cm wide, or onto light guides for strips of 10-cm width. The glue joints would be reinforced by gluing aluminum L-brackets onto the scintillator near the optical couplings of the phototubes.

We estimate that the preshower assembly and testing can be completed in 6 weeks by two people, after the scintillator is delivered by the vendor. Experience from the run-11 A_N DY implementation is that 14 weeks elapsed from when a requisition was initiated to when the scintillator was delivered. Two weeks are estimated for wrapping and gluing strips into quadrant planes. Two weeks are estimated for gluing photomultiplier tubes onto the scintillators. Finally, two weeks are estimated for tests and installation.

6 Vertex detector (Mid η)

An outstanding issue with the run-11 analysis that still needs to be resolved is a discrepancy between the calibration of the timing difference measurement and vertex-z. One method calibrates this by using single-beam backgrounds and the measured z-separation of the Yellow- and Blue-facing BBC. The other method uses the reconstructed z-vertex in the neutral pion mass reconstruction. These two methods disagree by $\sim 50\%$. Vertex-z plays a critical role for background suppression to access DY production through the matching of ECal and HCal clusters that require a line from the primary vertex and the matching to the preshower.


In addition, this detector is useful in identifying the type of interaction that produces the signals in the other detectors.

To assure a robust vertex-z calibration we propose to convert the 20-strip preshower detectors built for the run-11 A_N DY configuration into Mid η detectors. These detectors would be arranged in two planes to form one module. There would be two such modules mirror symmetrically arrayed left and right of the beam to assess systematic errors in the vertex-z measurement. In total, 80 existing XP2972 and Cockcroft-Walton bases would be coupled to both ends of the existing 40 narrow strips from the run-11 preshower. There would be 80 channels of QT readout required for our proposed Mid η

1354 detector and cabling to connect from PMT to QT.

1355 7 Forward Pion Detector (FPDY)

1356 As mentioned in the introduction, theoretical understanding of our prior mea-
1357 surements and other world data is still evolving, although the theory com-
1358 munity remains consistent regarding their prediction of a sign change for the
1359 analyzing power of Drell Yan production relative to measurements of trans-
1360 verse single spin asymmetries for semi-inclusive deep inelastic scattering. The
1361 A_N DY proponents believe it essential to close all possible theoretical loop-
1362 holes. One such loophole is whether a completed A_N DY experiment robustly
1363 understands the sign of the analyzing power measured for DY production.
1364 To close this loophole, we propose to concurrently measure the sign of the
1365 analyzing power for neutral pion production at x_F values that overlap both
1366 the DY production measurement we will do and prior neutral pion analyz-
1367 ing power measurements we have completed [6]. That measurement will be
1368 done by a forward pion detector (FPD) identical to the run-11 A_N DY ECal
1369 implementation. The only difference is that we will mirror the left/right sym-
1370 metric devices about the interaction point. This mirroring operation allows
1371 the run-13 A_N DY FPD to be positioned as far from the interaction point as
1372 possible (8.5m), thereby enabling a robust reconstruction of the neutral pion
1373 so as to measure its analyzing power to $x_F \sim 0.4$ (100 GeV neutral pions)
1374 without solely relying on identifying them through a deconvolution of highly
1375 overlapping showers in the lead glass.

1376 The run-13 A_N DY FPD enclosures include a shower maximum detector
1377 that allows robust separation of single photon clusters from di-photon clusters
1378 produced by neutral pions when the electromagnetic showers from the latter
1379 are overlapping within lead glass detectors, based on prior experience [5]. A
1380 picture of one such A_N DY FPD module is shown in Fig. 36. All equipment
1381 used in the run-11 A_N DY ECal implementation will be reused for the run-13
1382 A_N DY FPD. Unlike the transverse spin DY measurement, this really is an
1383 engineering measurement given our prior experience 

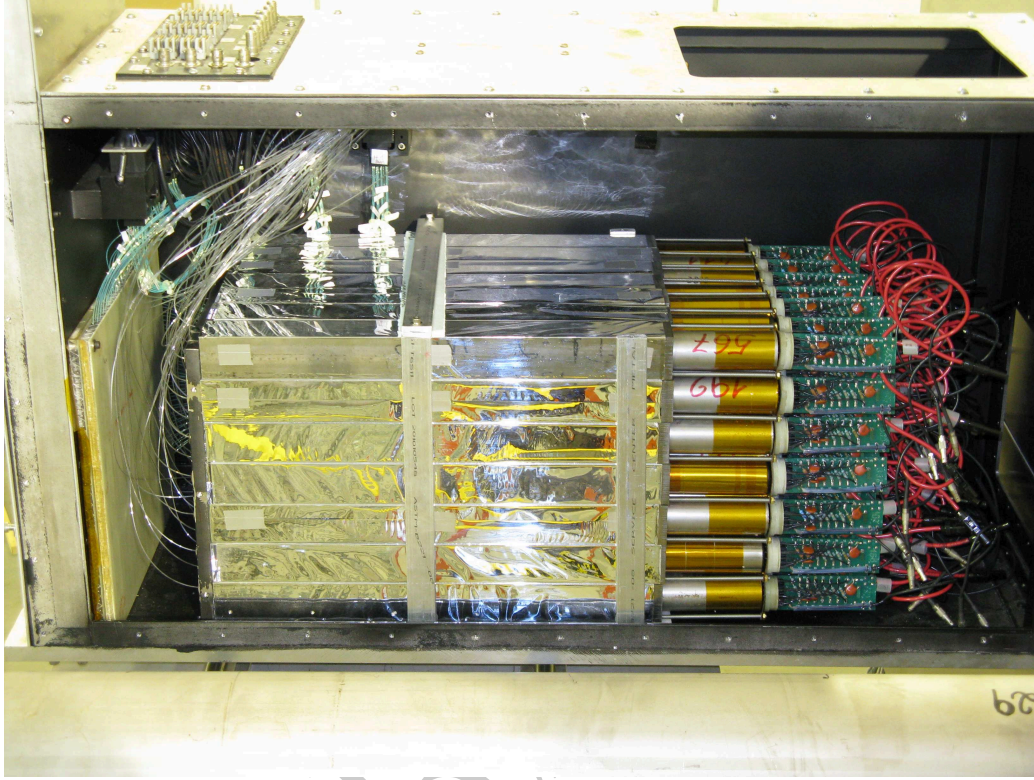


Figure 36: A picture of the run-11 A_N DY ECal before completion of its construction. This will become one module of the run-13 A_N DY FPD. The aluminum box at the far left houses a 7-element lead-glass preshower detector. Next to this box, within the primary calorimeter enclosure, are two planes of a scintillator-strip shower maximum detector. The wave-length shifting fibers are seen as green in this picture because the ambient light contains blue and ultraviolet wavelength photons. The next item seen are clear optical fibers that distribute light-emitting diode flashes to the lead glass for monitoring. A 7×7 stack of $40\text{mm}^2 \times 40\text{cm}$ lead glass detectors is the primary element of the FPD. Its operation in run-11 produced the invariant mass distributions in Fig. 19. The combined analysis of the lead glass response and di-photon finding based on the shower maximum detector will permit robust neutral pion identification to 100 GeV, where the electromagnetic showers in the lead glass are highly overlapping and the matrix response alone cannot be used for robust neutral pion identification.

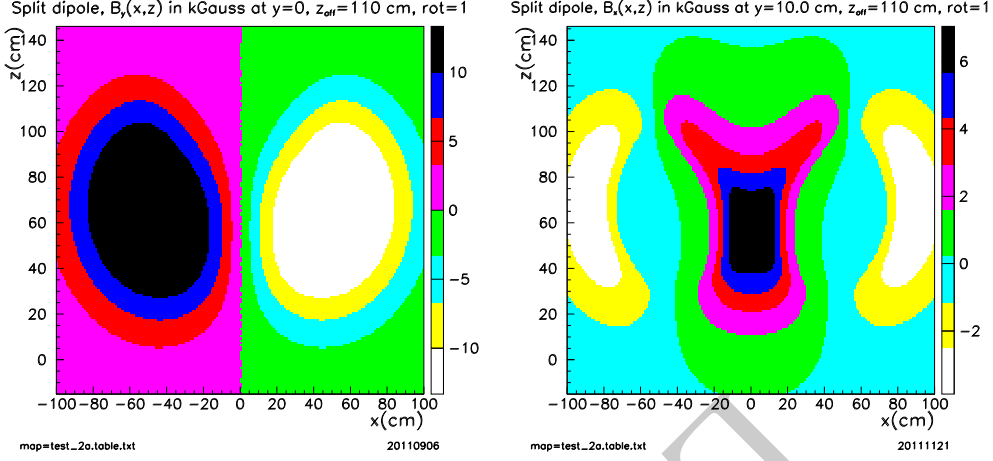


Figure 37: Calculated field map for the modified split-dipole magnet. The modification increases the gap between the poles from 15.7 to 31.4 cm.

1384 8 Gas Electron Multiplier Tracking (TRK)

1385 9 Run-14 magnet implementation

1386 Our initial plan [2] was to install a split-dipole magnet, designed by the
 1387 PHOBOS collaboration [15], and tracking chambers as the means of deter-
 1388 mining the sign of charged particles produced in the collision. Given the
 1389 optimization to the ECal and HCal acceptance, the gap between the poletips
 1390 for the split-dipole must get opened from 15.7 cm to 31.4 cm. As described
 1391 in Sec. 8, there is an exciting development to instrument A_NDY with high
 1392 resolution GEM tracking stations. Consequently, the section of the PAC
 1393 proposal related to the magnet has been superceded by this section.

1394 A new magnetic field calculation with this larger gap has been completed.
 1395 The central field strength decreases as expected and there is a small increase
 1396 in the fringe field. Figure 37 shows a map of $B_y(x, z)$ at $y = 0$ and $B_x(x, z)$
 1397 at $y = 10$ cm from a calculation with the poletip gap at 31.4 cm. As noted
 1398 below, charged particles produced by collisions that are within the ECal
 1399 acceptance are analyzed by both the x, y components of the magnetic field
 1400 from the split dipole.

1401 Tracking studies through the calculated field map have been completed
 1402 similar to those done for the PAC proposal. These studies establish the $x - y$
 1403 loci of events from DY production at the z location of each of three tracking

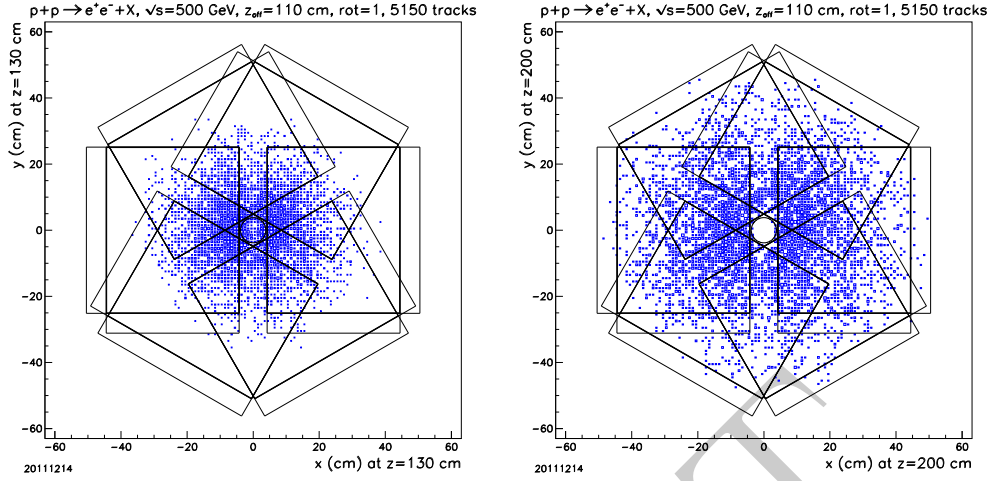


Figure 38: $x - y$ loci of points from DY production at the first two tracking stations. The GEM detectors are represented by the hexagonal tiling of a basic GEM module. Each GEM module is assumed to have two-sided electronics readout that is parallel to the active area of the detector.

stations. These loci impose the requirements on the transverse acceptance of the tracking stations at each z location. Figs. 38,39 show the $x - y$ loci for the proposed layout with the magnet at $z_{off}=110$ cm, the first tracking station at $z_1=130$ cm, the second tracking station at $z_2=200$ cm and the third tracking station at $z_3=390$ cm. The tracking studies are completed for e^+e^- pairs from Drell-Yan production that are within the acceptance of the ECal. In these studies, the vertex z location for each event is drawn at random from a Gaussian distribution with $\sigma=30$ cm. This diamond size has been projected by the Collider-Accelerator department for run 12 [16], and corresponds to optimizing the 9 MHz RF system used for accelerating the proton beams.

For completeness, the distribution of the z location that particles cross the beam pipe is shown in Fig. 40. This distribution is governed by the diamond size, with little impact from the magnetic field since the deflections due to the magnet are small.

Tracking studies through the split-dipole magnet gap opened to 31.4 cm have established deflections at each tracking station (Fig. 41). Deflection is the distance in the $x - y$ plane at the tracking station from the intercept of a straight line trajectory (zero field) to the intercept of the computed trajectory. Deflection crudely establishes the resolution requirements for each

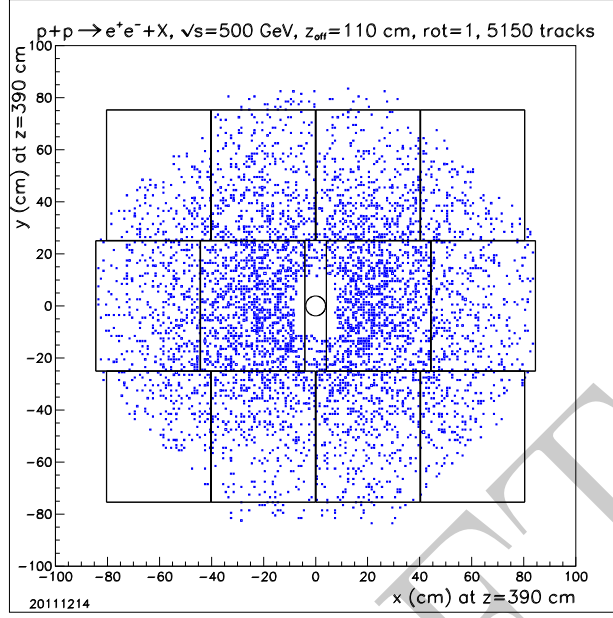


Figure 39: $x - y$ locus of points from DY production at the third tracking station. The GEM detectors are represented by a rectangular tiling of a basic GEM module. Each GEM module is assumed to have electronic readout boards that are perpendicular to the active area of the detector to enable close packing of the modules.

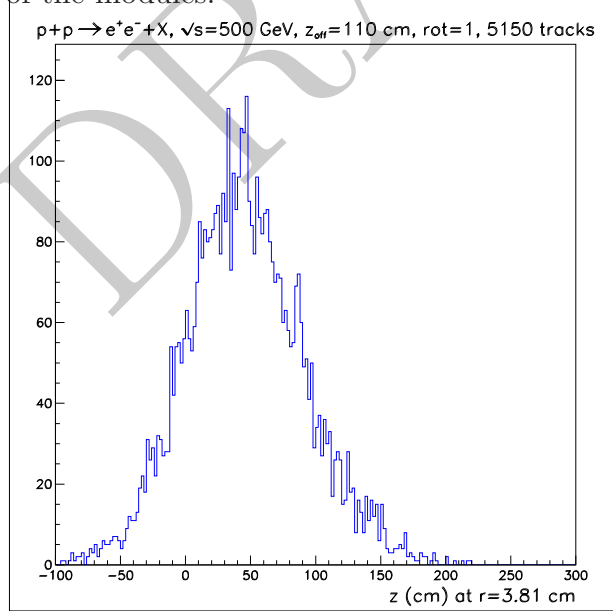


Figure 40: Distribution of the z location where DY daughters cross the beam pipe.

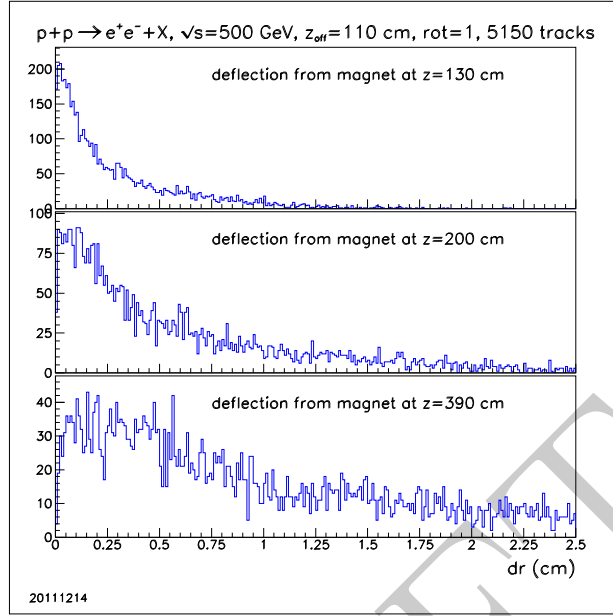


Figure 41: Distribution of deflections at each tracking station. Deflection is the distance in the $x - y$ plane of the tracking station between the field-free intercept and the full trajectory through the split-dipole.

1424 tracking station. The distributions hide the correlations that provide the keys
 1425 to track reconstruction. These correlations are described below. Suffice it to
 1426 say that the deflection distributions are little changed by the increased gap
 1427 between the poles of the split dipole.

1428 The deflection distributions (Fig. 41) have been analyzed to understand
 1429 how well the inclusion of the magnet will determine charge sign. This anal-
 1430 ysis has established that to a very good approximation, the deflections are
 1431 radial, meaning there are both x and y components. This can be understood
 1432 from the magnetic field map (Fig. 37), since the magnet effectively imposes
 1433 an impulse (momentum change) with x, y components $\Delta p_y \sim p_z B_x(y)$ and
 1434 $\Delta p_x \sim p_z B_y(x)$. Furthermore, the impulse is imparted in a spatial region
 1435 between the interaction point and the tracking stations. A radial impulse
 1436 is the intuitive expectation for a focussing element. Using the forward re-
 1437 gion of the split-dipole for momentum analysis results in the magnet acting
 1438 primarily as a focussing element.

1439 Given that the impulse is radial and the impulse is imparted between
 1440 the origin of the particle production and the tracking station, it is then clear
 1441 that charge sign discrimination is equivalent to the displacement between the

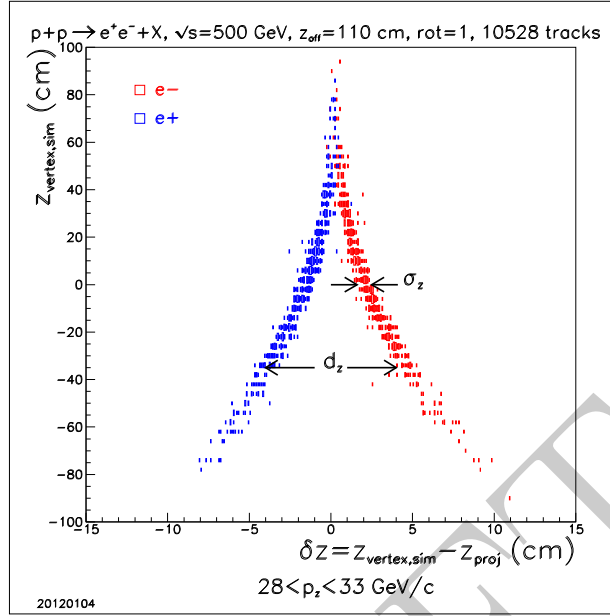


Figure 42: Correlation between the difference between z_{vertex} and the intercept of the fitted track with the beam line and z_{vertex} . Correlations for electrons and positrons are shown separately, to illustrate charge sign discrimination. The loci for each particle type results from the z_{vertex} dependence of radial impulses imparted by the split-dipole magnet.

1442 true z position of the vertex and the point that results from extrapolation
 1443 of a straight-line fit to the track back to the $x = 0, y = 0$ line (*i.e.*, the z
 1444 axis). To illustrate this, the x, y coordinates at each tracking station are de-
 1445 termined from the tracking through the split-dipole field. Afterwards, these
 1446 x, y coordinates are then smeared by Gaussian random numbers representing
 1447 the finite resolution of the GEMs. The resulting x, y points are then fit by
 1448 a linear least squares procedure, and the intercept with the z axis is found.
 1449 The resulting analysis for a bin in longitudinal momentum for the electrons
 1450 and positrons from DY production is shown in Fig. 42.

1451 These tracking studies presently include multiple scattering through the
 1452 beryllium and aluminum sections of the beam pipe, with the relevant material
 1453 determined by the z position where the tracked particle reaches $r=3.81$ cm.
 1454 Studies that include multiple scattering through the GEMs are underway.
 1455 Full GEANT simulations have established that multiple scattering effects
 1456 are dominated by crossing the beam pipe, in any case.

1457 The impulses that give rise to the z displacement in Fig. 42 depend on

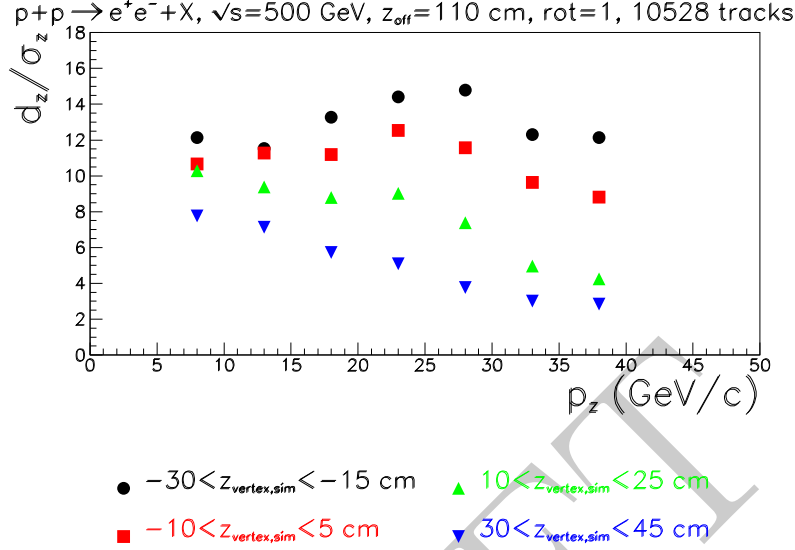


Figure 43: Analysis of Fig. 42 as a function of z_{vertex} and longitudinal momentum. The vertical axis of this plot is the significance of the charge sign discrimination, corresponding to the ratio of the separation for positive and negative charges (d_z) to the width (σ_z) of the distribution at fixed z_{vertex} and fixed momentum. The GEM tracking resolution provides significant charge sign discrimination for all source points over the momentum range spanned by DY production.

longitudinal momentum and on the z_{vertex} for the event. The latter dependence should be expected from the field maps (Fig. 37), since the split dipole has essential zero field ($B_x \sim 0$ and $B_y \sim 0$) near $x = 0$ and $y = 0$. When the event vertex is far from the magnet, the dilepton opening angle allows the electrons and positrons to sample high field regions of the split dipole. When the event vertex is close to the magnet, the electrons and positrons see smaller field strengths thereby resulting in smaller impulses.

A summary of the longitudinal momentum and z_{vertex} dependence of charge sign discrimination is given in Fig. 43. The vertical axis in this plot determines the *significance* of the z separation between positive and negative charges. The significance is a ratio between the physical separation in z (at fixed p_z and the true event origin), scaled by the width of the distribution for electrons (positrons). For source points most distant from the magnet, the beam pipe crossing is most grazing. Consequently, multiple scattering effects

in the beam pipe are the primary limitation in determining the significance (d_z/σ_z). Of course, these distant source points result in charged particles seeing the largest magnetic field strengths, so the significance of the charge sign measurement is large, but is found to be relatively independent of momentum. For source points close to the magnet, the beam pipe crossing is less grazing. Consequently, space point resolution is the primary limitation in determining charge sign. This analysis demonstrates that significance $> 2\sigma_z$ can be achieved for the entire momentum range relevant for Drell Yan production, even for source points close to the magnet, when GEM resolution is available.

The role of the magnet is to go beyond inference to measurement. The importance of charge sign measurement, rather than inference, for DY may be critical to the success of A_N DY. The replacement of inference by measurement is believed to be of critical importance for required ancillary measurements. For example, the heavy quark background to the dilepton production is presently understood within the context of a model (*i.e.*, PYTHIA). Potentially, intrinsic heavy quark components in the proton wave function will have very large impact on heavy quark production in the forward direction. Such contributions are not quantitatively modeled, so have not been assessed at this time. Measurements are required to distinguish DY production from dileptons from intrinsic heavy flavor components. Such measurements are possible in the b quark section by reconstruction of $B \rightarrow J/\psi K$ and by reconstruction of $\Lambda_b \rightarrow J/\psi \Lambda$. Known branching ratios can then be used to determine the background in the dilepton continuum from leptonic decays of heavy flavor. These ancillary measurements will greatly benefit from charge sign discrimination.

10 Triggers and Data Acquisition

The A_N DY trigger and data acquisition systems are based on hardware and software developed for STAR [18]. The STAR trigger electronics was designed to be flexible and has morphed seamlessly into both a portable test system for board and detector testing and into the A_N DY trigger/DAQ system. The trigger uses a hardware decision tree based on QT boards (Fig.44), DSM boards (Fig.45), a Trigger Control Unit (TCU), and an internal control and data network using a copy of the STAR STP system (acronyms defined in Sec.10.1). The DSM and TCU at A_N DY consist of first generation STAR

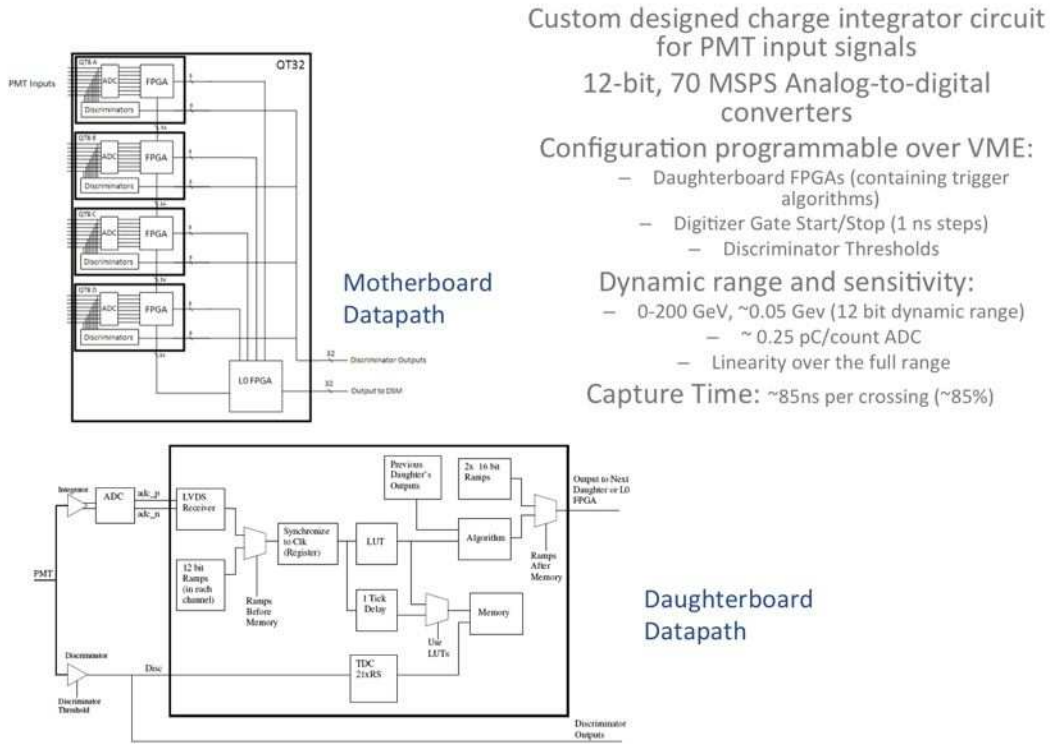


Figure 44: QT board details

hardware that has been superseded at STAR by newer versions that incorporate more powerful logic engines (FPGAs). The trigger needs for A_N DY are far simpler than the complex topological triggers used at STAR primarily because of the tightly focused physics goals of A_N DY.

A_N DY DAQ is a simple token-based readout scheme that is a natural extension of the hardware test routines used in STAR trigger development. Its operation is shown schematically in Fig.46. The A_N DY trigger/DAQ operates at rates exceeding 4 kHz and can take data at any time regardless of instantaneous luminosity. The upgrades from the run 11 to the run 13 configurations involve addition of more than 2k channels and 5 VME crates but involve only a few minor changes in software configuration files.

The A_N DY trigger uses minimum bias conditions based on the zero degree calorimeters and beam-beam counters as well as high-tower and cluster triggers combining various A_N DY FPDY, ECal and HCal modules. The A_N DY

Data Storage and Manipulation (DSM) Boards

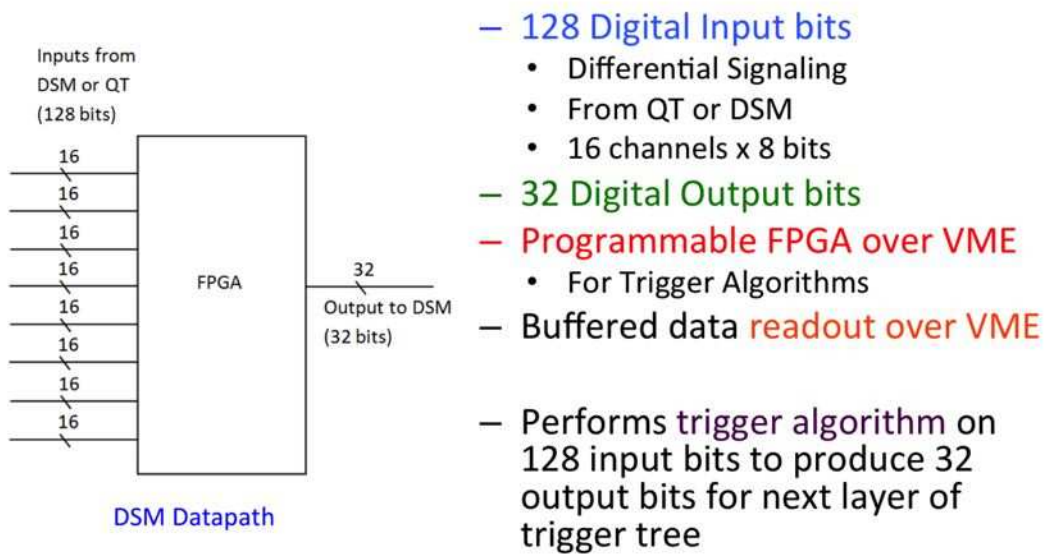


Figure 45: DSM board details

Trigger and DAQ System Overview

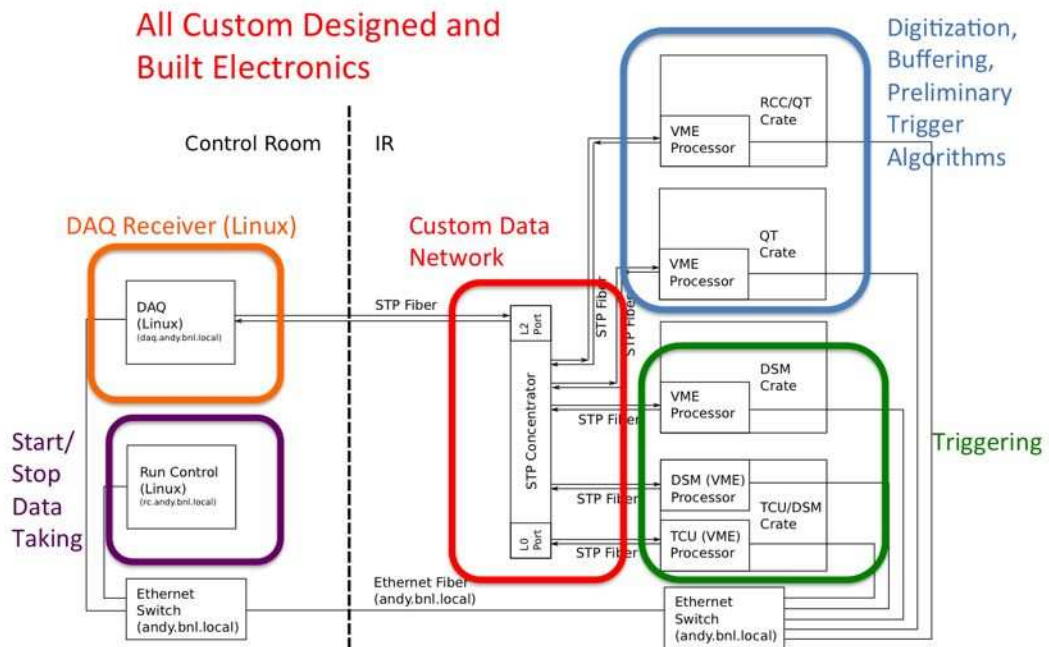


Figure 46: Schematic diagram of the trigger-daq system.

mid-rapidity array and preshower detectors are not involved in the trigger decision and are not connected to the decision tree. All detector systems use QT boards as their basic data input path; these board store 8B of data for each channel for 64k crossings in a cyclic memory. The QT boards also perform layer-0 logic manipulations as needed, such as finding the highest of the 32 channels of input ADC value (*e.g.* high tower), or summing input ADC values (*e.g.* cluster energy) and comparing to thresholds.

Up to 16 QT boards occupy a single 9U VME Crate as shown in Fig.48. QT boards for ZDC, BBC, FPD, ECal and HCal all have direct connections to send between 8 to 32 bits to a layer 1 DSM in the Level 1 crate. Each DSM board has connection to higher layer DSM boards in the L1 crate. For each RHIC crossing a trigger decision is made in the TCU based on combinations of up to 16 bits received from the last DSM board in the decision tree. There are no detector busy conditions at A_{NDY} ; the only busy condition arises if the TCU runs out of tokens indicating that the downstream trigger/DAQ resources are not available. We measured the token FIFO empty in two independent ways during run 11.

The TCU places its decision bits on the VME backplane where it is shipped to the TCD for distribution to each detector VME crate for distribution on the local backplane. Receipt of a trigger within each DSM board initiates copying the appropriate 8B of data for each channel from the indicated crossing address into a FIFO for readout. A TCU register is polled by the Level1 CPU and, when a trigger is detected, the L1 CPU initiates a data gather operation by having the STP broadcast the crossing number and an organizing token to each VME CPU. Each VME Crate has an STP PCMCIA card in the local crate-control CPU which is connected to an STP concentrator board. The Concentrator accepts data from the crates in a push architecture and ships it in packets up to 4kB in length to the level 2 Linux CPU. When L2 has received all of the token-tagged data from each crate it packages this and ships it over a Glink fiber to the DAQ computer for monitoring and storage. From local storage the data is asynchronously shipped to HPSS for archiving.

Cosmic ray triggers are used to test and tune the system using signals from ECal, HCal, and cosmic-ray scintillator paddles placed to form tracks within the various detector modules. LED triggers are taken at a rate of ~ 1 Hz to provide gain monitoring of all PS, ECal and HCal detectors. In run 13 A_{NDY} will operate with a small number of simultaneously active triggers to accept MB, HT, Esum, and cluster pair triggers.

1559 The DAQ system is operated from a web-based interface using unix-
1560 command controls. It is responsible for starting and stopping the TCU.
1561 The TCU operates continuously at the RHIC crossing frequency of 9.4 MHz.
1562 The DAQ system communicates with the RCC board to issue run and stop
1563 commands. Expanding the system for run 13 will involve rebuilding a num-
1564 ber of configuration files but will not entail architectural or significant code
1565 changes to the DAQ system.

1566 10.1 Acronyms

- 1567 • DSM: Data Storage and Manipulation: accepts 128 bits of input, passes
1568 the bits through an FPGA for logic decisions, and outputs up to 32 bits
- 1569 • FPGA: Field Programmable Gate Array: the logic engine on each
1570 board
- 1571 • TCU: Trigger Control Unit: accepts 16 “physics” input bits and up to
1572 16 detector busy bits (not used in A_NDY) to make a trigger decision
1573 for each RHIC crossing
- 1574 • TCD: Trigger-Clock Distribution: a 9U VME board connected to the
1575 output of the TCU to distribute token and crossing address to each of
1576 the detector VME Crates
- 1577 • TAC: Time-to-Amplitude-Converters: These are 16 channel boards
1578 converting time from a leading edge discriminator to the next RHIC
1579 Strobe into a current signal for digitization.
- 1580 • QT: 32 channel 9U VME charge and time digitizing boards
- 1581 • STP: STAR Trigger Pusher : this is a latency-free high-speed local
1582 network hardware and software protocol
- 1583 • PCMCIA: a hardware standard for plug-in boards into a VME CPU
- 1584 • GLink: Giga-link: a 2 Gb/s transmitter/receiver board for fiber-based
1585 communication
- 1586 • RCC: RHIC Clock and Control: a 9U VME board which provide the
1587 experiment clock and distributes this with the run/stop signal to all
1588 boards in the system

- 1589 • HPSS: High Priority Storage System: The mass storage used at the
1590 RHIC Computer Facility where the data is analyzed.
- 1591 • RS : RHIC Strobe: the 9.4 MHz RHIC clock, the rate at which beam
1592 bunches collide at the IP
- 1593 • IP : Interaction Point: the point at which the Blue and Yellow beams
1594 collide. This is really a diamond of length $\sigma \sim 30 \text{ cm}$.

1595 11 Simulations

1596 12 A_N DY Computing

1597 The computing needs for run-11 were met by employing a rack exccessed
1598 by the Atlas collaboration in their recent computing upgrade. The exccessed
1599 rack was moved to 1002D and the nodes were configured with Scientific Linux
1600 5.5 after increasing the memory for each computer to 4 gigabytes. Condor
1601 was installed and provides 116 slots for batch processing. One node had
1602 its internal disk NFS mounted across the cluster to provide 240 gigabytes
1603 of shared space. Inexpensive, commercial USB drives provided 8 terabytes
1604 of data storage. Standard analysis and simulation programs are in routine
1605 operation on this cluster. Database daemons, a local web server and online
1606 monitoring and analysis codes were routinely operated on this cluster during
1607 RHIC run 11. To date, offline analysis has been conducted on this computing
1608 cluster. Although extremely powerful for the run-11 test, the computing
1609 cluster has limitations in processing, disk storage and vulnerability to heat,
1610 especially during the summer months. Computing is a requirement for the
1611 A_N DY project, as discussed below.

1612 Analysis of the A_N DY data is the responsibility of individual members of
1613 the collaboration. Detector mapping and calibration files are shared among
1614 those doing the analysis. Physics results from A_N DY will require documenta-
1615 tion of the analysis and archiving of all coding used for the analysis. Multiple
1616 frameworks are available for analysis packages and we have employed essen-
1617 tially all options in prior work. Analysis of raw data or existing simulations
1618 are expected to represent only 15% of the total computing requirements for
1619 A_N DY. Instead, GEANT simulations are expected to be the most CPU in-
1620 tensive work. Analysis of the run-11 implementation of A_N DY was developed

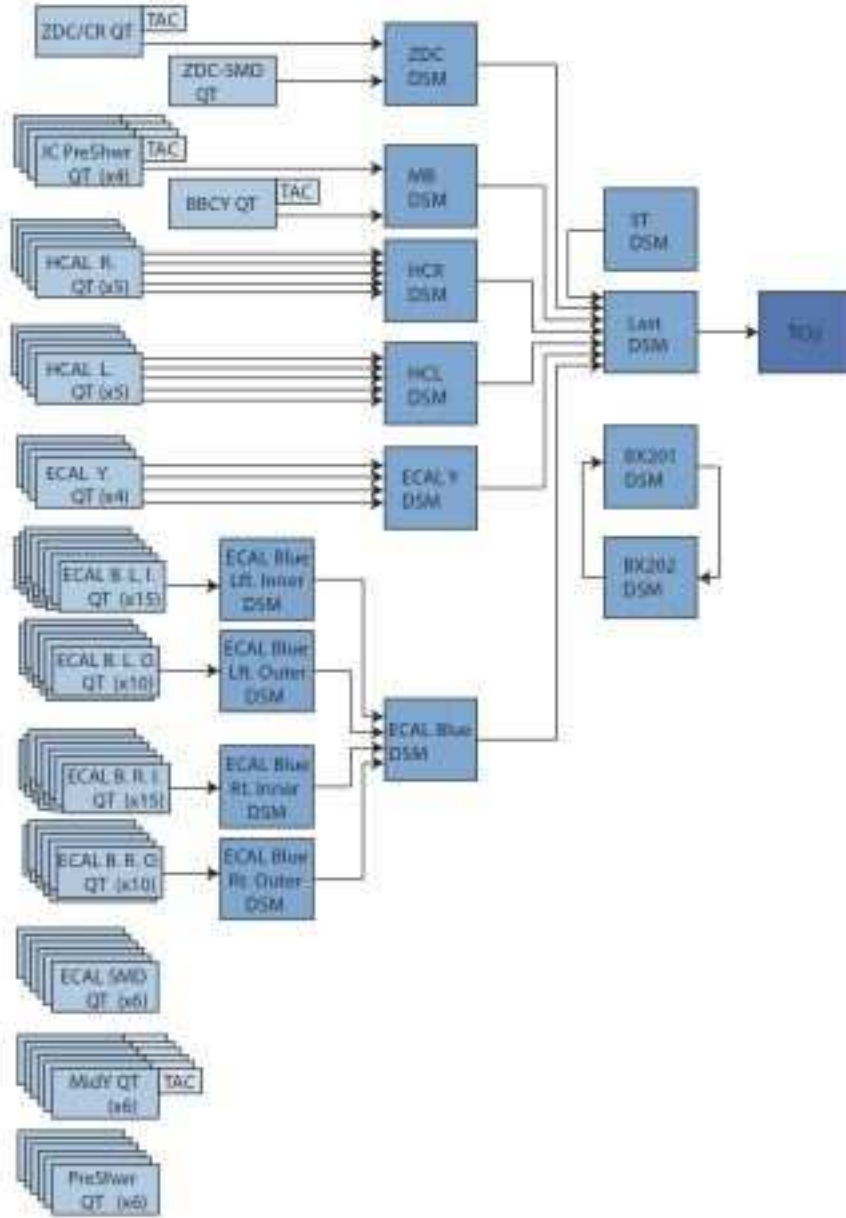


Figure 47: AnDY trigger decision tree for run13.

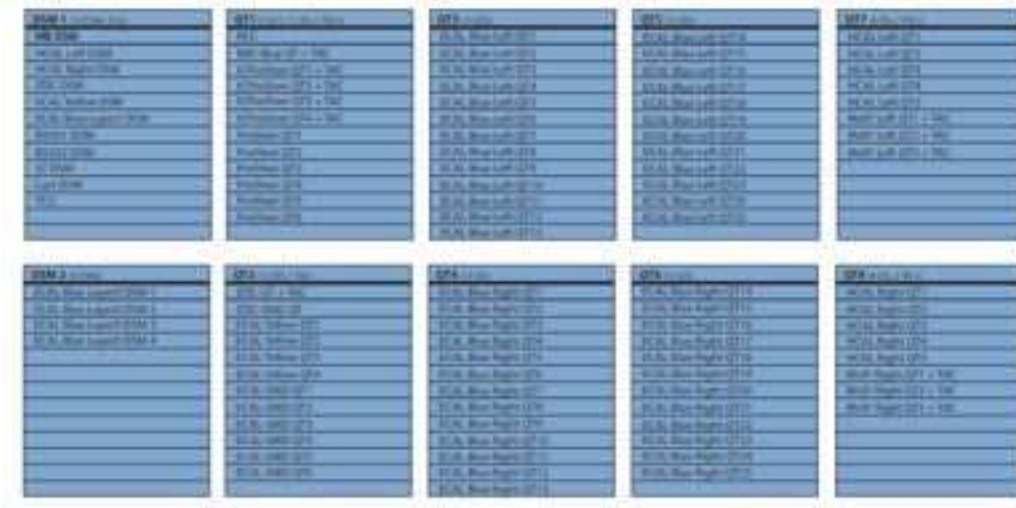


Figure 48: VME Crates for run13

on a 40M non-elastic, non-diffractive event sample, corresponding to $\sim 1nb^{-1}$ of data. This sample took 1 week of dedicated time on the 116-slot A_NDY cluster to generate through a GEANT model that accounts for all scintillating fibers within the run-11 HCal and requires 190 gigabytes of storage. At least two other event samples of comparable statistics await computing resources. They will employ filters on PYTHIA events (e.g., a jet filter and a large pair mass filter). The computing requirements for the run-13 A_NDY configuration are expected to be at least two orders of magnitude larger. A plan of how to meet the requirements awaits decisions at the RHIC/Atlas computing facility.

12.1 Near term computing needs

Immediate computing needs are cpu cycles for the analysis of the Run 11 data along with simulations incorporating the “as run” geometry required for the analysis, and space in HPSS for archiving the Run 11 data. Presumably the quickest way to obtain compute cycles with minimal impact on the resources of PHENIX and STAR is to reactivate some fraction of the 9 racks of PHENIX and STAR computers recently decommissioned. A new subnet can be obtained to isolate these machines from the PHENIX and STAR sys-

tems, and the standard operating system image can be reinstalled on the machines. There are clearly many factors which go into the decision as to what fraction of the decommissioned machines to reactivate, but, based on jobs run on recycled RCF machines in the A_N DY counting house, a minimum of two racks of machines with 4 GB of RAM is desired. If more machines can be made available, then the analysis can proceed faster.

Disk space is likely to be an issue in the short term. If sufficient centralized storage cannot be obtained to aid in the analysis and simulation of the Run 11 data, then the local disk on the reactivated compute nodes can be used both for local user file storage and for the Hadoop Distributed File System (HDFS). The latter will allow for efficient use of the distributed disk space on the farm nodes without requiring RACF manpower for the configuration or maintenance of the file system.

On the order of 4 Terabytes of HPSS storage is required for archiving the raw Run 11 data. The data is presently stored on consumer USB attached storage and a copy of the data should be made in HPSS as soon as possible to minimize the chance of losing the data. Analysis and simulation output should also get written into HPSS for archival purposes. The size of each of these outputs is expected to be comparable to the raw data set, so a total of 12 Terabytes of HPSS space is expected to be needed up to run 13. It is expected that new HPSS classes of service will be created for the raw and analyzed data.

12.2 Needs for run 13

Since data to be taken during run 13 will be written directly to HPSS, the optical fiber previously used by the BRAHMS experiments needs to be reactivated between the 1002D counting house and the RCF. Data will be transferred from the data acquisition system to HPSS at an estimated rate of 30 MB/s. An appropriately sized HPSS disk cache and sufficient tape drive resources need to be configured for the A_N DY classes of service to accommodate the data rate. The total raw data volume expected for run 13 is on the order of ~ 200 Terabytes.

12.3 Long term needs

It is assumed that the A_N DY experiment will take data through run 14 and then continue to analyze data for another two years after run 14. It is

further assumed that A_N DY will receive some fraction of the RCF allocation during these years and that current generation computers will be purchased for A_N DY so that the short term, reactivated compute nodes can be again retired. In addition, some amount of centralized storage needs to ramp up to facilitate the continued analysis of the data. NFS mounted central storage is probably sufficient for the needs of the experiment. The size of the centralized storage will depend on the allocation fraction for A_N DY.

As to the distributed disk on the compute nodes, it is not clear at this time if HDFS will continue to satisfy the needs of the experiment or if the need will arise to transition to an RCF managed distributed file system. The ultimate decision on the type of distributed file system, as with all of the computing infrastructure decisions for the experiment, will be made through active dialog with the computing facility staff.

13 Budget breakdown and resource-loaded timeline

Costs for staging A_N DY are driven primarily by electronics and cables for the various detectors and by installation of the magnet, although magnet costs need not be incurred prior to early 2014. There is at least 2 months of float in the schedule for preparing for run13 assuming that our funding is in place by 15 April 2012 and more than 4 months for run14. The resource-loaded timeline including cost and schedule are shown in Fig.49.

There are 7 detector systems in A_N DY in addition to the Zero Degree Calorimeters provided by CAD: the FPDY, BBCY, Mid η , TRK, PS, ECal, and HCAL detectors. The FPDY is a Yellow-beam-facing π^0 detector consisting of 98 PbGl cells as used in the run11 tests. It must be moved to the other side of the IP but incurs no costs. It is essential because it incorporates a Shower Maximum Detector that allows us to cleanly measure the π^0 anisotropy for the $\sqrt{s} = 500\text{GeV}$ beams. The BBCY is already in place from run11 and needs no further work. The change in expected luminosity forced an acceptance-increasing redesign that leads to elimination of the Blue-facing BBC that was used in run 11. This places a requirement on the central portion of the PreShower detector to provide timing signals used in minimum bias triggers. The Mid η detector is a scintillator hodoscope consisting of parts used in run11 augmented by TAC boards. The TRK is a set

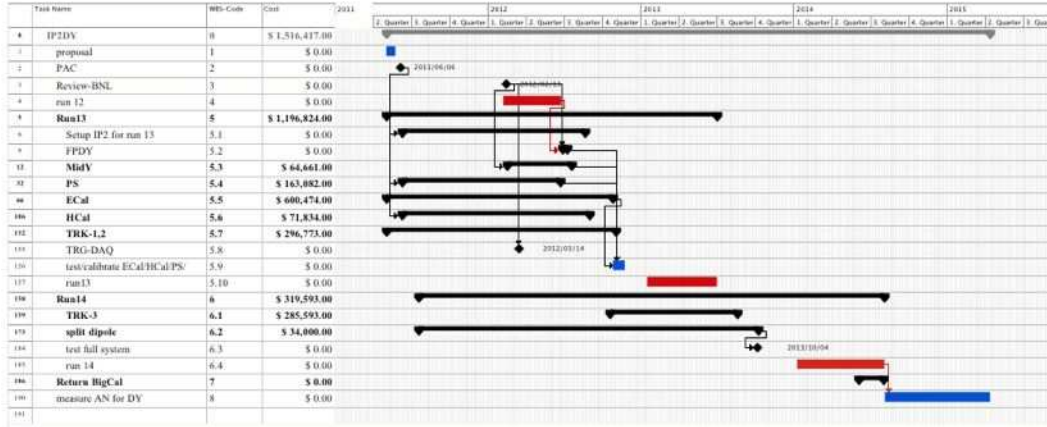


Figure 49: Resource loaded timeline

of 3 new GEM arrays called stations 1, 2 and 3 to provide charged particle tracking useful for photon elimination, electron identification, and detector calibration. The PS is a new scintillator slat detector whose fabrication includes all electronics for the readout. This is essential for both photon and hadron suppression. The ECal consists of 1596 lead-glass (PbGl) cells borrowed from JLab and needs a high voltage (HV) system and readout electronics. The HCal is taken from AGS-E864 and needs an HV control system to allow remote setting of HV values and additional channels over what we used in run11 to better match the full ECal detector acceptance.

The readout electronics for all of these detectors except TRK consists of QT boards (Fig.44) designed for STAR PMT readout. Fabrication of additional channels for A_N DY amounts to production of 70 additional boards and purchase of ~ 300 Positronix 8-channel connectors with long pigtailed. The QT fabrication is a standard production run, with organization taken care of by John Wolf and Jo Ann Bell at Berkeley. Total time from initial order to tested product delivery to BNL is 4 months, with most of that time taken in the actual production at the fabrication and loading firms. Cost and schedule are based on recent experience. We have over 100 of these boards in use at RHIC and have already informed our suppliers and fabricators of our intentions. Cost and delivery for the cables are based on quotes.

The Mid η detector consists of two modules of two planes of scintillator slats each. The scintillator, PMTs, and electronics were used in our run11 test: the scintillator slats are simply rearranged for this application. This

detector is used to determine z-vertex and mid-rapidity cluster multiplicity in coincidence with DY pairs. One module is placed on either side of the beam line to measure the z-vertex location for individual events. Each slat is $90\text{cm} \times 2.5\text{cm} \times 1\text{cm}$ viewed at both ends by an H5010 PMT with standard resistive bases fed from existing LeCroy High Voltage units. The module placement and plane separation are determined by diamond size. New mounting hardware and cabling must be provided for run 13. New TAC boards are needed to allow using time difference between PMTs on a single slat to give the plane 2-dimensional capability. This detector will be the responsibility of Xuan Li from Shandong University.

The TRK system consists of two stations of 6 GEM modules each and a third station consisting of 12 GEM modules. TRK1 and TRK2 provide tracks of the initial trajectories, while TRK3 sits just in front of PS to identify tracks entering ECal. Each module is $40 \times 50 \text{ cm}^2$ in area. The parts for the detector are purchased in large orders through BNL with assembly done at UVa. The electronics is purchased from CERN with the final readout stage provided by linux PCs.

The pre-shower detector consists of 228 scintillator slats of 3 different sizes as shown in Figure 34. The BC-408 equivalent scintillator is provided by Saint-Gobain cut and polished and ready for PMT mounting. The slats are similar to those used in run 11. Their cost is based on quotation from Saint-Gobain, with delivery specified as 6 weeks ARO. The slats will be viewed by XP2972 PMTs with Nunnemaker Cockroft-Walton bases that have been borrowed from Yale, originally used in AGS-E864. The PS detector will consist of 3 planes, two before a Pb converter and one following it. Each plane will be constructed in quadrants to facilitate mounting. The three planes will mount inside a preshower enclosure to simplify light-tight operation. This also minimizes material inhomogeneity thereby facilitating more accurate simulation of the $A_N\text{DY}$ detector. The HV consists of a distribution system originally designed for Phenix with a new computer-control for the distribution designed by UCB/SSL for $A_N\text{DY}$. Fabrication of the PS detector slats and mounting will be taken care of by Gunar Schnell and Charles Folz. The HV distribution boards will be fabricated through BNL by Akio Ogawa. The control system will be designed by Mike Ng and fabricated through UCB/SSL.

The ECal detector consists of PbI₂ cells viewed by FEU-84 PMTs. The system as borrowed from JLab consists of FEU-84 PMTs with two different kinds of bases: 596 cells (Yerevan) have standard resistive bases, and 1000 cells (Protvino) have resistive bases with booster power applied to the last

four dynodes. We intend to replace this hybrid HV system with CW bases for each PMT. This CW solution makes a simpler system which would then be used at JLab when the detector is returned. The CW design has been prototyped and tested. The bases will be fabricated in commercial firms with costs based on quotations and recent experience. The HV system will be taken care of by Jack Engelage and Michael Ng at UCB/SSL.

The ECal detector enclosure will be designed by Charles Folz and fabricated through BNL CAD. It will consist of two separate modules each mounted on Thompson bearings to allow them to move perpendicular to the beam line to facilitate access and enable operation close in the z direction to the HCal. It will include outer walls that are essentially covered with patch-panels for signals and for control. The QT readout electronics will mount very near this platform to allow Positronix connectors to go directly from the patch panels to the QT readout boards, thereby minimizing cable costs.

The HCal detector for A_NDY consists of two modules of 108 cells as used in run11 supplemented by 80 cells borrowed from Phenix to increase the acceptance and better match the ECal layout. The new cells will fill in the gap between the current 108-cell modules over and under the beam line, and will increase the vertical extent of the modules. Design and fabrication of the HCal housing and mechanical supports will be done by Charles Folz and CAD. Stacking the ~ 200 lb cells in their new arrangement will be accomplished by CAD personnel.

The data acquisition (DAQ) system will be a simple extension of the existing DAQ used for run11. This consists of first-generation electronics originally developed for STAR supplemented by QT boards. Addition of 5 VME crates does not alter the architecture and is a simple change in operating parameters. The DAQ system is the responsibility of Chris Perkins and UCB/SSL. It is expected to take less than 1 month of effort.

The costs shown include all overheads and burdens as well as contingency. The contingency for mechanical components being provided by CAD are 40%. The contingencies for items for which we have new quotes is 5%. Other items carry 20% - 30% contingency depending on their basis of estimate. Standard overheads and other burdens are shown in Table 1.

Resources enter many points in the Gannt chart and the cost tables. The basis for cost of each item is shown in Fig.50

Project milestones are shown in Table 2.

type	overhead multiplier	notes
BNL machine shop	1.349	for enclosures, mounts
BNL Purchase $\leq \$25k$	1.525	
BNL Purchases $\geq \$25k$	1.093	
BNL University contracts	1.175	to UVa. and UCB/SSL
UVa labor	1.52	GEM assembly
UCB/SSL	1.17	first \$25k : 0 after

Table 1: Institutional cost burdens

14 Management plan

The A_N DY management structure is based on the expectation that all collaborators will be involved in all aspects of the experiment, from design and implementation of hardware through operation of the detector and analysis of the data. Individual subsystems will typically have a lead individual with different aspects of the subsystem taken by different individuals. Analysis software is based on shared calibration and mapping routines with analysis code developed by individuals and groups for specific goals. Multiple approaches to analysis lead to more robust results and greatly aid in understanding subtleties as well as in uncovering errors. For organizing details, the lead persons for different aspects of the project are show first below:

- Spokesman: Les Bland
- Project Manager: Hank Crawford and Elke Aschenauer
- IP2 Infrastructure: C. Folz, BNL
- Computing and Networks: T. Throwe, BNL; A. Ogawa, BNL; J. Engelage, UCB/SSL
- Trigger/DAQ: Chris Perkins, UCB/SSL
- Electronics: J. Engelage, UCB/SSL; E. Judd, UCB/SSL
- ZDCs: Angelika Drees, BNL
- Preshower: Gunar Schnell, U. Basque; G. Simatovic, Zagreb; G. Igo, UCLA

	Subsystem	item	milestone date	description
1	ECal	detector at BNL	20110728	all BigCal cells are brought to BNL from JLab
2	ECal	CW Base prototype	20120109	test of Cockroft-Walton base for FUE84 PMTs
3	PS	HV control prototype	20120209	test remote HV control for Nun-nemaker base
4	DAQ	ready	20120314	daq ready for all tests and data
5	TRK	GEM module test	20120508	test 40x50 cm GEM at BNL
6	Midy	complete	20120702	detector for z-vertex ready
7	HCal	cells at IP2	20120723	move cells from Phenix to IP2
8	ECal-Y	complete	20120726	detector for π^0 w/ SMD ready
9	PS	complete	20120808	finish test of PS
10	HCal	complete	20120827	finish tests of HCal
11	ECal	complete	20120905	finish tests of ECal
12	MAG	split-dipole to IP2	20121002	magnet positioned at IP2
13	TRK	stations 1 and 2 complete	20121023	finish installation of stations 1 and 2
14	All	run13 system test	20121104	test of complete apparatus
15	MAG	complete	20120826	magnet installed and tested
14	TRK	station 3 complete	20130812	finish installation of station 3
15	All	run 14 system test	20131004	test of all systems
16	ECal	return to JLab	20120731	all cells, PMTs, bases to JLab

Table 2: AnDY project milestones

resource	base rate	overhead	local cost	contingency	Univ.contract	cost BOE/notes
cable-control-CW	1.00	1.53	1.53	1.05	1.00	1.61 14-conductor cable, \$50/100ft. IDC connectors, 50 cents in quantity.
cable-control-N	15.00	1.53	22.95	1.05	1.00	24.10
cable-RG58-bnc-bnc	18.00	1.53	27.54	1.05	1.00	28.92 KC quote
cable-positronix	75.00	1.53	114.75	1.05	1.00	120.49 KC quote
cable-positronix-k	147.00	1.53	224.91	1.05	1.00	236.16 KC quote
cable-patch-panel	950.00	1.53	1453.50	1.05	1.00	1526.18 digikey parts \$5.63 ea; 128/panel; 2 hr. BNL shop w/ burden @173/hr
cable-patch-panel-k-bnc	2115.00	1.53	3235.95	1.05	1.00	3397.75 digikey part \$25; 80/panel; (80*25*1.53+2*173)*1.05=3400
CW-control	500.00	1.17	585.00	1.30	1.17	889.79 recent build; each controls 128 bases
CW_bds	4.00	1.17	4.68	1.20	1.17	6.57 prototype build
CW-proto	5000.00	1.17	5850.00	1.20	1.17	8213.40 prototype cost - engineering
CW_parts	30.00	1.17	35.10	1.30	1.17	53.99 prototype cost
CW-Base	65.00	1.17	76.05	1.20	1.17	106.77 prototype cost
CW_id	25.00	1.17	29.25	1.20	1.17	41.07 prototype costs
GEMreadoutbds	2200.00	1.09	2398.00	1.05	1.00	2517.90 CERN quote; large purchase thru BNL
GEM-FEE	3.00	1.09	3.27	1.05	1.00	3.43 Uva recent fab; large purchase thru BNL
GEMfoils	1200.00	1.09	1308.00	1.05	1.00	1373.40 Uva recent fab; large purchase thru BNL
GEMframes	2600.00	1.09	2834.00	1.05	1.00	2975.70 Uva recent fab; large purchase thru BNL
GEMsupplies	1000.00	1.00	1000.00	1.05	1.17	1228.50 Uva recent fab
GEMsupport	2500.00	1.00	2500.00	1.05	1.17	3071.25 Uva recent fab
GEMassy	2000.00	1.53	3060.00	1.05	1.17	3759.21 Uva grad student, recent fab
GEMassy2	4000.00	1.53	6120.00	1.05	1.17	7518.42 Uva tech, recent fab
GEMgas	5000.00	1.00	5000.00	1.05	1.17	6142.50 Uva recent fab
GEMevblders	1000.00	1.53	1530.00	1.05	1.00	1606.50 recent purchase
flashers	50.00	1.17	58.50	1.05	1.17	71.87 recent fab
mech	1000.00	1.00	1000.00	1.00	1.00	1000.00 generic unit
Pb	1000.00	1.53	1530.00	1.05	1.00	1606.50 recent purchase
PMT-XP2972	0.00		0.00			0.00 borrowed from Yale
PMT-H5010	0.00		0.00			0.00 from AGS E896
Base-N	0.00		0.00			0.00 from Yale
HV-N-Control	100.00	1.17	117.00	1.20	1.17	164.27 estimate from similar fab
HV-N-proto	3000.00	1.17	3510.00	1.20	1.17	4928.04 estimate from similar fab
HV-N-Dist	347.00	1.17	405.99	1.05	1.17	498.76 recent fab
QT32-id	100.00	1.17	117.00	1.05	1.17	143.73 recent fab
QT32-bd	182.00	1.17	212.94	1.05	1.17	261.60 recent fab
QT32-pts	400.00	1.17	468.00	1.05	1.17	574.94 recent fab
QT8-bd	55.00	1.17	64.35	1.05	1.17	79.05 recent fab
QT8-pts	585.00	1.17	684.45	1.05	1.17	840.85 recent fab
QT8-id	100.00	1.17	117.00	1.05	1.17	143.73 recent fab
QTBOC	264.00	1.17	308.88	1.05	1.17	379.46 recent fab
QTAC-adapter	528.00	1.17	617.76	1.05	1.17	758.92 recent fab
scint	14524.00	1.53	22221.72	1.05	1.00	23332.81 St. Gobain quote
STP	667.00	1.17	780.39	1.05	1.17	958.71 recent fab
tac-bd	1957.00	1.17	2289.69	1.05	1.17	2812.88 recent fab
VME-QT	700.00	1.17	819.00	1.05	1.17	1006.14 recent backplane purchase

Figure 50: Resource rates including BOE.

- ECal: C. Perdrisat, College of William and Mary; E. Brash, TJNAI
- HCal: A. Ogawa, BNL
- Mid η : X. Li, Shandong
- FPD: L. Nogach, Protvino; N. Minaev, Protvino
- GEMs: N.Lyanage, K.Gnonvo, Uva.
- Magnet: W.Meng, C.Folz, BNL

Regarding memoranda of understanding (MOUs), we would seek to establish such documents as needed to aid our collaborators in their negotiations with their home institutions.

The lead institution for each detector is shown in Table 3.

Institutional responsibilities cover different aspects of each detector as shown in Table 4.

The FTE level of commitment for each member of A_NDY is shown in Table 5.

detector	institution	note
BBC	BNL	existing
Mid η	BNL/SU	existing scints and PMTs
FPDY	IHEP	existing
PS	Ikerbasque	all new
ECal	CWM/BNL	existing cells and PMTs
HCal	BNL	existing cells and PMTs
TRK	UVa	new GEM system

Table 3: Lead Institutions for Detectors

14.1 Key Performance Parameters (KPP)

We identify the following performance requirements for successful completion of the Drell-Yan A_N measurement:

1. ECal energy resolution: Calibration and resolution of each cell must be measured. Current cells provide $\sigma E/E \leq 15\%/\sqrt{E}$; mass resolution and signal/noise depend critically on knowing these.
2. ECal Cockcroft-Walton base operation: The bases must provide remotely settable High Voltage between 1300 – 1700 V with 12-bit accuracy. This is to allow gain matching of ECal FEU84 PMTs.
3. HCal and PS Nunnemaker base HV control: The HV-N controllers must provide remotely settable High Voltage between 1300 – 1700 V with 12-bit accuracy. This is to allow gain matching for the XP2972 PMTs on Hcal and PS which have Nunnemaker bases.
4. PS resolution: Each preshower slat must have charge resolution sufficient to resolve individual minimum ionizing particles (MIP) for 1, 2, 3 or more hits. This is essential to resolve single particles from early showers.
5. PS efficiency: Each slat must have efficiency $\geq 98\%$ for single MIPs. This is essential for photon rejection.

mechanical			
	HCal	cells	BNL
		enclosure	BNL
	ECal	cells	UWM
		enclosure	BNL
	PS	cells	Ikerbasque
		enclosure	BNL
	TRK	GEM modules	UVa
		enclosures	BNL
	Mid η	cells and mount	BNL
	FPDY	mount	BNL
Electronics			
	HCal	HV, QT, RO	UCB/SSL
	ECal	HV, QT, RO	UCB/SSL
	PS	HV, QT, RO	UCB/SSL
	TRK	HV	UVa
		FEE, RO	BNL
	BBC	HV	BNL
		QT, RO	UCB/SSL
	FPDY	HV	BNL
		QT, RO	UCB/SSL
	TRG/DAQ	DSM, TCU	UCB/SSL
Integration			
	installation	all	BNL
	operation	all	BNL, UCB/SSL
Software			
	data storage		BNL
	simulation		BNL, UCB/SSL
	Online		UCB/SSL
	Offline		BNL
	Analysis	TRK	UVa, BNL
		Calibration	IHEP, BNL, UVa
		J/Ψ , Υ	UCB/SSL, SUNYSB
		open charm	
		Λ_c	

Table 4: Institutional responsibilities

Support	Institution	Person	FY12	FY13	FY14	FY15
DOE	BNL	Aschenauer				
DOE	BNL	Bazilevsky				
DOE	BNL	Bland	0.95	0.95	0.95	0.95
DOE	BNL	Drees	0.15	0.15	0.15	0.15
DOE	BNL	Eyser	0.25	0.25	0.25	0.25
DOE	BNL	Folz	0.25	0.25	0.25	0.25
DOE	BNL	Makdisi	0.2	0.2	0.2	.2
DOE	BNL	Ogawa	0.9	0.9	0.9	0.7
DOE	BNL	Pile	0.15	0.15	0.15	0.15
DOE	BNL	Throwe	0.2	0.2	0.2	0.2
DOE	BNL/SHandong	Li	0.5			
	IHEP	Minaev	0.2	0.2	0.2	0.2
	IHEP	Nogach	0.5	0.5	0.5	0.5
EHU	Ikerbasque	Schnell	0.4	0.4	0.4	0.4
EHU	Ikerbasque	Van Hulse	0.25	0.25	0.25	0.25
	IU	Vossen				
DOE	LANL	Liu	0.1	0.1	0.1	0.1
NSF	NCU	Brash	0.1	0.1	0.1	0.1
NSF	NCU	Martin	0.5	0.5	0.5	0.5
NSF	NSU	Punjabi	0.1	0.1	0.1	0.1
	TJNF	Avakian				
DOE	UCB/SSL	Crawford	0.5	0.5	0.5	0.5
DOE	UCB/SSL	Engelage	0.25	0.25	0.25	0.25
DOE	UCB/SSL	Judd	0.1	0.1	0.1	0.1
DOE	UCB/SSL	Perkins	0.25	0.25	0.25	0.25
	UCLA	Igo				
DOE	UVa	Lyanage	0.1	0.1	0.1	0.1
DOE	UVa	Gnonvo	0.1	0.1	0.1	0.1
NSF	W&M	Perdrisat	0.1	0.1	0.1	0.1
	Yerevan	Abrahamyan				
	Yerevan	Shahinyan				
	Zagreb	Planinic	0.2	0.2	0.2	0.2
	Zagreb	Simatovic	.5	0.5	0.5	0.5

Table 5: Individual FTE commitments to AnDY

- 1858 6. GEM charge resolution: Single Mips must be distinguishable in individ-
1859 ual GEM modules. This is essential for distinguishing single particles
1860 from photon-conversion pairs.
- 1861 7. GEM position resolution: Each GEM module must have single particle
1862 position resolution of $\leq 500\mu$. This is essential for tracking individual
1863 electrons into the ECal cells.
- 1864 8. GEM efficiency: Each GEM module must provide $\geq 90\%$ efficiency
1865 for detecting a MIP. This is essential so that pairs of modules provide
1866 $\geq 99\%$ rejection power for photons.
- 1867 9. FPDY resolution FPDY must resolve π^0 up to $x_F \sim 0.4$. Need to cover
1868 the kinematics of DY and of our prior π^0 measurements.
- 1869 10. DAQ rates: The A_N DY data acquisition system must select, collect
1870 and store data at rates $\geq 2kHz$. Our signal rate is $\sim 0.005Hz$. We
1871 know from run11 that we must reject $\sim 10^5$ high-tower ECal triggers
1872 for each e^+e^- reconstructed: these occur at an average rate $\sim 500Hz$.
1873 We need higher daq rates to provide the minimum bias events needed
1874 for calibration and to allow flexibility in threshold selection for the
1875 high-tower or energy sum triggers.
- 1876 11. Dead time $\leq 10\%$ We need to sample full delivered luminosity. Thus
1877 our dead-time must be small.
- 1878 12. Magnet stability: The field must be stable to $\leq 2\%$ Bdl fluctuations over
1879 the course of the run. We measure charge sign by deflection at TRK-3,
1880 and a 2% change in Bdl will lead to an $\sim 0.3mm$ shift at TRK-3, while
1881 the deflections of typical DY electrons is $\sim mm$ at TRK-3.
- 1882 13. MidY Interaction z vertex: measured to $\leq 5cm$. The uncertainty in re-
1883 constructed mass scales directly with the uncertainty in vertex location
1884 - $dM \sim dz/z$.
- 1885 14. Diamond size: Small diamond is best. The mass resolution depends
1886 on knowing the vertex location, and our ability to measure deflection
1887 depends directly on diamond size.
- 1888 15. RHIC backgrounds: Beams must be tuned to have $\leq 10kHz$ of single
1889 beam backgrounds at IP2.

1890 16. RHIC Luminosity and Polarization: Average $\geq 10pb^{-1}/week$ with
 1891 $\geq 50\%$ polarization delivered to IP2. Our goal is to measure A_N to
 1892 better than 2% accuracy, implying a signal of $\geq 10^4$ reconstructed
 1893 pairs, approximately half at each transverse polarity, which requires
 1894 $100pb^{-1}$ delivered with polarization $\geq 50\%$.

1895 15 Safety

1896 A workplan has been established for the run-11 A_N DY effort (appendix 2).
 1897 It is a requirement that work to be done is understood and planned before it
 1898 is actually completed, to assure that it is completed safely. All components
 1899 of the proposed implementation at IP2 will be reviewed by the RHIC safety
 1900 committee. Action items from these reviews will be promptly resolved. It is
 1901 anticipated that the actual implementation will be subject to unannounced
 1902 inspections. Safety awareness is the primary consideration in prior work
 1903 done for forward calorimeter projects. This awareness is continually renewed
 1904 during weekly and daily planning meetings. The weekly meetings set broad
 1905 goals. The daily meetings are specific to the task at hand and assure com-
 1906 munication between those who will do the work.

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